

Atmospherically Processed and Stable Cs-Pb Based Perovskite Solar Cells

Shubhra Bansal, Michelle Chiu

Department of Mechanical Engineering, Center of Energy Research, University of Nevada Las Vegas, Las Vegas, Nevada 89154, U.S.A.

ABSTRACT

In this work, a planar heterojunction superstrate n-i-p device based on Zn(O,S) electron transport layer and CsPbI₂Br absorber material at 1.93 eV bandgap is presented. The CsPbI₂Br films are deposited using a 2-step atmospheric solution deposition process and characterized by X-ray diffraction (XRD), scanning electron microscopy (SEM), UV-vis spectroscopy and photoluminescence (PL). Best device with an efficiency of 12.34 % and 11.94% in reverse and forward scans respectively and stabilized power output of 12.14 mW/cm² has been demonstrated via atmospheric solution processing with minimal hysteresis between forward and reverse scans. The devices show voltage dependent current collection as well as light-dark crossover in forward bias. Light soaking tests at 65 °C and 1-sun at V_{oc}, resulted in open-circuit voltage and fill-factor degradation. Electroluminescence (EL) after 100 hours of light soaking shows a reduction in overall EL intensity as well as a shift in emission to lower wavelength. The devices exhibit a positive temperature coefficient of about 0.14 %/°C. It is found that Zn(O,S) is a viable alternative electron transport layer to replace TiO₂. By replacing methylammonium cation with cesium and addition of Br has improved the stability of the perovskite phase.

INTRODUCTION

Since the introduction of methylammonium lead halide based solar cells by Kojima *et al.* in 2009 [1], organic-inorganic hybrid perovskites have gained significant attention as an absorber layer in thin film solar cells. The characteristics exhibited by these materials such as high absorption coefficient [2], excellent transport properties [3], tunable bandgap [4], absence of deep trap states within the bandgap [5-6], low temperature processing etc. [7-9], make the materials suitable for low cost photovoltaic applications. Consequently, the photo-conversion efficiency of devices based on these materials has improved dramatically from 3.8% to 22.1% in a relatively short span of time [10-11]. Despite several advantages of perovskites, the application of these materials in commercial PV modules is seriously hindered by their stability. Perovskites have been shown to degrade rapidly upon exposure to heat, moisture, air, and light [4,8,12-14]. Limited success has been achieved in improving device stability by encapsulating devices [15,16]. The formation energy of MAPbX₃ has been reported to be very low, which makes these materials unstable at high temperatures and in humidity [17]. By replacing the organic cation with inorganic Cs, the intrinsic material stability can be improved [18]. Mixed A-cation [19-21] and mixed-halide systems have shown to be more structurally stable than MAPbI₃ due to increased formation energy and tolerance factors [22].

Conventionally used ETL and HTL materials are not completely immune to degradation either. TiO₂, which is a commonly used ETL material, has been known to be susceptible to UV induced degradation since late 1970s [23]. Upon exposure to UV radiation, oxygen desorption occurs creating oxygen vacancies (Ti³⁺ sites) on the surface of TiO₂ [24]. The surface energy levels resulting from changes in the electronic structure are found to be in the mid-gap region of the bulk

material. When perovskite absorber material gets photo-excited, electrons are injected into the conduction band of TiO_2 , where they become trapped into the mid-gap surface traps and readily recombine with holes. Leijtens et al. have reported performance degradation of TiO_2 based devices upon exposure to UV radiation [16, 25]. Improved device stability was observed using a UV filter and by replacing TiO_2 with Al_2O_3 [16]. Furthermore, the increased defect density at TiO_2 /perovskite interface has been shown to be associated with increased hysteresis effect in devices, as well [26]. Even though the degradation can be reduced using a UV filter, it is only achieved at a cost of photocurrent as well as complexity in device fabrication [16]. Both P3HT and Spiro-OMeTAD HTLs are also prone to degradation [13]. In the case of P3HT, the decrease in device performance is attributed to the photo-chemical degradation of the material [13] driven by UV, humidity and temperature. On the other hand, the decrease in V_{OC} and J_{SC} in the case of Spiro-OMeTAD observed by Matteocci et al. was attributed to changes in the oxidation state of Spiro-OMeTAD resulting in partial de-doping of the material [13]. Recently, attempts have been made to substitute the organic hole transport layers with CuSCN [27], NiO_x [28, 29], CuI [30], etc. and the electron transport layer TiO_2 with ZnO [29, 31]. You et al. [29] have reported a solar cell with ZnO as the ETL and NiO_x as the HTL with an efficiency of 16.1% and a significant improvement in device stability. Similarly, Huang et al. [32] have replaced TiO_2 with CdS , which significantly improved device stability, as compared to the case of TiO_2 , over a 12-hr period.

