



Steady periodic waves over a planar seabed: a global characterization

Matteo Antuono[†]

CNR-INM, Via di Vallerano 139, 00128 Rome, Italy

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The problem of propagation of steady periodic waves over a planar seabed is faced through the definition of a suitable semi-analytical iterative scheme. The latter is capable of describing highly nonlinear waves in deep, intermediate and shallow water conditions. Comparisons with the existing fifth-order theories show that the proposed model is accurate in all the regimes of motion and that it does not present any of the limitations affecting the Stokes and cnoidal wave solutions. Further, it also provides a reliable approximation of the dynamics of maximum-amplitude waves. The definition of the iterative scheme is preceded by a detailed study of the geometrical wave parameters. This latter part is aimed at the definition of a global scaling for water waves valid in all the regimes of motion.

Key words: surface gravity waves, coastal engineering

1. Introduction

The phenomenon of propagation of steady periodic waves over a planar seabed is one of the oldest and most fascinating problems in fluid dynamics, and dates back to the pioneering work of Stokes (1847). Beyond the interesting and challenging mathematical aspects, it embraces a wide variety of physical aspects (for example, the different behaviours over different wave depths, the interplay between dispersion and nonlinearity, the interactions with currents) that are important in many applied fields, like coastal and maritime engineering, naval engineering and environmental sciences. Despite the enormous amount of contributions on this subject, some aspects of wave propagation still deserve a dedicated study, especially in the field lying between practical applications and theory. The present work is dedicated to this latter aspect, providing a semi-analytical approach to the modelling of wave propagating over planar seabed that is accurate and,

†Email address for correspondence: matteo.antuono@cnr.it

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at the same time, provides solutions that are of practical interest for applications (e.g. numerical simulations and benchmarking). To better characterize the framework where the present work lies, we describe briefly the principal contributions in this field of research.

The analytical approach proposed by Stokes (1847, 1880) is based on the use of a perturbation expansion in a small parameter (the wave steepness) and applies to waves propagating over finite up to infinite depth. The third-order solution highlights some interesting features, like the increase of the wave celerity (namely, the wave translational velocity) as a function of the increase of the wave amplitude and the existence of a drift velocity in the same direction as the wave propagation. The main drawback of the Stokes approach is that the solution becomes singular as shallower depths are considered. Higher-order expansions (up to fifth-order) increase the range of validity of the Stokes solution slightly, but confirm the above features (see Fenton 1985).

The failure of the Stokes expansion is essentially caused by the singularity in shallow depths of the linear operator arising from the perturbation approach. In this regime of motion (which is characterized by the propagation of long waves in comparison with the still-water depth), the variations of the wave quantities are rather slow and, consequently, their order of magnitude is small. This results in a nonlinear interaction between the lowest mode of the steady periodic wave and its higher modes. The perturbation expansion at the basis of the cnoidal wave theory essentially stems from the above considerations and stands as a complementary approach in comparison to that used to derive the Stokes waves. At each order of the expansion, a nonlinear equation is obtained, whose solution is capable of describing the wave propagation up to the limit of infinite wavelength (i.e. solitary waves; Fenton 1972). On the other hand, the solution becomes inaccurate when the water depth becomes deeper, and higher-order solutions just mitigate this issue (see Fenton 1979).

In summary, the attainment of an exact or approximate analytical solution for the wave propagation in all the regimes of motion (i.e. shallow, intermediate and deep water) is still an open field of research, and despite the huge literature on this topic, no global analytical solution is currently available, to the author's knowledge.

The lack of such a global solution also results in a problem in practical applications, since the modelling of travelling waves in intermediate water often lies in the grey region where both the fifth-order theories (namely, Stokes and cnoidal waves) approach their own limits of validity.

The above considerations led many researchers to develop semi-analytical methods, that is, approaches where analytical modelling is provided up to a certain point, and a final part is left to numerical evaluation. In this case, explicit analytical solutions are generally not available but, in turn, it is possible to overcome many of the issues affecting the perturbation expansions and maintain a rigorous mathematical framework and an insight into the main physical aspects. To this class belong, for example, the works of Schwartz (1974), Williams (1981), Liao & Cheung (2003), Tao, Song & Chakrabarti (2007), Dyachenko, Lushnikov & Korotkevich (2016) and Zhong & Liao (2018). Some of those are devoted to the study of waves of maximum amplitude (also called waves of extreme height). These waves (whose existence in nature has been only conjectured) are predicted by the governing equations and are characterized by interesting specific features (both mathematical and physical), as the occurrence of a sharp angle at the crest (i.e. a singularity at the free surface). As shown in Henry (2006), Constantin (2006, 2012) and Constantin & Escher (2007), such a point behaves as an apparent stagnation point in the frame of reference moving with the wave speed, i.e. the velocity at the crest is null whereas the fluid particles are actually moving.

Despite a consistent number of theoretical works on extreme waves (Grant 1973; Norman 1974; Longuet-Higgins 1977; Toland 1978; Amick, Fraenkel & Toland 1982; Plotnikov 2002; Plotnikov & Toland 2004), their modelling is, at present, possible only through semi-analytical methods, whereas fifth-order wave theories cannot adequately represent such a high nonlinear phenomenon.

The present work belongs to the above-mentioned class of semi-analytical methods and is aimed at providing a fast and accurate model for wave propagation in all the regimes of motion, up to the limit of maximum-amplitude waves. The main target is to give a contribution to the theoretical description of steady periodic waves and to fill the gap between the theories in deep- and shallow-water conditions, supplying a reliable tool for practical applications (i.e. wave modelling, numerical benchmarking, etc.). The iterative scheme proposed in this work has been defined starting from the integral equation of Byatt-Smith (1970), which represents a rearrangement of the governing system of equations into a single equation in a suitable hodograph space. Special attention has been devoted to the physical characterization of the main parameters that arise in such a mathematical framework. The proposed iterative approach is intrinsically nonlinear and, consequently, is not affected by the singularity in the shallow-water limit that characterizes the linear operator of the Stokes wave.

The use of the integral equation of Byatt-Smith (1970) is motivated by its simple and compact representation of the wave problem that allows for a straightforward definition of the wave parameters and of the iterative scheme itself. Strangely, such an equation has received little attention in the literature, and to the author's knowledge, no theoretical findings can be enumerated for it at present. Incidentally, we highlight that other integral equations for steady waves have been proposed previously, as in in the pioneering works by Nekrasov (1921, 1928). In this latter case, a significant number of theoretical contributions exist (the interested reader can find an exhaustive summary of the literature on Nekrasov's equations in Kuznetsov 2021).

As a preparatory step to the definition of the iterative scheme, a study on the geometrical wave parameters is tackled in § 2 in order to define a global spatial scale associated with the vertical dynamics, that is, a significant length that represents the region where the fluid motion is essentially different from zero in all the regimes of motion. The use of such a scale allows for a global characterization of the solution without falling into a bias caused by the dependence on shallow- or deep-water lengths. It also facilitates the description of the phenomenon of wave propagation as a whole. Hence §§ 3 and 4 introduce the governing equation and the integral equation of Byatt-Smith (1970), and § 5 describes the iterative scheme. Finally, § 6 shows the outputs of the proposed scheme and the comparisons with the existing analytical solutions.

2. A global scaling for water waves

The phenomenon of wave propagation is characterized and therefore described by means of three spatial scales: a horizontal length scale x_0^* (related to the wavelength L^*), the wave amplitude a_0^* (half wave height), and the reference water depth h_0^* in the vertical direction. Using these quantities, it is possible to define the following dimensionless numbers:

$$\bar{\mu} = \frac{h_0^*}{x_0^*}, \quad \bar{\epsilon} = \frac{a_0^*}{h_0^*}, \quad \bar{\sigma} = \frac{a_0^*}{x_0^*} = \bar{\epsilon}\bar{\mu}.$$
(2.1*a*-*c*)

In particular, the first term, usually called dispersion parameter, is used to identify different regimes of motion: waves travel in shallow-water conditions if $\bar{\mu} \ll 1$ or in the deep-water

regime for $\bar{\mu} \gg 1$. Between these extrema, waves are said to propagate in intermediate depths. Finally, the second and third parameters in (2.1a-c) are called the nonlinearity and steepness parameter, respectively, and characterize the intensity of the wave dynamics.

From the last definition in (2.1a-c), it is clear that it is always possible to choose a pair of the above coefficients and deduce the remaining one. In particular, the pair $(\bar{\mu}, \bar{\epsilon})$ is used to described wave motion in shallow-water conditions, while the pair $(\bar{\mu}, \bar{\sigma})$ is generally adopted in deep water. The use of different pairs according to the different conditions of propagation is motivated by the fact that neither $\bar{\epsilon}$ nor $\bar{\sigma}$ can represent the wave dynamics globally over the three regimes of motion. In shallow depths, the condition $\bar{\mu} \ll 1$ (long waves) leads to $\bar{\sigma} = \bar{\mu}\bar{\epsilon} \ll \bar{\epsilon}$. This, along with the assumption of non-breaking waves (i.e. $\bar{\epsilon} \ll 1$), implies that the parameter $\bar{\sigma}$ is extremely small and, consequently, is not significant for the flow dynamics. The opposite occurs in deep water, where $\bar{\mu} \gg 1$ (short waves) leads to $\bar{\epsilon} = \bar{\sigma}/\bar{\mu} \ll \bar{\sigma}$. Again, the assumption that waves are non-breaking (i.e. $\bar{\sigma} \ll 1$ in this case) implies that $\bar{\epsilon}$ is not representative of the flow dynamics in deep water.

The definition of a global scaling across the regimes of motion is therefore the first step to derive a solution for waves of permanent shape travelling over a generic depth. The above considerations led us to introduce a further spatial scale, denoted z_0^* , that represents the scale associated with the vertical dynamics, that is, the length of the region where the fluid motion is essentially different from zero. Generally, this is a fraction of the wavelength in deep water, while it becomes comparable with the water depth in shallow-water conditions. Since z_0^* is a derived quantity (namely, it depends on the solution itself), it must depend on the remaining vertical scales, i.e. a_0^* and h_0^* . This implies the existence of the functional relation

$$\frac{z_0^*}{x_0^*} = f\left(\frac{a_0^*}{x_0^*}, \frac{h_0^*}{x_0^*}\right) = f\left(\bar{\sigma}, \bar{\mu}\right).$$
(2.2)

Since we restrict our analysis to non-breaking waves, we require $\bar{\sigma} \ll 1$ in deep water (where $1 \ll \bar{\mu}$) and $\bar{\epsilon} \ll 1$ in shallow water (where $\bar{\mu} \ll 1$). The former pair of inequalities implies $\bar{\sigma} \ll 1 \ll \bar{\mu}$, while the latter pair leads to $\bar{\sigma} = \bar{\epsilon}\bar{\mu} \ll \bar{\mu} \ll 1$. Hence for non-breaking waves, $\bar{\sigma} \ll \bar{\mu}$ holds true over all the regimes of motion. The above findings suggest that the dimensionless parameter $\bar{\mu}$ is the most meaningful to describe the relevant vertical scale for the wave propagation phenomenon. As a consequence, (2.2) can be simplified as

$$\frac{z_0^*}{x_0^*} \simeq f(0,\bar{\mu}) = \hat{f}(\bar{\mu}).$$
(2.3)

Starting from this result, we introduce new operative dimensionless parameters in analogy with those described in (2.1a-c):

$$\mu = \frac{z_0^*}{x_0^*} = \hat{f}(\bar{\mu}), \quad \epsilon = \frac{a_0^*}{z_0^*}.$$
(2.4*a*,*b*)

Note that there is no need to specify an equivalent definition involving $\bar{\sigma}$, since $\bar{\sigma} = \bar{\epsilon}\bar{\mu} = \epsilon\mu$. The most natural requirement for z_0^* is

$$z_0^* \to \begin{cases} h_0^* & \text{for } \bar{\mu} \to 0, \\ x_0^* & \text{for } \bar{\mu} \to +\infty, \end{cases} \quad \Rightarrow \quad \epsilon \to \begin{cases} \bar{\epsilon} & \text{for } \bar{\mu} \to 0, \\ \bar{\sigma} & \text{for } \bar{\mu} \to +\infty. \end{cases}$$
(2.5)

As a consequence of this requirement, z_0^* remains finite over all the regimes of motion, while h_0^* and x_0^* go to infinity in deep- and shallow-water conditions, respectively.

