

Electrical resistivity measurements on Ice Stream B, Antarctica

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ABSTRACT. Electrical resistivity sounding using the four-electrode Schlumberger array was carried out at station UpB on Ice Stream B to an electrode spacing of 3 km. Measured apparent resistivities were compared with theoretical models based on known relations between resistivity, density and temperature. Densities were measured in a pit and two coreholes; temperatures were measured in the upper 200 m of the ice stream and have been calculated for greater depth from an ice-stream temperature model. The resistivity, after correction for density and temperature, increases with depth down to 650–700 m. Below that is a marked decrease over the next 100 m or so that we correlate with the Holocene–Wisconsin transition zone. Still deeper there is an orders-of-magnitude increase to a value, in the basal ice, of 30 M Ω m or more. This extremely high resistivity is similar to that reported for temperate glaciers and deep in the Antarctic ice sheet elsewhere. We attribute it to the destruction, by extensive metamorphism, of impurity-conduction paths at two-grain boundaries.

INTRODUCTION

A geophysical and glaciological program to study the West Antarctic ice streams that drain into the Ross Ice Shelf was begun in 1983, with initial studies on Ice Stream B. Upstream B (UpB) camp, in the middle of Ice Stream B2, one of the main branches of Ice Stream B (Fig. 1), was established in November 1983.

Geophysical experiments conducted at and near UpB during the 1983–84 season included electrical resistivity sounding. Previous resistivity measurements at station RI, located in the flow band of Ice Stream E on the Ross Ice Shelf (Bentley, 1976; Shabtaie and Bentley, 1979, 1984), had revealed the existence of highly resistive ice at the base of the ice shelf with resistivity comparable to that reported for temperate glaciers (e.g. see Röthlisberger and Vögtli, 1967; Glen and Paren, 1975). Resistivity measurements were conducted on Ice Stream B particularly to obtain more information about the nature of this anomaly. Electrode separations covered a large enough range (2–3 km) to study the ice stream throughout its thickness.

The field methods, data analysis and modeling techniques are similar to those previously described (Bentley, 1977; Shabtaie and Bentley, 1979, 1984, in press a); we will not discuss them here. Here, we first calculate the resistivity as a function of depth in the ice, $\rho(z)$, to be expected from the variations of temperature and density. Then, we calculate the corresponding apparent resistivities, ρ_a , and find that they do not match the observations well. We find that we must introduce additional parameters, representing other physical and chemical changes in the ice, to find a satisfactorily fitting model.

MEASUREMENTS OF DENSITY AND TEMPERATURE

Density

Densities were measured on a 21 m core that was extracted at UpB camp near the resistivity profile during the 1983–84 field season. Cores were cut into pieces 10–20 cm long, and the densities (Fig. 2) were calculated by measuring the length, diameter and the weight of the individual pieces. To measure densities in the uppermost layers, density-tube samples were taken from the side walls of a 2.5 m pit at different depths and weighed. Each plotted density (Fig. 2) from the pit is the mean of four measurements at each depth. For densities at depths greater than 21 m, we have taken a smooth curve through spot measurements on a 100 m core from a hole drilled near the center of the resistivity profile during the 1984–85 field program (Alley and Bentley, 1988).

Temperatures

Temperature measurements were made in the pit and the 21 m corehole on 16 and 17 January, just before and after the resistivity measurements. Temperatures in the pit were read from pre-installed thermometers; those in the 21 m corehole were measured at intervals of 0.5–1 m using a Thermometric glass thermistor sensor. The temperature profile (Fig. 3a) shows the expected mid-summer shape.