EXPERIMENT

Superstrate n-i-p devices were fabricated using a modified 2-step process [33]. Glass with F:SnO₂ transparent conducting oxide (TCO) was first cleaned in an ultrasonic bath and 50 nm thick Zn(O,S) was deposited via chemical bath deposition using a Zn-sulfate chemistry [34]. Zn(O,S) was deposited in a bath of zinc-sulfate, thiourea and ammonium hydroxide at 80 °C while stirring at 500 rpm. The substrates were then cleaned with 2.5% ammonium hydroxide solution and water followed by air drying. The sunny side of glass was cleaned with dilute HCl to remove Zn(O,S). Zn(O,S) substrates were then pre-heated at 70 °C and partial electrolyte treatment was done using a 0.5 M PbCl_2 solution in methanol. The PbCl_2 treatment is expected to improve the doping density of the Zn(O,S) and help with nucleation of the perovskite film [33, 35]. A 0.5 M precursor of CsI, PbI_2 , CsBr, PbBr_2 in N, N-Dimethylformamide (DMF) was prepared in dry air conditions at 70 °C with 4 hours stir time. The intended I/(I+Br) ratios in the precursor was 0.67. 100 μL of solution was then spin-coated on the 1" x 1" substrates at 2000 rpm for 45 seconds. 100 μL of chlorobenzene applied at 25 seconds to improve the morphology of the films [19]. To get uniform film thickness and to incorporate enough Br in the films, a spray treatment of CsBr and PbBr_2 solution in methanol was done at the end of spin coating cycle. The films were then annealed at 350 °C for 10 minutes in dry air conditions. It is noted that without the CsBr and PbBr_2 treatments, we obtained the dark brown-black CsPbI_3 phase at 350 °C which quickly turned into the yellow phase when cooled to room temperature. It has been observed that CsPbI_3 is stable in non-perovskite orthorhombic (yellow) phase at room temperature and changes to the cubic perovskite (black) phase when heated above 300 °C [36, 37]. A carbon paste Acheson Electrodog 456C in MEK was then spin coated to form the hole transport layer followed by thermal evaporation of Au. The front contact was formed by indium paste applied directly to the TCO by lightly scratching the Zn(O,S) layer. X-ray diffraction on films was done using PANalytical X'PERT Pro X-ray Diffraction Spectrometer. UV-Vis measurements were conducted using a Shimadzu UV-2600 plus spectrophotometer. Photoluminescence measurements were performed

using a 488 nm Argon/Krypton Stabilite 2018 ion laser at 100 mW/cm². Solar cell J-V characteristics were measured using ATLAS XXL+ solar simulator calibrated using a standard c-Si cell. Quantum efficiency measurements were made using PV Measurements Inc. QEXL system. All solar cells and perovskite thin films were subjected to light soaking at 1-sun and 65 °C using an ATLAS XXL+ light soak chamber with Xenon-arc lamps. SCAPS-1D [38] device model for CH₃NH₃PbI₃ solar cells validates the treatment as superstrate n-i-p devices as explained elsewhere [39].