Further, ϵ always tends to the most 'significant' parameter according to the specific regime of motion. The above limits also imply the following asymptotic behaviours for μ :

$$\mu = \hat{f}(\bar{\mu}) \simeq \begin{cases} \bar{\mu} & \text{for } \bar{\mu} \to 0, \\ 1 & \text{for } \bar{\mu} \to +\infty. \end{cases}$$
(2.6)

At this point, we have to specify a choice for both z_0^* and x_0^* . The most natural choice for z_0^* comes from the first-order solution for the wave celerity in Stokes waves, $c_{s,0}^*$, which is generally regarded as a reliable approximation over all the three regimes of motion. In particular, this reads

$$\left(c_{s,0}^{*}\right)^{2} = g^{*} \, \frac{\tanh(\kappa_{0}^{*}h_{0}^{*})}{\kappa_{0}^{*}},\tag{2.7}$$

where $\kappa_0^* = 2\pi/L^*$ is the wavenumber, and g^* is the gravity acceleration. Defining the scale for the wave celerity as $\sqrt{g^* z_0^*}$, we require

$$\frac{(c_{s,0}^*)^2}{g^* z_0^*} = 1 \quad \Rightarrow \quad \mu = \frac{\tanh(\kappa_0^* h_0^*)}{\kappa_0^* x_0^*}.$$
(2.8)

Then, choosing $x_0^* = 1/\kappa_0^*$, we finally obtain

$$\mu = \tanh\left(\bar{\mu}\right),\tag{2.9}$$

which is consistent with the functional definition of μ given in (2.4*a*,*b*), and with the requirements in (2.6).

Using (2.9), it is possible to give simple theoretical definitions of the shallow- and deep-water regimes. Specifically, based on the asymptotic behaviours of μ , we introduce the quantities:

$$\bar{\mu}_S = \sup \{ \bar{\mu} \text{ such that } |\bar{\mu} - \tanh(\bar{\mu})| < 0.01 \},$$
 (2.10)

$$\bar{\mu}_D = \inf \{ \bar{\mu} \text{ such that } |1 - \tanh(\bar{\mu})| < 0.01 \}.$$
 (2.11)

Accordingly, we say that the shallow-water regime occurs for $\bar{\mu} < \bar{\mu}_S$, and the deep-water regime for $\bar{\mu} > \bar{\mu}_D$. For practical applications, these quantities can be approximated by $\bar{\mu}_S \simeq \pi/10$ and $\bar{\mu}_D \simeq 5\pi/6$. Remarkably, these values are close to those usually adopted in the literature (see, for example, Dean & Dalrymple 1991, pp. 64–65). Figure 1(*a*) displays μ as a function of $\bar{\mu}$, and the values of $\bar{\mu}_S$ and $\bar{\mu}_D$. Conversely, figure 1(*b*) shows the behaviours of the parameters $\bar{\sigma}$ and $\bar{\epsilon}$ as functions of μ for a chosen value of ϵ (namely, $\epsilon = 0.25$). In particular, using the definitions in (2.4*a*,*b*) and (2.9), we obtain

$$\bar{\sigma} = \epsilon \mu, \quad \bar{\epsilon} = \frac{\epsilon \mu}{\operatorname{arctanh}(\mu)}.$$
 (2.12*a*,*b*)

Figure 1 confirms the fact that neither $\bar{\sigma}$ nor $\bar{\epsilon}$ can be used as reference parameter for the description of the wave propagation over the three regimes of motion.

To give an example of the physical meaning and utility of the scale ϵ , we consider the propagation of waves of maximum amplitude. The existence of these waves is conjectured based on numerical and theoretical results even though, from a practical point of view, waves in the real world are expected to break before reaching their theoretical maximum amplitude. Through the years, different works have been devoted to the derivation of relations for the maximum amplitude of a wave propagating over a certain water depth.



Figure 1. (a) Parameter μ as a function of $\bar{\mu}$ (solid line). The dashed lines indicate the values of $\bar{\mu}_S$ and $\bar{\mu}_D$, while the dotted lines denote the asymptotic behaviours of μ . (b) Parameters $\bar{\sigma}$ and $\bar{\epsilon}$ as functions of μ for $\epsilon = 0.25$.

Among the most widespread, we report here the relation provided in Fenton (1990) based on the results by Williams (1981),

$$\bar{\epsilon}_{max} = \frac{1}{2} \left(\frac{0.141063\zeta + 0.0095721\zeta^2 + 0.0077829\zeta^3}{1 + 0.0788340\zeta + 0.0317567\zeta^2 + 0.0093407\zeta^3} \right),$$
(2.13)

and the expression proposed recently by Zhong & Liao (2018),

$$\bar{\epsilon}_{max} = \frac{1}{2} \left(\frac{0.14109\zeta + 0.00804\zeta^2 + 0.00949\zeta^3}{1 + 0.09671\zeta + 0.02695\zeta^2 + 0.01139\zeta^3} \right), \tag{2.14}$$

where $\zeta = 2\pi/\bar{\mu}$ varies in [1, 100]. In both the above formulas, the range of variation of $\bar{\epsilon}_{max}$ is quite large. In Williams (1981), $\bar{\epsilon}_{max}$ varies from 0.0707 for $\zeta = 1$ (deep water) to 0.4082 for $\zeta = 100$ (shallow water), and similar values are predicted by Zhong & Liao (2018), namely 0.0699 for $\zeta = 1$, and 0.4106 for $\zeta = 100$ (see figure 2*a*). Conversely, the range of variation becomes much more narrow if we use the new scaling and introduce $\epsilon_{max} = (\bar{\mu}/\mu)\bar{\epsilon}_{max}$. As shown in figure 2(*b*), ϵ_{max} ranges between 0.395 and 0.445. Hence the new scaling represents a significant order of magnitude for the maximum wave amplitude over all the regimes of motion.

Before proceeding to the next section, we underline that the parameter μ in (2.9) was already derived in Kirby (1998) by following a different approach. In that case, the aim of the work was to remedy an inconsistency in the scaling proposed in Beji (1995). A first application involving the use of the above parameter is reported in Janssen, Herbers & Battjes (2006) in the context of spectral models.



Figure 2. Laws for the maximum wave amplitude as proposed by Williams (1981) (dashed lines) and by Zhong & Liao (2018) (solid lines), displayed by using the parameters (a) $\bar{\epsilon}$, and (b) ϵ .

3. Governing equations

The dynamics of an inviscid and irrotational free-surface flow over a uniform seabed is described by the following system (in dimensional variables):

$$\begin{aligned} \phi_{z^{*}}^{*} &= \eta_{t^{*}}^{*} + \phi_{x^{*}}^{*} \eta_{x^{*}}^{*}, \\ \phi_{t^{*}}^{*} &+ \frac{\phi_{x^{*}}^{*2}}{2} + \frac{\phi_{z^{*}}^{*2}}{2} + g^{*} \eta^{*} = B^{*}, \\ \phi_{x^{*}x^{*}}^{*} &+ \phi_{z^{*}z^{*}}^{*} = 0, \\ \phi_{z^{*}}^{*} &= 0, \\ \phi_{z^{*}}^{*} &= 0, \\ z^{*} &= -h_{0}^{*}, \end{aligned}$$

$$(3.1)$$

where ϕ^* is the velocity potential, η^* and h_0^* are the free-surface elevation and the level of the horizontal seabed, g^* is the gravity acceleration, and B^* is the Bernoulli constant. The horizontal and vertical spatial coordinates, namely (x^*, z^*) , form a right-handed frame of reference where z^* points upwards and $z^* = 0$ indicates the still-water level.

Now let us consider waves that translate with a constant velocity c^* in the direction of decreasing values of x^* . This is equivalent to assuming that all quantities depend on t^* and x^* through the composite variable $x^* + c^*t^*$. Hence the system (3.1) becomes:

$$\begin{array}{l}
\left. \phi_{z^{*}}^{*} = \left(\phi_{x^{*}}^{*} + c^{*}\right)\eta_{x^{*}}^{*}, \\
\left. c^{*}\phi_{x^{*}}^{*} + \frac{\phi_{x^{*}}^{*2}}{2} + \frac{\phi_{z^{*}}^{*2}}{2} + g^{*}\eta^{*} = B^{*}, \\
\left. \phi_{x^{*}x^{*}}^{*} + \phi_{z^{*}z^{*}}^{*} = 0, \\
\left. \phi_{z^{*}}^{*} = 0, \\
\right. \\
\left. z^{*} = -h_{0}^{*}. \\
\end{array} \right\}$$

$$(3.2)$$

The dimensional scaling of the above equations is introduced by using the definitions in (2.4a,b). In particular, we use z_0^* for both the horizontal and vertical scale, since this length

is always finite over all the three regimes of motion. Then we set

$$x^* = z_0^* x, \quad z^* = z_0^* z, \quad t^* = t_0^* t = \sqrt{\frac{z_0^*}{g}} t,$$
 (3.3*a*-*c*)

and finally,

$$\eta^* = a_0^* \eta, \quad c^* = \sqrt{g z_0^*} F, \quad \phi^* = \phi_0^* \phi = c^* z_0^* \phi.$$
 (3.4*a*-*c*)

We also introduce the conjugate potential ψ^* and the pair of conjugate potentials Φ and Ψ that represent the fluid motion in the frame of reference moving with the wave:

$$\Phi = \phi + x, \quad \Psi = \psi + z + H. \tag{3.5a,b}$$

Here, $\psi_0^* = \phi_0^*$ has been used to scale Ψ and ψ , while $H = h_0^*/z_0^*$. As a consequence of the definition in (2.9), it follows that

$$H = \frac{\bar{\mu}}{\tanh(\bar{\mu})} = \frac{\arctan(\mu)}{\mu}.$$
(3.6)

Then, the system (3.2) becomes:

$$\begin{aligned}
\Phi_{z} &= \epsilon \frac{\mathrm{d}\eta}{\mathrm{d}x} \Phi_{x}, \\
\Phi_{x}^{2} &+ \Phi_{z}^{2} = \frac{2\epsilon^{2}B + F^{2} - 2\epsilon\eta}{F^{2}}, \\
\Phi_{xx} &+ \Phi_{zz} = 0, \\
\Phi_{z} &= 0, \\
\end{aligned}$$

$$(3.7)$$

$$(3.7)$$

where $B = B^*/(u_0^*)^2$, and $u_0^* = \epsilon \sqrt{g^* z_0^*}$ is the reference scale for the horizontal velocity. Before proceeding to the analysis, it is worth noting that Ψ is constant along the bottom and the free surface (see, for example, Byatt-Smith 1970). Without any loss of generality, we assume $\Psi = \Psi_s$ at the free surface and $\Psi = 0$ at the bottom.