During December 1983, a string of 15 thermistor sensors was emplaced at depths between 25 and 200 m in a hot-water drillhole. Temperature measurements (Fig. 3b) were made in January 1985, 13 months after

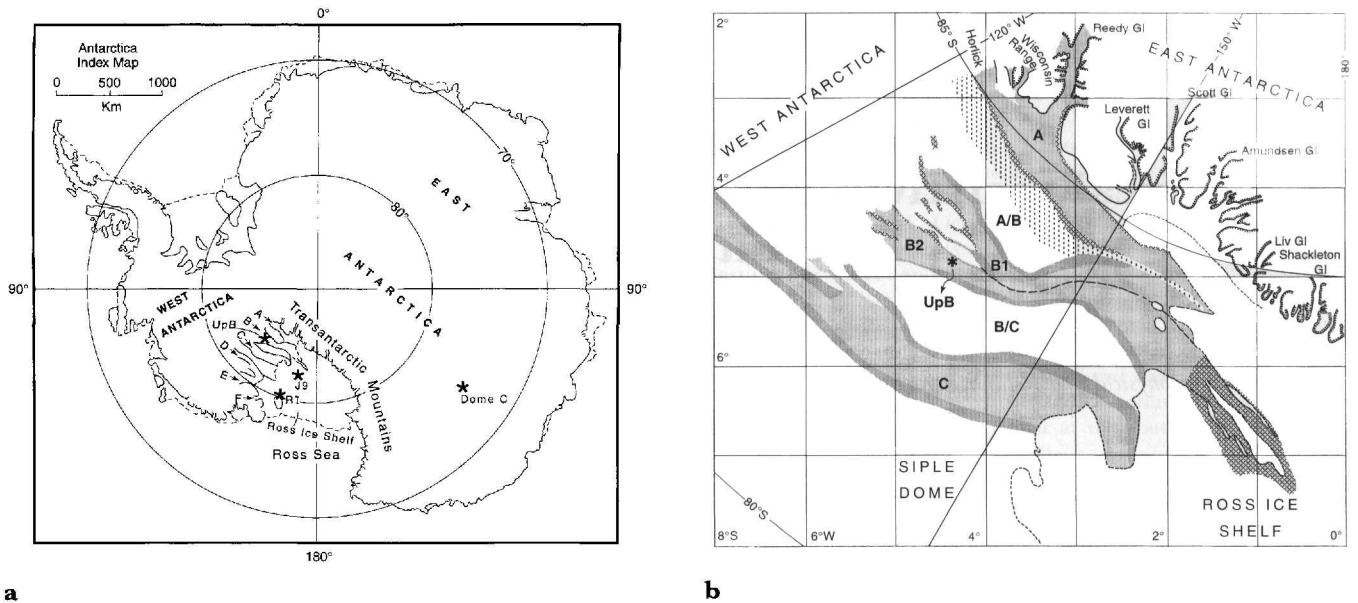


Fig. 1. a. Antarctica, showing the resistivity stations discussed in the text (reproduced from Drewry (1983) with modification). b. Map of West Antarctic Ice Streams A, B and C showing the location of UpB camp. Hachured strips denote zones of concentrated crevasses. The short-dashed line indicates the grounding line (junction with the Ross Ice Shelf). The long-dashed line shows the boundaries between two ice streams. For more detail, see Shabtaie and Bentley (1988).

the installation of the sensors (long enough to eliminate transient temperature changes). The temperatures show no negative gradient below the winter cold wave in the upper 10 m, which implies that the advection of cold ice from upstream is insignificant.

Temperatures deeper than 200 m were calculated from the model discussed next.

TEMPERATURE MODELING

Basal boundary condition

The temperature gradient, G_g , that would be produced at the base of a stationary ice sheet by the geothermal flux, q_g , in the absence of melting is

$$G_g = \frac{q_g}{\theta_c} \tag{1}$$

where θ_c ($7 \times 10^7 \text{ J m}^{-1} \text{ a}^{-1} \text{ }^\circ\text{C}^{-1}$) is the thermal conductivity of ice. For moving ice, the contribution from basal shear heating must be added to the geothermal heat to obtain the basal temperature gradient, G_b ; then, still with no melting,