RESULTS AND DISCUSSION

Figure 1a shows the absorption and photoluminescence (PL) spectra of CsPbI₂Br films. The absorption peak is observed at about 640 nm corresponding to a bandgap of 1.93 eV and PL shows variation in peak position between 625 nm to 640 nm, indicating presence of micro non-uniformities. For $x = 0.67$, the measured XRD spectrum (Figure 1b) agrees well with literature data for CsPbI₂Br for the (100) and (200) peak positions and a preferred orientation in (100) direction [35, 40]. The peak positions are shifted to high 2θ as compared to cubic CsPbI₃ due to lattice contraction resulting from Br substitution of I. As previously observed (110) and (111) peaks are negligible intensity for CsPbI₂Br perovskite phase. To check the stability of the films under solar cell operating conditions, thin films were subjected to 1-sun and 65 °C anneal conditions. XRD pattern for as deposited, after 50 and 100 hours anneal are shown in Figure 1b. XRD data shows onset of peak at 20° indicating formation of CsPbI₃ after 100 hours of light soaking, resulting from halide segregation or non-uniformities. Hoke et al. [41] have reported segregation of iodine rich domains with lower bandgap and increased sub-bandgap absorption leading to red shift in PL. CsPbI₂Br has been shown to be stable at temperatures more than 85 °C [35], therefore, we believe the perovskite phase is intrinsically stable and further process and crystal control is needed to obtain long-term stability. Results from six devices are summarized in Table 1 including the best device at 12.34% power conversion efficiency and a stabilized power output (SPO) of 12.14 mW/cm². The SPO was measured by holding the device at 1.1 V under standard test conditions (STC). We could obtain a tight distribution in device efficiency indicating the robustness of the process. All devices were measured in ambient conditions (< 25% RH) without encapsulation.

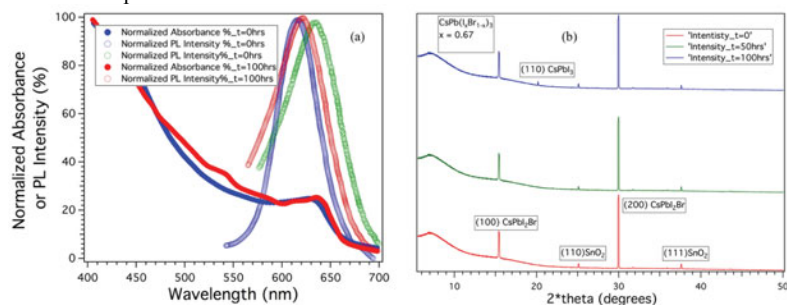


Figure 1. (a) Normalized absorbance and PL data for CsPbI₂Br films at 0 and 100 hours. PL measurements at 0 hours indicate regions of micro-nonuniformities in composition. (b) XRD scans on CsPbI₂Br after 0, 50 and 100 hours anneal at 1-sun and 65 °C.

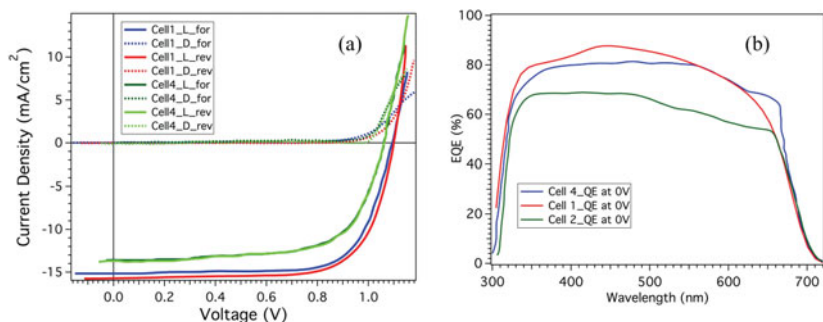


Figure 2. (a) J-V scans in forward and reverse bias for 5 mm x 5 mm CsPbI₂Br devices. Both cells indicate voltage dependent current collection in light curves and light-dark crossover in forward bias. (b) Experimental quantum efficiency for CsPbI₂Br devices indicates losses in longer wavelength region likely due to incomplete carrier collection that may be attributed to non-uniform perovskite thickness.