3.1. The integral equation

In this subsection, we recall briefly the work by Byatt-Smith (1970) where the system (3.7) is recast into a single integral equation for the free-surface elevation η through a hodograph transformation.

Balancing the continuity equation, we obtain the scale for the vertical velocity, namely $w_0^* = u_0^*$. Through this, we write the usual Cauchy–Riemann relations between Φ and Ψ :

$$\begin{aligned}
\Phi_x &= \Psi_z = U, \\
\Phi_z &= -\Psi_x = W,
\end{aligned}$$
(3.8)

where U and W are the velocity components in the frame of reference moving with the wave. Now let us consider (Φ, Ψ) as independent variables. Then the following identities

hold true:

$$\Phi \equiv \Phi \left(x(\Phi, \Psi), z(\Phi, \Psi) \right), \quad \Psi \equiv \Psi \left(x(\Phi, \Psi), z(\Phi, \Psi) \right).$$
(3.9*a*,*b*)

Differentiating by Φ , we find

$$1 \equiv U \frac{\partial x}{\partial \Phi} + W \frac{\partial z}{\partial \Phi}, \quad 0 \equiv -W \frac{\partial x}{\partial \Phi} + U \frac{\partial z}{\partial \Phi}, \quad (3.10a,b)$$

and rearranging, we obtain the expressions

$$\frac{\partial x}{\partial \Phi} = \frac{U}{U^2 + W^2}, \quad \frac{\partial z}{\partial \Phi} = \frac{W}{U^2 + W^2}, \quad (3.11a,b)$$

from which we derive the following equation:

$$\left(\frac{\partial x}{\partial \Phi}\right)^2 + \left(\frac{\partial z}{\partial \Phi}\right)^2 = \frac{1}{U^2 + W^2}.$$
(3.12)

Conversely, differentiating the relations in (3.9a,b) by Ψ , we find

$$\frac{\partial x}{\partial \Psi} = -\frac{W}{U^2 + W^2}, \quad \frac{\partial z}{\partial \Psi} = \frac{U}{U^2 + W^2}, \quad (3.13a,b)$$

and by comparison with the equations in (3.11a,b), we derive the Cauchy–Riemann relations in the (Φ, Ψ) -space:

$$\frac{\partial x}{\partial \Phi} = \frac{\partial z}{\partial \Psi},
\frac{\partial x}{\partial \Psi} = -\frac{\partial z}{\partial \Phi}.$$
(3.14)

Using the results above, we rewrite system (3.7) as

$$z = \epsilon \eta,$$

$$\left(\frac{\mathrm{d} x_s}{\mathrm{d} \Phi}\right)^2 = \frac{F^2}{2\epsilon^2 B + F^2 - 2\epsilon \eta} - \left(\frac{\mathrm{d} z_s}{\mathrm{d} \Phi}\right)^2, \qquad \Psi = \Psi_s,$$

$$\frac{\partial^2 z}{\partial \Phi^2} + \frac{\partial^2 z}{\partial \Psi^2} = 0,$$

$$z = -H \qquad \qquad \Psi = 0,$$

$$(3.15)$$

where $d/d\Phi$ denotes the derivative along the free surface (i.e. at $\Psi = \Psi_s$). Accordingly, the subscript *s* indicates the variables evaluated at the free surface, namely $x_s = x(\Phi, \Psi_s)$ and $z_s = z(\Phi, \Psi_s)$. The solution of the Laplace equation for $z(\Psi, \Phi)$ that satisfies the impermeability condition at the seabed is

$$z(\Phi, \Psi) = -H + \int_{\mathbb{R}} \exp(-iy\Phi) A(y) \sinh(y\Psi) \,\mathrm{d}y, \qquad (3.16)$$

and through the Cauchy–Riemann relations in (3.14), we obtain that along Ψ_s ,

$$\frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi} = \left(\frac{\partial x}{\partial\Phi}\right)_{\Psi_s} = \left(\frac{\partial z}{\partial\Psi}\right)_{\Psi_s} = \int_{\mathbb{R}} \exp(-\mathrm{i}\,y\,\Phi) y A(y) \cosh\left(y\Psi_s\right) \mathrm{d}y. \tag{3.17}$$

The expression for $dx_s/d\Phi$ comes from the second equation of the system (3.15). In particular, we choose the positive sign, since we suppose x and Φ to be oriented in the same way. After substitution inside (3.17), we apply the inverse Fourier transform and obtain A(y). Substituting A(y) in (3.16) and evaluating the overall expression at $\Psi = \Psi$, we obtain

$$\epsilon \eta + H = \int_{\mathbb{R}} \exp(-iy\Phi) \mathscr{F}^{-1} \left[\frac{\mathrm{d} x_s}{\mathrm{d}\Phi} \right] \frac{\tanh\left(y\Psi_s\right)}{y} \,\mathrm{d}y,$$

$$\frac{\mathrm{d} x_s}{\mathrm{d}\Phi} = \left[\frac{F^2}{2\epsilon^2 B + F^2 - 2\epsilon\eta} - \epsilon^2 \left(\frac{\mathrm{d}\eta}{\mathrm{d}\Phi}\right)^2 \right]^{1/2},$$
(3.18)

where the identity $z_s(\Phi) = \epsilon \eta(\Phi)$ has been used. Finally, applying the convolution theorem and using (A2*a*,*b*) and (A8) of Appendix A, we find

$$\epsilon \eta + H = -\frac{1}{\pi} \int_{\mathbb{R}} \left[\frac{F^2}{2\epsilon^2 B + F^2 - 2\epsilon \eta} - \epsilon^2 \left(\frac{\mathrm{d}\eta}{\mathrm{d}\Phi} \right)^2 \right]^{1/2} \log \left(\tanh \left(\frac{\pi |\xi - \Phi|}{4\Psi_s} \right) \right) \mathrm{d}\xi.$$
(3.19)

Apart from the presence of the Bernoulli constant *B*, the above expression coincides with the integral equation for η defined by Byatt-Smith (1970). In the sequel, we will sometimes abbreviate (3.19) as

$$\epsilon \eta + H = \mathscr{C}\left[\frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi}\right], \quad \text{where } \frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi} = \left[\frac{F^2}{2\epsilon^2 B + F^2 - 2\epsilon\eta} - \epsilon^2 \left(\frac{\mathrm{d}\eta}{\mathrm{d}\Phi}\right)^2\right]^{1/2}, \quad (3.20)$$

and \mathscr{C} is the convolution operator described in Appendix A. In brief, the action of this operator and of its inverse on a generic periodic function $p = \sum_{n \in \mathbb{Z}} P_n \exp(in\mu\Phi)$ is

$$\mathscr{C}[p] = \Psi_s P_0 + \sum_{n \in \mathbb{Z}_0} \frac{\tanh(n\mu\Psi_s)}{n\mu} P_n \exp(in\mu\Phi), \qquad (3.21)$$

$$\mathscr{C}^{-1}[p] = \frac{P_0}{\Psi_s} + \sum_{n \in \mathbb{Z}_0} \frac{n\mu P_n}{\tanh(n\mu\Psi_s)} \exp(in\mu\Phi).$$
(3.22)

More details are given in Appendix A.

4. Periodic solutions

Equation (3.19) represents an elegant rearrangement of the initial system (3.2) into a single integral expression in the hodograph space (Φ , Ψ). In Byatt-Smith (1970), the parameter Ψ_s was used to select different kinds of waves (from Stokes to cnoidal and solitary waves), and its dependence on the remaining parameters (namely ϵ , μ , F and H in the present case) was obtained as an approximation from the solution itself. Differently from this approach, we derive here the explicit expression for Ψ_s for periodic waves. The value for the solitary wave is that associated with a periodic wave of infinite length.

Let us focus on periodic solutions in the dimensionless physical space. The dimensionless wavelength is $\ell_0 = L^*/z_0^*$, and since $L^* = 2\pi x_0^*$, we obtain $\ell_0 = 2\pi/\mu$. In the transformed space (Φ, Ψ) , this is equivalent to assuming that there exists a transformed wavelength $\hat{\ell}$ such that

$$\eta(\Phi + \widehat{\ell}) = \eta(\Phi)$$
 and $x_s(\Phi + \widehat{\ell}) = x_s(\Phi) + \ell_0.$ (4.1*a*,*b*)

The integral equation for η can be recast as

$$\epsilon \eta + H = -\frac{1}{\pi} \int_{\mathbb{R}} \frac{\mathrm{d} x_s(\xi)}{\mathrm{d}\xi} \log\left(\tanh\left(\frac{\pi|\xi - \Phi|}{4\Psi_s}\right) \right) \mathrm{d}\xi$$
$$= -\frac{1}{\pi} \int_{\mathbb{R}} \frac{\mathrm{d} x_s(\xi + \Phi)}{\mathrm{d}\xi} \log\left(\tanh\left(\frac{\pi|\xi|}{4\Psi_s}\right) \right) \mathrm{d}\xi, \tag{4.2}$$

so that, averaging over the period $\hat{\ell}$ (and denoting by η_0 the mean value of η in the hodograph plane), we find

$$\epsilon \eta_{0} + H = -\frac{1}{\pi} \int_{\mathbb{R}} d\xi \log \left(\tanh \left(\frac{\pi |\xi|}{4\Psi_{s}} \right) \right) \frac{1}{\hat{\ell}} \int_{0}^{\hat{\ell}} \frac{\mathrm{d}x_{s}(\xi + \Phi)}{\mathrm{d}\Phi} \,\mathrm{d}\Phi$$
$$= -\frac{1}{\pi} \int_{\mathbb{R}} \mathrm{d}\xi \log \left(\tanh \left(\frac{\pi |\xi|}{4\Psi_{s}} \right) \right) \left[\frac{x_{s}(\xi + \hat{\ell}) - x_{s}(\xi)}{\hat{\ell}} \right] \mathrm{d}\Phi$$
$$= \left[-\frac{1}{\pi} \int_{\mathbb{R}} \mathrm{d}\xi \log \left(\tanh \left(\frac{\pi |\xi|}{4\Psi_{s}} \right) \right) \mathrm{d}\Phi \right] \left(\frac{\ell_{0}}{\hat{\ell}} \right). \tag{4.3}$$

Using the properties of the convolution integral (see Appendix A), it follows that

$$\epsilon \eta_0 + H = \Psi_s \left(\frac{\ell_0}{\overline{\ell}}\right). \tag{4.4}$$

At this point, we need to derive an expression for ℓ . This comes from the definition of Φ in (3.5*a*,*b*). In particular,

 $\Phi(x+\ell_0,z) = \phi(x+\ell_0,z) + (x+\ell_0) = \phi(x,z) + (x+\ell_0) = \Phi(x,z) + \ell_0, \quad (4.5)$

and by comparison with (4.1*a*,*b*), it follows that $\hat{\ell} = \ell_0$. After substitution in (4.4), we finally obtain

$$\Psi_s = H + \epsilon \eta_0, \tag{4.6}$$

which represents the desired relation for Ψ_s .

