WEAK GALAXY-GALAXY LENSING IN HST DATA

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Description

We expand on the work of Brainerd, Blandford & Smail (1996) (BBS), using a larger archival WFPC-2 dataset including many galaxy redshifts. It is clear from our data that the ellipticity distribution of images changes substantially with observed magnitude (figure 1, *left*) which is shown by simulations to be mainly the result of detection effects. We have detected a lensing signal, using a similar selection method to BBS and are also implementing a maximum likelihood method to constrain halo parameters. Using simulations, we show in figure 1 (*right*) that the signal is consistant with typical halo velocity dispersions of $\sigma^* \sim 70 - 100 \ km s^{-1}$ However, the radial extent of the halos is less well constrained.



Figure 1. left: Change in the ellipticity distribution with magnitude (solid, dotted, dashed & long dashed lines: I = 19 - 23, 23 - 24, 24 - 25, 25 - 26. right: Lensing signal from data (solid line) and simulations using $\sigma * = 100 \& 70 km s^{-1}$ (short & long dashes)

References

Brainerd, T.G., Blandgord, R.D. & Smail, I., 1996, ApJ., 466, 623.

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K. Sato (ed.), Cosmological Parameters and the Evolution of the Universe, 247. © 1999 IAU. Printed in the Netherlands.