Table 1. Cell parameters for six Zn(O,S)/ CsPbI₂Br devices

Cell #	V _{oc} (V)	J _{sc} (mA/cm ²)	FF (%)	Eff (%)	Stabilized P _{max} (mW/cm ²)	QE J _{sc} (mA/cm ²)	est. Eff (%) w/ QE J _{sc}	V _{oc} at 65 °C	Estimated Temperature Coefficient (%/°C)
1	1.1	15.7	71.43	12.33	12.14	14.38	11.3	1.165	0.15
2	1.245	10.1	72.05	9.06	9.11	9.96	8.94	1.3	0.12
3	1.07	13.22	68.4	9.66	9.54	12.68	9.28	1.123	0.12
4	1.063	13.68	69.91	10.23	10.4	12.75	9.54	1.123	0.14
5	1.09	14.85	70.02	11.3	11.33	14.52	11.08	1.15	0.14
6	1.12	15.05	70.88	11.95	12.01	14.21	11.28	1.188	0.15

As shown in Figure 2a, the best device had an efficiency of 12.34% in a reverse scan and 11.94% in forward scan with V_{oc} of 1.1 V and 1.094 V respectively. Light generated current (J_L) during forward scan is 15.15 mA/cm² and 15.7 mA/cm² for the reverse scan. Further, we observed some voltage dependent collection in the light curves as well as light-dark crossover in forward bias. Due to the voltage dependent collection, J_L is different from J_{sc}, which agrees well with the QE estimated J_{sc}. All values for the six devices are reported in Table 1. The best device was held at a constant bias of 1.1 V to yield a stabilized power output of 12.14 mW/cm², which is reasonable as minimal hysteresis is observed between forward and reverse scans. The efficiency reported here shows an improvement over two previous reports of CsPbI₂Br devices [35,40] which may be attributed to improved J_{sc} and FF. Our devices use F:SnO₂/Zn(O,S) as well as a PbCl₂ pre-treatment before perovskite deposition among key differences. The V_{oc} deficit is high which is being addressed by tailoring of Zn(O,S) /perovskite interface, thickness control of perovskite layer and a better HTL such as NiO or CsSnI₃. As shown in the focused ion beam cross-section we see potential leakage paths which may lead to weak diodes and decrease in V_{oc} because of non-uniform deposition of Zn(O,S). It can also be seen from FIB-SEM images that thickness and film quality of CsPbI₂Br needs further improvement (Figure 3c).

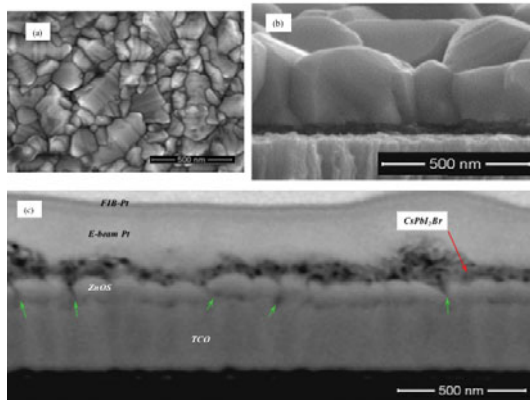


Figure 3. (a) Top-down SEM image for solution processed CsPbI₂Br film on SnO₂ (b) Cross-sectional view of cleaved sample of CsPbI₂Br film on SnO₂ (c) FIB-SEM cross-sectional image of TCO/ Zn(O,S)/ CsPbI₂Br film stack. Perovskite layer appears not fully crystallized and Zn(O,S) film shows pin-holes.

Three of the devices were subjected to light soak testing at 1-sun (1000 W/m²) and 65 °C for 100 hours under ambient conditions and in-situ JV measurements were made at regular intervals. The V_{oc} of the device measured at 65 °C was higher than measured at 25 °C indicative of a positive temperature coefficient of these materials. It has been reported bandgap of perovskite materials increases with temperature [42]. The open circuit voltage at 65 °C and estimated temperature coefficients are listed in Table 1. Figure 4 shows degradation data of cell parameters when the devices were stressed at 1-sun, 65 °C and open-circuit voltage conditions. The devices show degradation in air due to development of shunts and significant decrease in V_{oc} and FF.