In the next section and in Appendix B, we clarify the meaning of η_0 . For this aim, we require that the mean water level is zero in the physical space, that is,

$$\int_{0}^{\ell_{0}} \eta(x) \, \mathrm{d}x = 0 \quad \Rightarrow \quad \int_{0}^{\ell_{0}} \eta(\Phi) \left(\frac{\mathrm{d}x_{s}}{\mathrm{d}\Phi}\right) \mathrm{d}\Phi = 0. \tag{4.7}$$

This condition is also needed to have consistent definitions of the vertical scales. Indeed, an eventual set-up/set-down of the wave in the physical space would make the values of $\bar{\epsilon}$ and $\bar{\mu}$ imprecise.

Before proceeding, we address a further interesting behaviour of the integral equation (3.19). Following the proof given in the appendix of the work of Byatt-Smith (1970), it is simple to verify that any solution satisfying $\max(\epsilon \eta) = F^2/2 + \epsilon^2 B$ (i.e. any solution admitting a point where the denominator inside the integral (3.19) nullifies) has a corner of 120°, as conjectured by Stokes (1847). This case corresponds to the wave of maximum amplitude that is admissible for a given value of μ .

4.1. Mass flux and Stokes drift

The mass flux in the frame of reference of the Earth is given by

$$Q = \int_{-H}^{\epsilon \eta} u \, \mathrm{d}z = \frac{F}{\epsilon} \int_{-H}^{\epsilon \eta} \psi_z \, \mathrm{d}z, \qquad (4.8)$$

where $Q = Q^*/(z_0^*u_0^*)$. Using the definition (3.5*a*,*b*), we find

$$Q = \frac{F}{\epsilon} \left(\Psi_s - \epsilon \eta - H \right), \tag{4.9}$$

and averaging over the wavelength and imposing η to have a null mean (see (4.7)), we obtain

$$\bar{Q} = \frac{F}{\epsilon} \left(\Psi_s - H \right), \tag{4.10}$$

where \bar{Q} denotes the mean of Q in the physical space. Substituting (4.6) for Ψ_s , we finally find

$$\bar{Q} = F\eta_0, \tag{4.11}$$

which clarifies the role of η_0 . To avoid misunderstanding, we recall that η_0 is obtained by averaging η in the hodograph space, while \overline{Q} indicates an average in the physical plane. Following Fenton (1990), the drift velocity according to the second definition of Stokes is given by $U_D = \overline{Q}/H = F\eta_0/H$. In Appendix B, we show that $\eta_0 = O(\epsilon)$ and that $\eta_0 \le 0$. The latter result confirms that the Stokes drift occurs in the same direction as the wave propagation (i.e. in the direction of decreasing values of x^* for the present case).

Incidentally, we highlight that the definition of the drift velocity is not a trivial matter (see, for example, the discussion in Teles Da Silva & Peregrine 1988). This was first recognized by Stokes himself who introduced the drift velocity in two different ways, namely as a mean horizontal velocity (first definition) and as a depth-averaged velocity (second definition, as used in the present work). In Constantin (2013), the former value is proved to be always larger than the latter, and the difference between them is related to the excess kinetic energy as described in Henry (2021). The latter work also shows that the equipartition between the kinetic and potential energies predicted by the linear theory is no longer valid for nonlinear waves.

4.2. Singularity of the linear operator for vanishing depths

Before proceeding to the description of the iterative scheme, it is important to highlight the reasons that led us to develop this approach. To this purpose, we recall briefly the hypotheses that are the basis of the Stokes and cnoidal wave theories.

The approach used to obtain the Stokes waves is to assume $\epsilon \ll 1$ and expand (3.19) and the related parameters through perturbation series as follows:

$$\eta(\Phi) = \sum_{k=0}^{+\infty} \eta^{(k)}(\Phi) \,\epsilon^k, \quad F = \sum_{k=0}^{+\infty} F_k \epsilon^k, \quad B = \sum_{k=0}^{+\infty} B_k \epsilon^k.$$
(4.12*a*-*c*)

At the zeroth order, this yields the linear equation

$$\mathscr{L}[\eta^{(0)}] = \eta^{(0)} - \frac{\mathscr{C}\left[\eta^{(0)}\right]}{F_0^2} = 0, \tag{4.13}$$

where \mathscr{C} is the convolution operator in (3.20) with $\Psi_s = H$ (see Appendix A for more details). To obtain non-trivial solutions, the kernel of the linear operator \mathscr{L} has to be non-empty, and this is achieved by selecting $F_0^2 = \tanh(\mu H)/\mu$ (incidentally, we observe that $F_0 = 1$ as a consequence of the adopted scaling). The latter choice implies that \mathscr{L} admits a complete set of eigenfunctions $v_n = \exp(in\mu\Phi)$ with eigenvalues

$$\lambda_n = 1 - \frac{\tanh(n\mu H)}{n\tanh(\mu H)}.$$
(4.14)

Due to the presence of nonlinearities, the equation to solve at the *k*th order has the structure

$$\mathscr{L}[\eta^{(k)}] = \Xi^{(k)}(\Phi), \quad \text{with} \Xi^{(k)}(\Phi) = \sum_{n=2}^{N_k} \xi_n^{(k)} v_n(\Phi), \tag{4.15}$$

where N_k is the number of modes of the source term $\Xi^{(k)}$. To avoid the occurrence of resonant solutions, the eigenvectors $v_0 = const$ and $v_1(\Phi)$ are removed from $\Xi^{(k)}$ through a proper choice of the parameters F_k and B_k . Hence the general structure of the solution at the *k*th order is

$$\eta^{(k)} = e_0^{(k)} + e_1^{(k)} v_1(\Phi) + \sum_{n=2}^{N_k} \frac{\xi_n^{(k)}}{\lambda_n} v_n(\Phi), \qquad (4.16)$$

where $e_0^{(k)}$ and $e_1^{(k)}$ are constant factors. When $\mu \ll 1$ (i.e. waves propagate in the shallow water regime), the eigenvalues λ_n tend to zero, and consequently the expansion becomes singular, regardless of the magnitude of nonlinearities. This latter observation implies that any approach based on the use of the linear theory is inaccurate for the description of waves in shallow-water conditions. The limit of the Stokes wave theory is generally inspected through the Ursell number, U_r , that represents the ratio between the free-surface amplitudes at second order and first order. The request for this ratio to be small gives the following range of validity of the Stokes wave solution:

$$U_r = 8\pi^2 \frac{\bar{\epsilon}}{\bar{\mu}^2} \ll \frac{32}{3} \pi^2.$$
(4.17)

Hereinafter, we prefer to normalize the Ursell number as $\hat{U}_r = 3U_r/(32\pi^2)$, so that the above condition simplifies as $\hat{U}_r \ll 1$.

The cnoidal wave theory represents the counterpart in shallow-water conditions of the Stokes waves (see Fenton 1990). Roughly speaking, this relies on the assumption that the wave steepness ς is small so that the kernel in the integral of (3.18) can be formally expanded in a Taylor series:

$$\frac{\tanh\left(y\Psi_{s}\right)}{y} = \Psi_{s} - \frac{\Psi_{s}^{3}}{3}y^{2} + \frac{2\Psi_{s}^{5}}{15}y^{4} + O(y^{6}), \quad \text{with } y = O(\varsigma) \ll 1.$$
(4.18)

Through this expansion, the convolution integral in (3.19) can be represented as a series of differential operators (see, for example, Byatt-Smith 1970; Fenton 1972, 1979), and the assumption of small wave steepness (namely, $\varsigma^2 = O(\epsilon)$) leads to a rearrangement of the leading-order terms in the form of nonlinear equations. This circumvents the use of the linear operator, and consequently avoids the occurrence of singularities for $\bar{\mu}$ going to zero. Unfortunately, this approach is valid in very shallow depths or, equivalently, for very long

waves, as tsunami. Higher modes of the wave are, in fact, out of the hypothesis at the basis of the expansion in (4.18), and consequently the cnoidal wave theory cannot accurately describe waves propagating in intermediate and deep-water conditions. The limit of the cnoidal wave theory can be inspected by using the solution of Fenton (1979). This is represented by an expansion of even powers of the Jacobi cosine function $cn(\xi \mid m)$, where $m \in (0, 1)$ is the modulus of the Jacobi function, and $\xi = \alpha x^*/h_T^*$ with α a dimensionless parameter, and h_T^* the water depth at the wave trough. As described in Fenton (1979), α and h_T^* depend on both m and $\bar{\epsilon}$. The wavelength of the Jacobi cosine function is

$$L^* = 2 \frac{h_T^*(m,\bar{\epsilon})}{\alpha(m,\bar{\epsilon})} K(m) \quad \Rightarrow \quad \frac{1}{\bar{\mu}} = \frac{1}{\alpha(m,\bar{\epsilon})} \frac{h_T^*(m,\bar{\epsilon})}{h_0^*} \frac{K(m)}{\pi}, \tag{4.19}$$

where K is the complete elliptic integral of the first kind. Using the results in Fenton (1979), the above expression can be simplified through the leading-order relations

$$\frac{h_T^*}{h_0^*} = 1 + O(\bar{\epsilon}), \quad \alpha = \frac{\sqrt{3\bar{\epsilon}}}{2} \left(1 + O(\bar{\epsilon})\right), \tag{4.20a,b}$$

which lead to

$$\bar{\mu} = \frac{\pi}{2K(m)}\sqrt{3\bar{\epsilon}} + O(\bar{\epsilon}^{3/2}). \tag{4.21}$$

Since K(m) is a strictly increasing function, the shallow-water limit is attained for *m* tending to 1. In this case, K(m) diverges to infinity, and therefore $\bar{\mu}$ goes to zero (solitary wave limit). Conversely, the deep-water limit is reached for *m* approaching 0, where $K(0) = \pi/2$. In this case, we obtain the maximum value that can be assumed by $\bar{\mu}$. This corresponds to the bound

$$\bar{\mu} \le \sqrt{3\bar{\epsilon}} + O(\bar{\epsilon}^{3/2}) \quad \Rightarrow \quad \hat{U}_r \gtrsim 1/4,$$
(4.22)

which confirms the cnoidal wave theory to be complementary to the Stokes waves theory.

Summarizing, if we want to derive an approach that is reliable for all the fluid regimes, then it is of crucial importance to include the nonlinear terms straightforwardly and to avoid any sort of expansion based on small-amplitude or long-wave approximations. The iterative scheme described in the next section represents a valid strategy to overcome the above limitations.

5. The iterative scheme

In this section, we describe the iterative scheme used to obtain the solution for steady periodic waves. Introducing the variable $\theta = \eta - \eta_0$, we rewrite (3.20) as

$$\epsilon\theta + \Psi_s = \mathscr{C}\left[\frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi}\right], \quad \text{where } \frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi} = \left[\frac{F^2}{2\epsilon^2\beta + F^2 - 2\epsilon\theta} - \epsilon^2\dot{\theta}^2\right]^{1/2}, \tag{5.1}$$

where $\beta = B - \eta_0/\epsilon$, and $\dot{\theta} = d\theta/d\Phi$ is used for the sake of the notation. We recall that $\eta_0/\epsilon = O(1)$, as described in Appendix B. To achieve a more manageable expression, we first apply the inverse operator \mathscr{C}^{-1} and square, obtaining

$$[(\epsilon \mathscr{C}^{-1}[\theta] + 1)^2 + \epsilon^2 \dot{\theta}^2](2\epsilon^2\beta + F^2 - 2\epsilon\theta) = F^2.$$
(5.2)

Incidentally, we observe that the above expression is similar to (2.13) of Williams (1981), which, however, applies just to waves of maximum amplitude.