$$G_b = G_g + \tau_b V_b / \theta_c \tag{2}$$

where τ_b is basal shear stress and V_b is the sliding velocity. (The shear-heating term in Equation (2) is independent of how the basal shear stress is accommodated by the ice sheet and its bed, so long as the heat is released near the interface.) We assume that the basal shear stress is equal to the driving stress, τ , and that V_b is equal to surface velocity, V . These assumptions maximize both τ_b and V_b , and therefore maximize our calculated rate of strain heating. The driving stress is given by

$$\tau = \bar{\gamma}_i g H \sin \bar{\alpha} \tag{3}$$

where $\bar{\gamma}_i$ is the mean ice-column density, g is the acceleration of gravity, H is the ice thickness and $\bar{\alpha}$ is the surface slope averaged over a distance of $20H$ to eliminate the effect of longitudinal stress gradients.

At Byrd Station, the assumption of a negligible basal melt rate is acceptable. The value of G_b measured in the corchole there was $0.0325^\circ\text{C m}^{-1}$ (Gow and others, 1968) and also $\tau_b V_b / \theta_c = 0.0065^\circ\text{C m}^{-1}$ (Budd and others, 1971). Equation (2) then gives $G_g = 0.026^\circ\text{C m}^{-1}$. This is close to the value of G_g , $0.023^\circ\text{C m}^{-1}$, given by Equation (1) for a global mean geothermal flux of $1.6 \times 10^6 \text{ J m}^{-2} \text{ a}^{-1}$. Lacking any better value, we initially assume $G_g = 0.026^\circ\text{C m}^{-1}$ also at UpB.

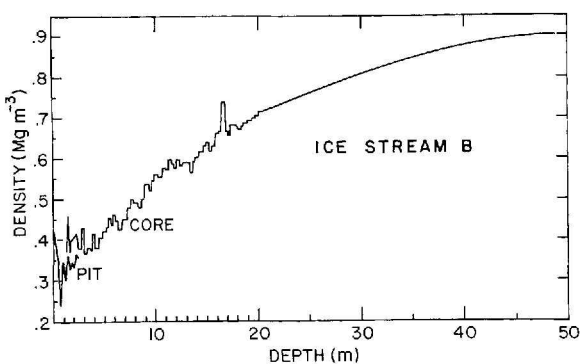


Fig. 2. Density versus depth at UpB. The 1983–84 pit and core densities are indicated. The smooth curve at depths greater than 20 m shows mean densities from the 1984–85 corehole.