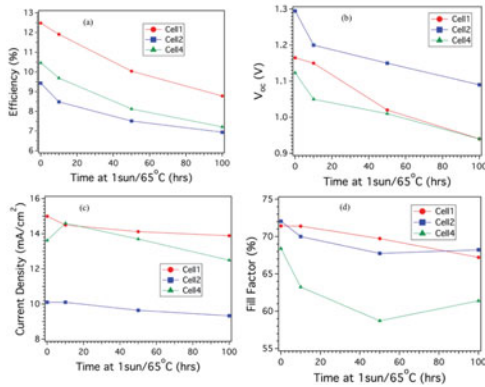


Figure 4. Degradation data of cell parameters (a) efficiency (b) open-circuit voltage (c) fill factor (d) current density for three devices stressed at 1-sun, 65 °C and open-circuit voltage conditions.

The shunts formation may be due to disintegration of CsPbI₂Br into CsPbI₃ or PbI₂ as indicated by Sutton et al. [43]. Another speculation is the probes used for J-V measurements may cause scratching and hot-spot formation. Electroluminescence (EL) images on Cell 4 before and

after 100 hours at 1-sun and 65 °C are shown in Figure 5. Images were taken at a current density of 60 mA/cm² equivalent to about 4 suns using a 0.5x objective. EL images indicate lower intensity consistent with appearance of shunts and lower V_{oc} with stress.

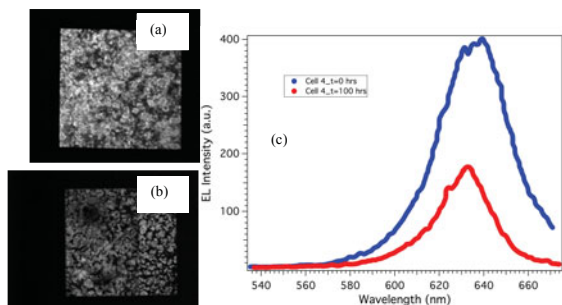


Figure 5. Electroluminescence images of Cell 4 (a) before and (b) after light soak test. (c) Emission shifts from 640 nm to 633 nm after 100 hours of light soaking test.

CONCLUSION

Here we have demonstrated inorganic planar perovskite devices with power conversion efficiency of 12.34% and stabilized maximum power of 12.14 mW/cm², which represents highest performance reported yet for CsPbI₂Br perovskite solar cells. These inorganic lead halide perovskites show a positive temperature coefficient, which needs further investigation. Positive temperature coefficient can possibly make these materials extremely attractive for high solar cell operating temperatures. The bandgap demonstrated here of 1.93 eV is a great candidate for high efficiency tandem device structures with devices such as c-Si and CIGS [44, 45]. The structural stability, uniform deposition, microstructural control, micro non-uniformities are still open questions which need further fundamental understanding. We have also demonstrated Zn(O,S) as an alternate electron transport material to address the stability issues of TiO₂. Future research is needed to optimize the Zn(O,S) process to obtain extremely uniform thin films for improved V_{oc} and J_{sc} . Atomic layer deposition of Zn(O,S) has been demonstrated at manufacturing scale by thin film manufacturers, making it a viable solution for substrate configuration needed for tandem device structures.