Expanding (5.2) and applying the operator \mathscr{C} , we find

$$\theta = \frac{\mathscr{C}[f]}{\chi}, \quad \text{where } f = \theta - \epsilon\beta + 2\epsilon\theta \,\mathscr{C}^{-1}[\theta] - \frac{\epsilon}{2} \,\chi S + \epsilon^2 \theta S, \tag{5.3}$$

where $\chi = 2\epsilon^2\beta + F^2$ and

$$S = (\mathscr{C}^{-1}[\theta])^2 + \dot{\theta}^2.$$
(5.4)

Equation (5.3) has the same linear operator of the integral equation (3.19) as a consequence of the subsequent application of the operators \mathscr{C}^{-1} and \mathscr{C} . In particular, using the results of Appendix A, it is possible to prove that the linear operator on the right-hand side of (5.3) admits a set of positive decreasing eigenvalues λ_n with $|\lambda_1| = 1$ (the last equality comes from the linear problem at the leading order). This property is a necessary condition for the convergence of the iterative scheme; however, it is not sufficient. In fact, the scheme converges if the term $(2\epsilon^2\beta + F^2 - 2\epsilon\theta)$ in (5.2) remains strictly greater than zero during the iterations. More details on this aspect are provided at the end of §§ 5.1 and 6.1.

A further advantage of the form in (5.3) is that only power terms of θ , θ and their counterparts transformed through \mathscr{C} appear. This allows for a straightforward derivation of the analytical expressions of the parameters χ , β and η_0 . In particular, these are obtained by imposing the following conditions.

(i) We normalize θ to unity, namely

$$[[\theta]] = \max(\theta) - \min(\theta) = 2.$$
(5.5)

This is a consequence of the definition of ϵ as a function of the wave amplitude a_0^* (namely, half the wave height). Hereinafter we assume that the maximum and minimum values occur at $\Phi = 0$ (wave crest) and $\Phi = \pi/\mu$ (wave trough). Consequently, the above condition is replaced by

$$[[\theta]] = \theta|_{\phi=0} - \theta|_{\phi=\pi/\mu} = 2.$$
(5.6)

Applying the operator [[\cdot]] to (5.3), we find the expression for χ :

$$\chi = \frac{2\left[\left[\mathscr{C}[\theta]\right]\right] + 4\epsilon\left[\left[\mathscr{C}\left[\theta\,\mathscr{C}^{-1}\left[\theta\right]\right]\right]\right] + 2\epsilon^{2}\left[\left[\mathscr{C}[\theta S]\right]\right]}{4 + \epsilon\left[\left[\mathscr{C}[S]\right]\right]}.$$
(5.7)

(ii) The mean of θ is zero in the hodograph space, namely $\langle \theta \rangle = 0$. This is a consequence of the definition of θ . As shown in Appendix B, this implies $\langle f \rangle = 0$ when the average is applied to (5.3). This latter condition gives the expression for β :

$$\beta = 2\langle \theta \, \mathscr{C}^{-1} \left[\theta \right] \rangle - \frac{\chi}{2} \, \langle S \rangle + \epsilon \, \langle \theta S \rangle.$$
(5.8)

(iii) The mean of η is zero in the physical space (see (4.7)). As described in Appendix B, this condition is equivalent to

$$\eta_0 = -\epsilon \langle \theta \, \mathscr{C}^{-1}[\theta] \rangle. \tag{5.9}$$

The corresponding expression for Ψ_s derives from (4.6).

This knowledge of χ , β and η_0 immediately allows for the evaluation of the physical parameters *F*, *B* and *U*_D (or equivalently, \bar{Q}).

We are now in position to introduce the iterative scheme. To this aim, we formally represent (5.3) as

$$\theta = \mathcal{N} \left[\theta \,|\, \Psi_s, \beta, \chi \right], \tag{5.10}$$

so that the iterative scheme is represented as

$$\theta_{m+1} = \mathscr{N}[\theta_m \,|\, \Psi_s^{(m)}, \, \beta_{m+1}, \, \chi_{m+1}].$$
(5.11)

Specifically, χ_{m+1} and β_{m+1} are obtained from θ_m and $\Psi_s^{(m)}$ through (5.7) and (5.8). After χ_{m+1} and β_{m+1} are evaluated, the solution for θ_{m+1} is computed by means of (5.11), and finally $\eta_0^{(m+1)}$ is obtained through (5.9). Incidentally, we highlight that the operators \mathscr{C} and \mathscr{C}^{-1} used on the right-hand side of (5.11) depend on $\Psi_s^{(m)} = H + \epsilon \eta_0^{(m)}$, and this means that they change slightly after each iteration (see (3.21) and (3.22)). As an initial guess for the iterative scheme, we choose

$$\theta_0 = \cos(\mu \Phi), \quad \eta_0^{(0)} = 0, \quad \Psi_s^{(0)} = H.$$
 (5.12*a*-*c*)

Consistently with (5.6), the above function satisfies the conditions $\max \theta = \theta(\Phi = 0)$, $\min \theta = \theta(\Phi = \pi/\mu)$ and $[[\theta]] = 2$.

5.1. Details on the numerical implementation

Starting from the single-mode initial guess $\theta_0 = \cos(\mu \Phi)$, additional modes are generated through the nonlinear terms in the right-hand side of (5.11). At the *m*-iteration, we obtain 3^m modes, and the *n*th mode of the iterative solution η has the form

$$E_n^{(m)} = \begin{cases} \epsilon^{n-1} G_n^{(m)}(\mu, \epsilon) & \text{for } n \le 3^m, \\ 0 & \text{elsewhere,} \end{cases}$$
(5.13)

where $G_n^{(m)} = O(1)$. Obviously, only a finite number of modes N can be considered for the numerical solution. The specific choice for N is motivated not only by the order of magnitude of the Nth mode but also by the wavelength that we want to describe. For example, in shallow-water conditions, a larger number of modes is needed, since the wave crest is very 'localized' in space in comparison to the wavelength. This is also pointed out in Zhong & Liao (2018), where 50 000 modes are used to model a wave of maximum amplitude with $\bar{\mu} = 1.0053 \times 10^{-2}$, while 5000 modes are necessary for a wave of maximum amplitude in infinite depth with equal accuracy. In the present case, the total number of modes N is set so that the energy stored at the convergence in the highest one-tenth of modes is less than a certain threshold value, namely

$$\left(\sum_{n=R}^{N} E_n^2\right) \left(\sum_{n=1}^{N} E_n^2\right)^{-1} < 10^{-8},$$
(5.14)

where $R = \lfloor 9N/10 \rceil$, and $\lfloor \cdot \rceil$ denotes the nearest integer (the superscript *m* is dropped here for ease of notation).

It is easy to convince oneself that the analytical solution obtained from the initial guess in (5.12*a*-*c*), as well as any single term in (5.3), can be represented through a cosine Fourier series. Assuming that the coefficients of θ_m have been computed already, the coefficients of $\dot{\theta}_m$, $\mathscr{C}[\theta_m]$ and $\mathscr{C}^{-1}[\theta_m]$ are obtained straightforwardly by using (3.21) and

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(3.22). Then the computation of four convolution series is required to find the coefficients of $\dot{\theta}_m^2$, $(\mathscr{C}^{-1}[\theta_m])^2$, $\theta_m S_m$ and $\theta_m \mathscr{C}^{-1}[\theta_m]$. This represents the most computationally expensive part of the scheme. Conversely, the computational cost related to the evaluation of the operators $\langle \cdot \rangle$ and $[[\cdot]]$ is relatively small. More details are given in Appendix C.

Incidentally, we underline that the condition $\chi_m - 2\epsilon \theta_m > 0$ has to be satisfied during the iterations to guarantee that the denominator inside the integral of (5.1) remains positive, to avoid the occurrence of non-physical solutions. In practice it is sufficient to check that the parameter $\nu_m = \chi_m - 2\epsilon \max \theta_m$ is strictly positive.

The convergence rate of the iterative scheme is computed as

$$q_m = \frac{\log (C_m / C_{m-1})}{\log (C_{m-1} / C_{m-2})}, \quad \text{where } C_m = \max_i \left[c_m^{(i)} \right], \tag{5.15}$$

and

$$c_m^{(1)} = \frac{\|\theta_m - \theta_{m-1}\|}{\|\theta_m\|}, \quad c_m^{(2)} = \frac{|\chi_m - \chi_{m-1}|}{|\chi_m|}, \tag{5.16a,b}$$

$$c_m^{(3)} = \frac{|\beta_m - \beta_{m-1}|}{|\beta_m|}, \quad c_m^{(4)} = \frac{|\eta_0^{(m)} - \eta_0^{(m-1)}|}{|\eta_0^{(m)}|}.$$
 (5.17*a*,*b*)

The convergence of the iterative scheme is assessed at the iteration M such that $C_M < 10^{-9}$. We anticipate here that the convergence rate of the present iterative scheme is approximately linear, and that it is slightly higher in deep-water conditions.

6. Numerical solutions and applications

The iterative scheme is implemented in Fortran 90 by using double precision, and the numerical solutions are run on a 48-core Intel Xeon(R), Silver 4214 CPU, 2.20 GHz. The OpenMP communication protocol is used to accelerate the computations. After the modes of η are computed, the solution for $x_s(\Phi)$ is obtained through the formulas described in Appendix D.

6.1. Maximum-amplitude solutions

We start the description of the numerical results from the theoretical limit of the proposed scheme, that is, when the parameter $v_m = \chi_m - 2\epsilon \max \theta_m$ tends to zero. This corresponds to the maximum-amplitude solutions that are admissible for a chosen wavelength. Numerically, this limit is approximated by selecting a value for $\bar{\mu}$ and increasing ϵ up to a maximal value, say ϵ_{max} , before the parameter v_m becomes negative for the first time during the iterations. The parameter ϵ_{max} is computed with four-digit accuracy, and the corresponding value of ν at convergence is denoted by v_M . Table 1 reports ϵ_{max} , the number of modes N, the total number of iterations M and v_M for $\bar{\mu}$ ranging from deep to shallow water. The values of v_M are $O(10^{-2})$ in all the regimes of motion. Similarly to v_M , the total number of iterations is influenced weakly by the changes of $\bar{\mu}$, and varies between 615 and 727. On the contrary, the number of modes tends to infinity rapidly as $\bar{\mu}$ goes to zero, while it is constant in deep water.