REFERENCES

1. A. Kojima, K. Teshima, Y. Shirai, T. Miyasaka, *J. Am. Chem. Soc.*, 131, 17, 6050 (2009).
2. N.-G. Park, *Mater. Today*, 18, 65 (2015).
3. S. D. Stranks, G. E. Eperon, G. Grancini, C. Menelaou, M. J. Alcocer, T. Leijtens, L. M. Herz, A. Petrozza, H. J. Snaith, *Science*, 342, 6156, 341 (2013).
4. J. Noh, S. H. Im, J. H. Heo, T. N. Mandal, S. I. Seok, *Nanoletters*, 13, 1764 (2013).
5. W-J Yin, T. Shi, Y. Yan, *Appl. Phys. Lett.*, 104, 063903 (2014).
6. J. Kim, S. H. Lee, K. H. Hong, *J. Phys. Chem. Lett.*, 5, 8, 1312 (2014).
7. W.S. Yang, J. H. Noh, N. J. Jeon, Y. C. Kim, S. Ryu, J. Seo, S. Seok, *Science*, 348, 1234 (2015).
8. J. Burschka, N. Pellet, S. J. Moon, R. H-Baker, P. Gao, M. K. Nazeeruddin, M. Grätzel, *Nature*, 499, 7458, 316 (2013).

9. H. J. Snaith, *J. Phys. Chem. Lett.*, 4, 21, 3623 (2013).
10. NREL efficiency chart http://www.nrel.gov/ncpv/images/efficiency_chart.jpg (Accessed 23 April 2017).
11. M. Green, K. Emery, Y. Hishikawa, W. Warta, E. D. Dunlop, *Prog. Photovol: Res. Appl.*, 24, 1, 3 (2016).
12. T. Leijtens, G. E. Eperon, N. K. Noel, S. N. Habisreutinger, A. Petrozza, H. J. Snaith, *Adv. Energy Mater.*, 5, 1500963 (2015).
13. F. Matteocci, S. Razza, F. D. Giacomo, S. Casaluci, G. Mincuzzi, T. M. Brown, A. D'Epifanio, S. Licocchia, A. D. Carlo, *Phys. Chem. Chem. Phys.*, 16, 3918 (2014).
14. A. B. Djuricic, F. Liu, A. M. C. Ng, Q. Dong, M. K. Wong, A. Ng, C. Surya, *Phys. Status Solidi RRL*, 10, 4, 281, (2016).
15. Y. Han, S. Meyer, Y. Dkhissi, K. Weber, J. M Pringle, U. Bach, L. Spiccia, Y. B. Cheng, *J. Mat. Chem. A*, 3, 8139 (2015).
16. T. Leijtens, G. E. Eperon, S. Pathak, A. Abate, M. M. Lee, H. J. Snaith, *Nat. Comm.*, 4, 3885 (2013).
17. D. B. Mitzi, *Prog. Inorg. Chem.*, 48, 1 (2007).
18. A. Amat, E. Mosconi, E. Ronca, C. Quarti, P. Umari, M. K. Nazeeruddin, M. Grätzel, and F. De Angelis, *Nano Lett.*, 14, 3608 (2014).
19. Z. Li, M. Yang, J-S. Park, S-H. Wei, J. J. Berry, K. Zhu, *Chem. Mater.*, 28, 1, 284 (2016).
20. D. P. McMeekin, G. Sadoughi, W. Rehman, G. E. Eperon, M. Saliba, M. T. Hörlantner, A. Haghighirad, N. Sakai, L. Korte, B. Rech, M. B. Johnston, L. M. Herz, H. J. Snaith, *Science*, 351, 6269, 151 (2016).
21. C. Yi, J. Luo, S. Meloni, A. Boziki, N. A-Astani, C. Grätzel, S. M. Zakeeruddin, U. Röthlisberger, M. Grätzel, *Energy Environ. Sci.*, 9, 656 (2016).
22. W. Travis, E. N. K. Glover, H. Bronstein, D. O. Scanlon, R. G. Palgrave, *Chem. Sci.*, 7, 4548 (2016).
23. G. Munura , V. Rives-Arna , and A. Saucedo , *J. Chem. Soc., Faraday Trans. 1*, 75 , 736 (1979).
24. M.J. Llansola-Portoles, J.J. Bergkamp, D. Finkelstein-Shapiro, B. D. Sherman, G. Kodis, N. M. Dimitrijevic, D. Gust, T. A. Moore, and A. L. Moore, *J. Phys. Chem. A*, 118, 10631 (2014).
25. S.K. Pathak, A. Abate, T. Leijtens, D. J. Hollman, J. Teuscher, L. Pazos, P. Docampo, U. Steiner, H. J. Snaith, *Adv. Energy Mater.*, 8, 1301667 (2014).
26. J.M. Azpiroz, E. Mosconi, J. Bisquert and F. D. Angelis, *Energy Environ. Sci.*, 8, 2118 (2015).
27. P. Qin, S. Tanaka, S. Ito, N. Tetreault, K. Manabe, H. Nishino, M.K. Nazeeruddin and M. Grätzel, *Nature Commun.*, 5, 3834 (2014).
28. Z. Zhu, Y. Bai, T. Zhang, Z. Liu, X. Long, Z. Wei, Z. Wang, L. Zhang, J. Wang, F. Yan, S. Yang, *Angew. Chem. Int. Ed.*, 53, 12571 (2014).
29. J. You, L. Meng, T-B. Song, T-F. Guo, Y. Yang, W-H. Chang, Z. Hong, H. Chen, H. Zhou, Q. Chen, Y. Liu, N. D. Marco, Y. Yang, *Nat. Nanotech.*, 11, 75 (2016).
30. J. A. Christians, R. C. M. Fung, P. V. Kamat, *J. Am. Chem. Soc.*, 136, 2, 758 (2014).
31. D. Y. Liu and T. L. Kelly, *Nat. Photon.*, 8, 2, 133 (2014).
32. I. Huang, K. Yong, *ACS Appl. Mater. Inter.*, 8, 4226 (2016).
33. M. Kulbak, D. Cahen, G. Hodes, *J. Phys. Chem. Lett.*, 6, 2452 (2015).
34. L. Qi, G. Mao, J. Ao, *Appl. Surf. Sci.*, 254, 18, 5711 (2008).

35. R. J. Sutton, G. E. Eperon, L. Miranda, E. L. Parrott, B. A. Kamino, J. B. Patel, M. T. Hörantner, M. B. Johnston, A. S. Haghighirad, D. T. Moore, H. J. Snaith, *Adv. Energy Mater.*, pp. 1502458 (2016).
36. C. K. Moller, *Nature*, 182, 1436 (1958).
37. S. Sharma, N. Weiden, A. Weiss, *Z. Phys. Chem.*, 175, 63 (1992).
38. A. Niemegeers and M. Burgelman, *Proc. 25th IEEE Photovoltaic Spec. Conf.*, Washington DC, IEEE, 901 (1996).
39. S. Bansal, P. Aryal, *Proc. 43rd IEEE Photovoltaic Spec. Conf.*, Portland, OR (2016).
40. R. E. Beal, D. J. Slotcavage, T. Leijtens, A. R. Bowring, R. A. Belisle, W. H. Nguyen, G. F. Burkhard, E. T. Hoke, M. D. McGehee, *J. Phys. Chem. Lett.*, 7, 746 (2016).
41. E. T. Hoke, D. J. Slotcavage, E. R. Dohner, A. R. Bowring, H. I. Karunadasa, M. D. McGehee, *Chem. Sci.*, 6, 613 (2015).
42. O. Dupré, R. Vaillon, M. A. Green, *Sol. Energy Mat. Sol. Cells*, 140, 92 (2015).
43. R. J. Sutton, G. E. Eperon, L. Miranda, E. S. Parrott, B. A. Kamino, J. B. Patel, M. T. Hörantner, M. B. Johnston, A. A. Haghighirad, D. T. Moore, H. J. Snaith, *Adv. Energy Mat.*, 6, 8 (2016).
44. M. H. Futscher and B. Ehrler, *ACS Energy Lett.*, 1, 4, 863 (2016).
45. T. Todorov, T. Gershon, O. Gunawan, Y. S. Lee. C. Sturdevant, L.-Y. Chang, S. Guha, *Adv. Energy Mat.*, 5, 23, 1500799 (2015).