Differently from Williams (1981) and Zhong & Liao (2018), where the problem of steady periodic waves was formulated directly for waves of maximum amplitude (namely, for $\nu = 0$), the present approach can be regarded as a limit 'from below' (even though the limiting value is never reached in practice). For this reason, ϵ_{max} obtained through

Ν $\bar{\mu}$ ϵ_{max} M v_M 719 $\pi/512$ 0.4084 60 0 00 0.0331 $\pi/256$ 0.4076 30 0 0 0 693 0.0332 711 $\pi/128$ 0.4062 15 000 0.0327 $\pi/64$ 0.4034 10 0 00 727 0.0328 0.3987 719 $\pi/32$ 5500 0.0318 $\pi/16$ 0.3919 2500 703 0.0304 0.0296 $\pi/12$ 0.3892 2000 695 $\pi/8$ 1500 660 0.0287 0.3875 650 0.0279 $\pi/6$ 0.3897 1200 $\pi/5$ 0.3932 1000 645 0.0273 $\pi/4$ 0.3997 750 623 0.0266 $\pi/3$ 750 629 0.0260 0.4107 $\pi/2$ 700 627 0.0252 0.4262 $2\pi/3$ 0.4333 500 615 0.0249 500 622 0.0246 0.4371 π 500 623 0.0245 $4\pi/3$ 0.4376 $3\pi/2$ 0.4377 500 645 0.0243 2π 0.4377 500 645 0.0244 0.4377 500 645 0.0244 4π

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Table	1.	Computed	values of ϵ_{max} ,	number	of modes	N, total	number	of iteration	ıs M,	and 1	v at the
convergence for $\bar{\mu}$ ranging from deep to shallow water.											

the present scheme is expected to underestimate the numerical results from Williams (1981) and Zhong & Liao (2018). Fortunately, the discrepancies are rather small, as illustrated in table 2, where the results of Williams (1981) and Zhong & Liao (2018) are compared with those predicted by the proposed model, and in figure 3(a). In this latter case, the outputs of the proposed scheme are indicated with black dots, while the black solid line indicates a third-order polynomial fitting curves. In the same plot, the red dashed line and the green solid line indicate the solutions for ϵ_{max} predicted in Williams (1981) and Zhong & Liao (2018), respectively (see (2.13) and (2.14)). The relative error in the L^{∞} -norm between the present solution and that of Williams (1981) is below 1.7%, while for that of Zhong & Liao (2018) it is below 4.1%. We highlight that the use of the parameter $\mu = \tanh(\bar{\mu})$ on the ordinate axis gives a more regular and somehow meaningful trend of the results in comparison to that shown in figure 2. In particular, ϵ_{max} attains a minimum (i.e. $\epsilon_{max} = 0.3875$) at about $\mu = 0.3737$ (namely, $\bar{\mu} = \pi/8$), which corresponds to intermediate-water conditions (close to the boundary with the shallow-water regime; see figure 1). Figures 3(b-d) display the behaviour of the parameters F, B and Q corresponding to the values of ϵ_{max} in figure 3(a). Again, the black dots indicate the numerical outputs and the solid lines are proper fitting curves. Both Fand Q are strictly decreasing with μ , while B attains a maximum (that is, B = 0.063) at about $\mu = 0.5569$ (namely, $\bar{\mu} = \pi/5$). The fitting curves are

$$\epsilon_{max} = -0.0427\mu^3 + 0.1960\mu^2 - 0.1248\mu + 0.4091, \tag{6.1}$$

$$F = \frac{1.2922 - 0.1903\mu - 0.2656\mu^2}{1 + 0.2188\mu - 0.4536\mu^2},$$
(6.2)



Figure 3. Numerical solutions (black dots) and fitting curves (black solid lines) for (a) ϵ_{max} and for the corresponding values of (b) F, (c) B and (d) \overline{Q} , as functions of μ . The red dashed line and the green solid line in (a) indicate the solutions for ϵ_{max} predicted in Williams (1981) and Zhong & Liao (2018), respectively.

$$B = -0.0594\mu^4 + 0.0506\mu^3 - 0.2235\mu^2 + 0.2323\mu,$$
(6.3)

$$\bar{Q} = -\frac{0.3261\mu - 0.2434\mu^2}{1 + 0.3902\mu - 0.8770\mu^2}.$$
(6.4)

It is interesting to analyse the values predicted for infinite depth ($\mu \rightarrow 1$) and infinite wavelength ($\mu \rightarrow 0$). In the former case, (6.1) gives $\epsilon_{max} = 0.4376$, while $\epsilon_{max} = 0.4432$ is obtained by using (2.13) and (2.14) of Williams (1981) and Zhong & Liao (2018), respectively. Conversely, for infinite wavelength (namely, for solitary waves), (6.1) gives $\epsilon_{max} = 0.4091$, while $\epsilon_{max} = 0.4166$ is computed for Williams (1981) and Zhong & Liao (2018). In this latter case, the most accurate results are likely those described in Fenton (1972), where a ninth-order expansion is derived for solitary waves and $\epsilon_{max} = 0.425$ is predicted. Remarkably, the differences on the Froude number are even smaller, since F = 1.30 is computed in Fenton (1972) and F = 1.2922 is obtained through (6.2).

Table 3 displays the convergence of the Froude number at $\bar{\mu} = 1.0053 \times 10^{-2}$, as predicted by the present scheme and by the model of Zhong & Liao (2018). Incidentally, we observe that in the latter work, the parameter ϵ_{max} is an output of the scheme, and consequently it varies as the number of modes increases. The values of ϵ_{max} for the model of Zhong & Liao (2018) are reported in the table caption. Conversely, in the present model, $\epsilon_{max} = 0.4079$ for all the cases. Both the schemes prove to be in good agreement even when a relatively low number of modes is used.

Figure 4(*a*) displays the free-surface solutions of maximum amplitude for $\bar{\mu} = 4\pi$ (blue solid line), $\bar{\mu} = \pi/4$ (red solid line) and $\bar{\mu} = \pi/64$ (green solid line) scaled over one period (namely, $\ell_0 = 2\pi/\mu$), along with the theoretical slope $\pi/3$ at the crests (dashed lines). In all the regimes of motion, the agreement with the theoretical slope is good, proving that the angle at the crest is 120° as conjectured by Stokes (1847). Incidentally,

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μ	Williams (1981)	Zhong & Liao (2018)	Present model
1.0000	0.4432	0.4432	0.4376
0.9801	0.4396	0.4418	0.4349
0.9229	0.4306	0.4374	0.4273
0.8356	0.4208	0.4303	0.4168
0.7246	0.4110	0.4222	0.4053
0.5987	0.4025	0.4150	0.3955
0.4706	0.3969	0.4072	0.3893
0.3425	0.3951	0.4035	0.3876
0.2194	0.3975	0.4047	0.3907
0.1052	0.4047	0.4084	0.3981
0.0512		0.4127	0.4032
0.0305	_	0.4138	0.4055
0.0100		0.4150	0.4079

Table 2. Values of ϵ_{max} as reported in Williams (1981) and in Zhong & Liao (2018), and as predicted by the proposed model.

	N (modes)	21 000	28 000	35 000	40 000	50 000	55 000
	Zhong & Liao (2018)	1.2832	1.2843	1.2849	1.2853	1.2857	1.2857
	Present model	1.2872	1.2874	1.2875	1.2875	1.2876	1.2876
Table 3. Co (2018), the 0.4134, 0.414 the cases.	nvergence of F for parameter ϵ_{max} vari 11, 0.4145, 0.4148, 0.4150	$\bar{\mu} = 1.0$ ties as the p. 0.4150.	053 × 10 ⁻ he numb Conversel	⁻² . We ber of m y, in the	highlight nodes inc present	that in greases. So model, ϵ_i	Zhong & Liao Specifically, $\epsilon_{max} =$ max = 0.4079 in all

we highlight that the solutions obtained with the present scheme just approximate this theoretical angle since for $\nu > 0$, the derivative at the crest is always zero. This is shown in figure 4(*b*), where a detail of the deep-water case is displayed. Finally, figure 5 shows the solution in deep water (i.e. $\bar{\mu} = 4\pi$) along with the numerical results obtained in Dyachenko *et al.* (2016) (green dots) and Schwartz (1974) (red diamonds) for waves in infinite depth. In comparison to these outputs, the present solution displays a small shift downwards along the vertical direction. In any case, the match with the outputs of Dyachenko *et al.* (2016) and Schwartz (1974) appears good if the free surface is translated upwards (dashed line), proving that the solution with $\mu = \pi/4$ is a good approximation of the solution for waves over infinite depth. Remarkably, the integral over the period of the translated solution gives 0.011 against -4.84×10^{-4} of the original signal. This proves that (4.7) is satisfied accurately by the proposed iterative scheme, while the results by Dyachenko *et al.* (2016) and Schwartz (1974) are actually characterized by a slight set-up.

Before proceeding to the next section, we address some details about the numerical simulations. The most computationally expensive simulations are those characterized by the maximum wave amplitude in shallow-water conditions, since they require a large number of modes to satisfy the energy condition in (5.14). In particular, the simulation with $\bar{\mu} = \pi/512 = 6.1359 \times 10^{-3}$ (i.e. the shallowest case considered in the present work), $\epsilon_{max} = 0.4084$ and 60 000 modes reaches the convergence in 719 iterations and takes about 800 seconds with 42 cores. Incidentally, we observe that in Zhong & Liao (2018), the shallowest case was characterized by $\bar{\mu} = 1.0053 \times 10^{-2}$ and 50 000 modes, and a



Figure 4. (a) The maximum-amplitude free-surface solutions for $\bar{\mu} = 4\pi$ (blue solid line), $\bar{\mu} = \pi/4$ (red solid line) and $\bar{\mu} = \pi/64$ (green solid line). (b) Detail of the solution for $\bar{\mu} = 4\pi$ (solid line). The dashed lines indicate the theoretical slope $\pi/3$ at the crests.



Figure 5. The solution for $\bar{\mu} = 4\pi$ (solid line) and the same signal with the addition of a set-up (dashed line) to match the results by Dyachenko *et al.* (2016) (green dots) and Schwartz (1974) (red diamonds).

supercomputer was needed for the computations. This suggests that the present scheme can be regarded as a fast and accurate approximation of the singular solution (i.e. $\nu = 0$) described in Zhong & Liao (2018). With respect to this, figure 6 displays the comparison with the shallowest case considered in Zhong & Liao (2018), that is, $\bar{\mu} = 1.0053 \times 10^{-2}$. The solutions show overall good agreement (figure 6*a*), although the proposed iterative scheme slightly underestimates the crest height ($\epsilon = 0.4085$ against $\epsilon = 0.4150$) and is characterized by a rounded peak instead of a sharp profile (figure 6*b*). Again, we underline that this behaviour is a consequence of the fact that the proposed model represents solutions with $\nu > 0$.

We highlight that the number of iterations to reach the convergence drastically reduces far from the conditions of maximum wave amplitude. For example, the shallowest case (i.e. $\bar{\mu} = \pi/512$), with $\epsilon = 0.35$ and 60 000 modes, reaches the convergence in 137 iterations and takes about 102 seconds with 42 cores. Regardless, 8500 modes are enough to satisfy the condition (5.14) on the energy. In this latter case, the convergence is reached in the same number of iterations and takes just 2.4 seconds with 42 cores, and 55.7 seconds with a single core.



Figure 6. (a) The solution for $\bar{\mu} = 1.0053 \times 10^{-2}$ as predicted by the homotopy analysis method in Zhong & Liao (2018) (red solid lines) and by the present iterative method (black dashed lines). (b) Detail in a region close to the crest.

6.2. Comparisons with the existing theories

In this subsection, we consider highly nonlinear waves propagating in three different water regimes. Specifically, we set $\epsilon = 0.35$ and choose $\bar{\mu} = \pi/16$ (shallow-water regime), $\bar{\mu} = \pi/4$ (intermediate-water regime) and $\bar{\mu} = \pi$ (deep-water regime).

Figure 7 displays the evolution of the parameters F, B, \overline{Q} and C_m as functions of the iterations. The rate of convergence is slower in the shallow-water limit, while it becomes faster as the depth increases (see the behaviour of C_m in figure 7d). Generally, after a short transient, the parameters F, B and \overline{Q} reach a plateau and the subsequent iterations are necessary mainly for convergence of the highest modes of θ . The overall distributions of the mode amplitudes E_n at the convergence are displayed in figure 8. For all the cases, the amplitude of the highest modes is of order 10^{-15} . Incidentally, we observe that the amplitudes are all positive.

To better highlight the benefits of the proposed solution, we consider a comparison with fifth-order Stokes and cnoidal wave theories described in Fenton (1985, 1979). Figure 9 displays the solutions for the free-surface elevation η as predicted by the present model and by the above-mentioned theories. For $\bar{\mu} = \pi$ (deep water), figure 9(*a*) illustrates the comparison with the fifth-order Stokes solution (red dashed line). The overall match is very good, and the proposed solution practically coincides with the Stokes one. In this regime, the solution from the cnoidal wave theory is not available, since it is out of its range of validity. A comparison with both the theories is, to some extent, possible for $\bar{\mu} = \pi/4$ (intermediate water, figure 9*b*). In this case, the fifth-order cnoidal wave (dash-dot green line) is close to that predicted by the present model (solid black line), except for a slight set-down. On the contrary, the Stokes solution shows non-physical oscillations caused by its second-order component. Finally, figure 9*c* displays the shallow-water case (namely, $\bar{\mu} = \pi/16$). In this regime, the Stokes solution is out of its range of validity and therefore is not shown. On the contrary, the match between the present model and the cnoidal wave is very good, and the signals are practically overlapped.

A more addressing example of the reliability of the proposed solution is obtained by comparing the solutions for F and U_D with those predicted by the fifth-order Stokes and cnoidal wave theories. In particular, in figure 10, we consider three cases with decreasing values of ϵ . For $\epsilon = 0.35$ (figures 10*a*,*b*), each of the above-mentioned theories gives results in fairly good agreement with the present model in its own range of applicability, but diverges rapidly out of it. The overall behaviour is confirmed for smaller values of ϵ .



Figure 7. Convergence of the solution for $\epsilon = 0.35$ (fixed) and $\bar{\mu} = \pi/16$ (red lines), $\bar{\mu} = \pi/4$ (green lines) and $\bar{\mu} = \pi$ (blues lines).



Figure 8. Amplitude E_n of the modes at the convergence for $\epsilon = 0.35$ (fixed) and $\bar{\mu} = \pi/16$ (red line), $\bar{\mu} = \pi/4$ (green line) and $\bar{\mu} = \pi$ (blue line).

Specifically, the match with the present model improves in deep water for the Stokes wave solution and in shallow water for the cnoidal wave one, but in intermediate water, their singularities are almost unaffected by the decrease of ϵ . In this case, the most evident phenomenon is the shift of the singularities towards shallower depths caused by the decrease of the Ursell number. This behaviour is consistent with the bounds described in § 4.2 that predict the Stokes wave theory to hold true for $\hat{U}_r \ll 1$, and the cnoidal theory to be applicable for $\hat{U}_r > 1/4$.

Summarizing, the proposed iterative model provides an accurate global solution all over the regimes of propagation and displays an overall good match with the fifth-order shallowand deep-water theories in their respective regions of applicability.



Figure 9. Solutions for the free-surface elevation η for $\epsilon = 0.35$ (fixed) and (a) $\bar{\mu} = \pi$, (b) $\bar{\mu} = \pi/4$, and (c) $\bar{\mu} = \pi/16$. The black solid lines indicate the proposed solution, the red dashed lines denote the firth-order Stokes solutions, and the green dash-dotted lines represent the fifth-order cnoidal theory solutions.

6.3. An explicit formula for the wave celerity

The results described so far are obtained by assuming that the wavelength of the propagating wave is known. This allows for the evaluation of $\bar{\mu}$ and consequently of all the related spatial scales described in §§ 2 and 3. Nonetheless, in practical applications, the wave period T^* is generally known instead of the wavelength, therefore a way to deduce $\bar{\mu}$ from it is needed. The knowledge of an explicit expression for *F* as a function of μ and ϵ is the key point to obtain $\bar{\mu}$ from the wave period. We start from the dispersion relation

$$c^* = \frac{\omega^*}{k_0^*} \quad \Rightarrow \quad \omega = \mu F(\mu, \epsilon),$$
 (6.5)

where $\omega = \omega^* t_0^*$ is the angular wave frequency, i.e. $\omega^* = 2\pi/T^*$. Since t_0^* depends on z_0^* (and therefore on the wavelength), it is convenient to introduce the shallow-water scale $\bar{\omega} = \omega^* \sqrt{h_0^*/g^*}$ and $\epsilon = \bar{\epsilon} H(\mu)$, so that (6.5) becomes

$$\bar{\omega} = \sqrt{\bar{\mu} \tanh(\bar{\mu})} F\left(\tanh(\bar{\mu}), \frac{\bar{\epsilon}\bar{\mu}}{\tanh(\bar{\mu})}\right).$$
(6.6)

Here, $\bar{\epsilon}$ and $\bar{\omega}$ are known, while $\bar{\mu}$ has to be obtained by inversion. Since $1 \le F \le 1.3$ (the upper bound is attained for a solitary wave of maximum amplitude), the inversion can be performed by applying the bisection method over the segment $(\bar{\mu}_1, \bar{\mu}_2)$, where $\bar{\mu}_1 = (\bar{\omega}/1.3)^2$ and $\bar{\mu}_2 = (\bar{\omega}^2 + \bar{\omega}\sqrt{\bar{\omega}^2 + 4})/2$. For these bounds, we used the relations $\tanh(\bar{\mu}) \le 1$ and $\tanh(\bar{\mu}) \ge \bar{\mu}/(1 + \bar{\mu})$.

The final step is to provide an explicit expression for F as a function of μ and ϵ . Since an explicit solution is not available, we provide a formula obtained through a regression



Figure 10. Solutions for (a,c,e) F and (b,d,f) U_D , as predicted by the fifth-order theories for Stokes waves (red dashed lines; Fenton 1985) and cnoidal waves (green dash-dotted lines; Fenton 1979). The solid black lines are the outputs obtained through the proposed iterative scheme.

analysis by using about 300 sample points in the (μ, ϵ) -plane. This reads

$$F(\mu,\epsilon) = 1 + \epsilon^2 \left(\frac{a+b\mu+c\mu^2}{\epsilon+d\mu+e\mu^2} \right), \tag{6.7}$$

where

$$a = -15.2479\epsilon^{3} + 10.9266\epsilon^{2} - 3.0256\epsilon + 1.1712,$$

$$b = 26.9416\epsilon^{3} - 17.1964\epsilon^{2} + 5.6883\epsilon - 2.2527,$$

$$c = -18.2407\epsilon^{3} + 11.7909\epsilon^{2} - 4.4316\epsilon + 1.4138,$$

(6.8)

and

$$d = 141.0995\epsilon^{4} - 119.6816\epsilon^{3} + 38.0355\epsilon^{2} - 5.6719\epsilon + 0.3527, e = -119.3300\epsilon^{4} + 84.4075\epsilon^{3} - 19.0126\epsilon^{2} - 0.0593\epsilon + 0.3693.$$
(6.9)

The structure of (6.7) is conceived to represent the asymptotic behaviours of F in deep- and shallow-water regimes. Specifically, $F = 1 + O(\epsilon)$ for $\mu \ll 1$ as predicted by the cnoidal wave theory (Fenton 1979), while $F = 1 + O(\epsilon^2)$ for $\mu = O(1)$ as predicted by the Stokes theory (Fenton 1985).

Figure 11(*a*) displays the contour lines of (6.7) in the (μ, ϵ) -plane. Specifically, the solution is drawn in the domain $(\mu, \epsilon) \in [0, 1] \times [0, \epsilon_{max}(\mu)]$, where ϵ_{max} is given by (6.1). The plot illustrates clearly the increase of the Froude number for increasing amplitudes and for smaller values of μ . To verify the accuracy of the proposed formula, we consider the absolute relative errors with respect to the Stokes and cnoidal wave theories, that is,

$$\mathcal{E}_{stokes} = \left| 1 - \frac{F_{stokes}}{F} \right|, \quad \mathcal{E}_{cnoidal} = \left| 1 - \frac{F_{cnoidal}}{F} \right|, \quad (6.10a,b)$$

where *F* is computed using (6.7). These are displayed in figures 11(*b*,*c*), respectively. In both cases, the grey regions indicate the ranges where the above-mentioned theories are not applicable. Generally, the overall error maintains below 0.5% in the largest part of the range of validity for both the theories, and increases slightly close to the upper bound of the domain (namely, for ϵ close to ϵ_{max}), where it grows up to 1% for the Stokes solution (figure 11*b*) and up to 2% for the cnoidal wave solution (figure 11*c*). Incidentally, we highlight that in such a region, the accuracy of the fifth-order theories decreases rapidly, since $\epsilon \simeq 0.39$ and the truncation error becomes of order $\epsilon^6 \simeq 3 \times 10^{-3}$.

7. Conclusions

Starting from the integral equation of Byatt-Smith (1970), we defined an iterative scheme to solve the problem of steady periodic waves propagating over a planar seabed. The proposed approach allowed for the modelling of the wave dynamics from deep- to shallow-water conditions, and up to the limit of maximum-amplitude waves.

The outputs of the iterative schemes were compared with the existing fifth-order theories in shallow- and deep-water conditions, and with the results (both theoretical and numerical) for maximum-amplitude waves available in the literature. In all cases, the proposed model displayed overall good agreement, proving to be reliable and accurate. Further, the proposed iterative scheme can be employed in practical applications concerning wave modelling and numerical benchmarking, thanks to its limited computational cost.

The global description of the wave dynamics over the different regimes of motion was possible through the definition of a new scaling capable of identifying the vertical length representative of the wave motion. This also allowed for a compact description and presentation of the results provided in this work.

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Author ORCIDs.

Matteo Antuono https://orcid.org/0000-0002-9489-6922.



Steady periodic waves over a planar seabed

Figure 11. Contours of (*a*) formula (6.7), and absolute relative errors with respect to (*b*) the Stokes wave theory (Fenton 1985) and (*c*) the cnoidal wave theory (Fenton 1979). The grey regions indicate the ranges where the above-mentioned theories are not applicable.

Appendix A. The convolution operator $\mathscr C$

In this appendix, we show the derivation of the convolution operator \mathscr{C} and recall briefly some useful theoretical results. We define the Fourier transform and its inverse as

$$\mathscr{F}(f(x))(y) = \int_{\mathbb{R}} \overline{e}^{iyx} f(x) \, \mathrm{d}x, \quad \mathscr{F}^{-1}(f(x))(y) = \frac{1}{2\pi} \int_{\mathbb{R}} e^{iyx} f(x) \, \mathrm{d}x. \quad (A1a,b)$$

Consequently, the following results hold true for convolution integrals:

$$\mathscr{F}(f * g) = \mathscr{F}(f)\mathscr{F}(g), \quad \mathscr{F}^{-1}(f * g) = 2\pi\mathscr{F}^{-1}(f)\mathscr{F}^{-1}(g), \tag{A2a,b}$$

where the symbol * denotes the convolution operator. Now, let us consider the function

 $f(x) = \log \left(|\tanh \left(\alpha x \right)| \right) = \log \left(\tanh \left(\alpha |x| \right) \right), \tag{A3}$

where α is a positive real constant. Differentiating this expression, we find

$$f(x) = 2\alpha \operatorname{cosech}(2\alpha x). \tag{A4}$$

Then we multiply by x and apply the inverse Fourier transform, obtaining

$$\mathscr{F}^{-1}\left(x\dot{f}(x)\right)(y) = \frac{\gamma}{2}\operatorname{sech}^{2}(\gamma y),\tag{A5}$$

where $\gamma = \pi/(4\alpha)$ and the result on the right-hand side is demonstrated in Appendix D of Sinolakis (1986). Using the properties of the inverse Fourier transform, (A5) leads to the differential equation

$$y\frac{\mathrm{d}\mathscr{F}^{-1}(f)}{\mathrm{d}y} + \mathscr{F}^{-1}(f) + \frac{\gamma}{2}\operatorname{sech}^2(\gamma y) = 0, \tag{A6}$$

whose solution is

$$\mathscr{F}^{-1}(f)(y) = -\frac{\tanh(\gamma y)}{2y} + \frac{C_0}{y},\tag{A7}$$

where C_0 is a constant. Since the inverse transform of a real even function is a real even function, the constant C_0 has to be zero. Then applying the Fourier transform and using (A5), we find, finally,

$$\mathscr{F}\left(\frac{\tanh(\gamma y)}{y}\right)(x) = -2\log\left(\tanh\left(\frac{\pi|x|}{4\gamma}\right)\right),\tag{A8}$$

where $\alpha = \pi/(4\gamma)$ has been used. The above results also lead to the identities

$$-2\int_{\mathbb{R}}\sin(\xi x)\log\left(\tanh\left(\frac{\pi|x|}{4\gamma}\right)\right)dx = 0,$$

$$-2\int_{\mathbb{R}}\cos(\xi x)\log\left(\tanh\left(\frac{\pi|x|}{4\gamma}\right)\right)dx = 2\pi\frac{\tanh(\gamma\xi)}{\xi},$$
(A9)

where $\xi \in \mathbb{R}$. Now we introduce the convolution operator

$$\mathscr{C}[f] = f(\cdot) * \left[-\frac{1}{\pi} \log \left(\tanh \left(\frac{\pi |\cdot|}{4\gamma} \right) \right) \right].$$
(A10)

From the results obtained above, it immediately follows that

$$\mathscr{C}[e^{i\xi x}](y) = \frac{\tanh(\gamma\xi)}{\xi} e^{i\xi y} \text{ and } \mathscr{C}[1](y) = \gamma.$$
(A11*a*,*b*)

Hence the definition of the inverse operator is straightforward:

$$\mathscr{C}^{-1}[e^{i\xi x}](y) = \frac{\xi}{\tanh\left(\gamma\xi\right)} e^{i\xi y} \quad \text{and} \quad \mathscr{C}^{-1}[1](y) = \frac{1}{\gamma}.$$
 (A12*a*,*b*)

Note that, differently from \mathscr{C} , the inverse operator cannot be represented as a convolution integral.

Appendix B. The analytical expression for η_0

Let us assume that θ is a periodic even function with zero mean:

$$\theta = \sum_{n=1}^{\infty} E_n \cos(n\mu\Phi) \quad \text{and} \quad \sum_{n=1}^{\infty} n |E_n| < \infty.$$
(B1*a*,*b*)

Then, applying the operator \mathscr{C} and its inverse, we immediately obtain

$$\mathscr{C}[\theta] = \sum_{n=1}^{\infty} \frac{\tanh(n\mu\Psi_s)}{n\mu} E_n \cos(n\mu\Phi), \quad \mathscr{C}^{-1}[\theta] = \sum_{n=1}^{\infty} \frac{n\mu E_n}{\tanh(n\mu\Psi_s)} \cos(n\mu\Phi),$$
(B2*a*,*b*)

and consequently we find

$$\langle \mathscr{C}[\theta] \rangle = 0, \quad \langle \mathscr{C}^{-1}[\theta] \rangle = 0,$$
 (B3*a*,*b*)

where the symbol $\langle \cdot \rangle$ denotes the average over the wave period in the hodograph plane. More generally, we observe that for a generic periodic function f, the following relation holds true:

$$\langle f \rangle = 0 \quad \Leftrightarrow \quad \langle \mathscr{C}[f] \rangle = 0 \quad \Leftrightarrow \quad \langle \mathscr{C}^{-1}[f] \rangle = 0.$$
 (B4)

Now let us consider the condition (4.7). Applying the inverse operator \mathscr{C}^{-1} to (5.1), we obtain

$$\frac{\mathrm{d}\,x_s}{\mathrm{d}\Phi} = 1 + \epsilon\,\mathscr{C}^{-1}[\theta].\tag{B5}$$

Then substituting in (4.7) and recalling that $\langle \theta \rangle = 0$ and $\eta = \theta + \eta_0$, we obtain

$$\eta_0 = -\epsilon \ \langle \theta \ \mathscr{C}^{-1}[\theta] \rangle. \tag{B6}$$

Finally, substituting the expression for θ , we find

$$\langle \theta \, \mathscr{C}^{-1}[\theta] \rangle = \sum_{n=1}^{\infty} \frac{n\mu}{\tanh(n\mu\Psi_s)} \, \frac{E_n^2}{2} \ge 0. \tag{B7}$$

This implies that $\eta_0 \leq 0$, or equivalently, $\Psi_s \leq H$.

Appendix C. Details on the convolution terms

Let us assume that that both θ_m and $\Psi_s^{(m)}$ are known, and that the solution for θ_m is expressed through a cosine series with (N + 1) modes (including n = 0). This is written in the form

$$\theta_m = \sum_{|n| \le N} e_n^{(m)} \exp(in\mu\Phi), \quad \text{where } e_n^{(m)} = e_{-n}^{(m)} \text{ and } e_0^{(m)} = 0.$$
(C1)

Accordingly, the coefficients of the cosine series in (A1*a*,*b*) are $E_n^{(m)} = 2e_n^{(m)}$ and $E_0^{(m)} = e_0^{(m)}$. It is simple to verify that all the terms in the iterative models are, in fact, represented by cosine series. As a consequence, it is sufficient to compute the modes for $0 \le n \le N$ and derive the remaining ones by symmetry. For non-negative values of *n*, the coefficients

of the convolution terms are in the form of summations that span from the mode p = n - N to the mode p = N. In particular, the term S_m in (5.4) is given by

$$S_m = \sum_{|n| \le N} s_n^{(m)} \exp(in\mu\Phi), \text{ where } s_n^{(m)} = s_{-n}^{(m)},$$
(C2)

and the coefficients for $n \ge 0$ are

$$s_n^{(m)} = \mu^2 \sum_{p=n-N}^{N} p(n-p) \left[\frac{1}{\tanh\left(p\mu\Psi_s^{(m)}\right) \tanh\left((n-p)\mu\Psi_s^{(m)}\right)} - 1 \right] e_p^{(m)} e_{n-p}^{(m)}.$$
(C3)

Similarly, for $\theta_m S_m$ we obtain

$$\theta_m S_m = \sum_{|n| \le N} r_n^{(m)} \exp\left(in\mu\Phi\right), \quad \text{where } r_n^{(m)} = r_{-n}^{(m)}, \tag{C4}$$

and the coefficients for $n \ge 0$ are

$$r_n^{(m)} = \frac{1}{2} \sum_{p=n-N}^{N} [s_p^{(m)} e_{n-p}^{(m)} + e_p^{(m)} s_{n-p}^{(m)}].$$
 (C5)

In this latter case, a symmetric form has been used for the convolution. The last term to be computed is

$$\theta_m \mathscr{C}^{-1}[\theta_m] = \sum_{|n| \le N} q_n^{(m)} \exp\left(in\mu \Phi\right), \quad \text{where } q_n^{(m)} = q_{-n}^{(m)}, \tag{C6}$$

whose coefficients (in symmetric form) for $n \ge 0$ are

$$q_n^{(m)} = \frac{\mu}{2} \sum_{p=n-N}^{N} \left[\frac{p}{\tanh\left(p\mu\Psi_s^{(m)}\right)} + \frac{n-p}{\tanh\left((n-p)\mu\Psi_s^{(m)}\right)} \right] e_p^{(m)} e_{n-p}^{(m)}.$$
 (C7)

After all the coefficients of the convolution integrals are computed, the evaluation of χ_{m+1} , β_{m+1} , θ_{m+1} and $\eta_0^{(m+1)}$ through (5.7), (5.8), (5.11) and (5.9) is straightforward. For example, [[S]] and $\langle S \rangle$ are obtained immediately as

$$[[S_m]] = \sum_{|n| \le N} s_n^{(m)} \left[1 - (-1)^n \right], \quad \langle S_m \rangle = s_0^{(m)}.$$
(C8*a*,*b*)

Appendix D. The analytical expressions for $z(\Phi, \Psi)$ and $x(\Phi, \Psi)$

Substituting (B5) in the integral expression (3.16), we obtain

$$z(\Phi, \Psi) = -H + \Psi + \epsilon \int_{\mathbb{R}} \exp(-iy\Phi) \,\mathscr{F}^{-1}[\theta] \,\frac{\sinh(y\Psi)}{\sinh(y\Psi_s)} \,\mathrm{d}y,\tag{D1}$$

and, through the first equation in the system (3.14), we find

$$x(\Phi, \Psi) = x_0 + \Phi + i\epsilon \int_{\mathbb{R}} \exp(-iy\Phi) \mathscr{F}^{-1}[\theta] \frac{\cosh(y\Psi)}{\sinh(y\Psi_s)} \,\mathrm{d}y. \tag{D2}$$

If θ is given by the series (**B**1*a*,*b*), then we obtain

$$z(\Phi, \Psi) = -H + \Psi + \epsilon \sum_{n=1}^{\infty} E_n \frac{\sinh(n\mu\Psi)}{\sinh(n\mu\Psi_s)} \cos(n\mu\Phi),$$
(D3)

$$x(\Phi, \Psi) = x_0 + \Phi + \epsilon \sum_{n=1}^{\infty} E_n \frac{\cosh(n\mu\Psi)}{\sinh(n\mu\Psi_s)} \sin(n\mu\Phi),$$
(D4)

and finally,

$$x_s(\Phi) = x(\Phi, \Psi_s) = x_0 + \Phi + \epsilon \sum_{n=1}^{\infty} \frac{E_n}{\tanh(n\mu\Psi_s)} \sin(n\mu\Phi).$$
 (D5)

The solution for the velocity field in the moving frame of reference is obtained from (3.10a,b) and reads

$$U = \left[\left(\frac{\partial x}{\partial \Phi} \right)^2 + \left(\frac{\partial z}{\partial \Phi} \right)^2 \right]^{-1} \frac{\partial x}{\partial \Phi}, \quad W = \left[\left(\frac{\partial x}{\partial \Phi} \right)^2 + \left(\frac{\partial z}{\partial \Phi} \right)^2 \right]^{-1} \frac{\partial z}{\partial \Phi}.$$
 (D6*a*,*b*)

Finally, using (3.15), we obtain the values at the free surface:

$$U_s = \left(\frac{F^2 + 2\epsilon^2 B - 2\epsilon\eta}{F^2}\right) \frac{\mathrm{d}x_s}{\mathrm{d}\Phi}, \quad W_s = \epsilon \left(\frac{F^2 + 2\epsilon^2 B - 2\epsilon\eta}{F^2}\right) \frac{\mathrm{d}\eta}{\mathrm{d}\Phi}, \qquad (D7a,b)$$

where $dz_s/d\Phi = \epsilon d\eta/d\Phi$ has been used in the last equality.

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