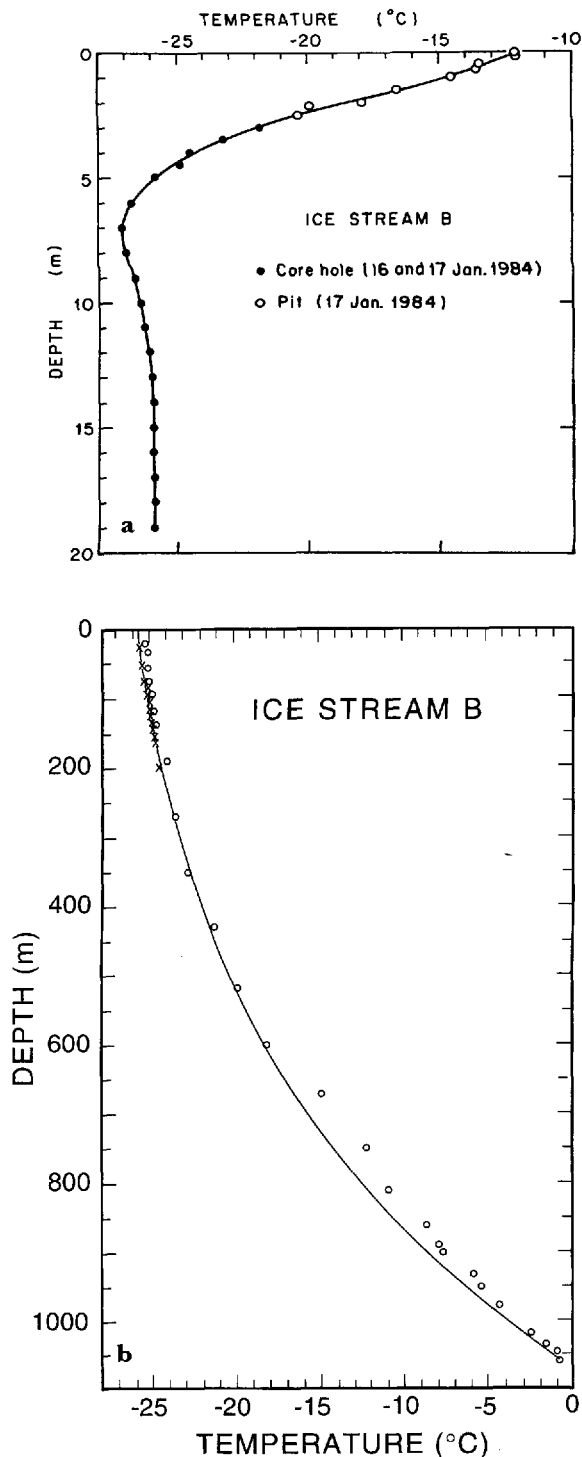


Fig. 3. Temperature versus depth at UpB. a. Temperatures in the upper 20 m measured in the pit and shallow corehole at UpB. b. Temperatures from two different hot-water drilled holes. Crosses show the measurements in the 1983-84 200 m hole; circles show measurements in a hole drilled to the bed (personal communication from H. Engelhardt, 1993). The solid line is the temperature model described in the text.

For an ice stream, the basal melt rate, \dot{b}_H , is not negligible, so Equation (2) is not valid; instead,

$$\dot{b}_H = \frac{1}{\gamma_i L} [V_b \tau_b + \theta_c (G_g - G_b)] \quad (4)$$

where L (335 J g^{-1}) is the specific latent energy of fusion and γ_i (0.917 Mg m^{-3}) is the density of ice. Equation (4) is a simple generalization of an equation given by Weertman (1961). G_b and \dot{b}_H are both unknown and cannot be estimated independently, so we treat \dot{b}_H as an adjustable parameter in the heat equation. At UpB, $V = 440 \text{ m a}^{-1}$ (Whillans and others, 1987), $H = 1060 \text{ m}$, $\alpha = 2.4 \times 10^{-3}$ (Shabtaie and others, 1987) and $\bar{\gamma}_i = 0.90$ (calculated from Figure 2); then $\tau_b = \tau = 22 \text{ kPa}$, from Equation (3).

Ice-stream temperature model

The equation of heat conduction in an ice sheet is (see Paterson, 1981, chapter 10):

$$\frac{\partial T}{\partial t} = \theta_d \frac{\partial^2 T}{\partial z^2} - V \frac{\partial T}{\partial x} - W \frac{\partial T}{\partial z} + 2\dot{\epsilon}_{xy} \tau_{xy} / \gamma_i C \quad (5)$$

where T is the temperature, t is time, θ_d is the thermal diffusivity, z is the vertical direction (positive downward), x is the horizontal direction along the flow, y is the horizontal direction transverse to flow, W is the vertical velocity of an ice particle, $\dot{\epsilon}_{xy}$ and τ_{xy} are the xy components of strain rate and shear stress, respectively, and C is the specific-heat capacity of the ice. In solving Equation (5), we assume: (a) the firn layer can be ignored, (b) horizontal temperature gradients are negligible and (c) internal strain heatings are small compared with the basal shear heating. Furthermore, we can neglect the horizontal advection term ($\partial T / \partial x$) at UpB, because there is such a small regional difference in mean annual 10 m temperatures (-26° to -27°C , 300 km downstream of UpB on the ice shelf (Thomas, 1976) compared with -24° to -26°C , 200-300 km upstream (Bentley and others, 1964)) and because, as already mentioned, our temperature measurements (Fig. 3b) show that advection is not significant.

The vertical velocity (W) can be approximated by (Weertman, 1961)

$$W = \dot{b}_0 + \frac{z}{H} (\dot{b}_H - \dot{b}_0) \quad (6)$$

where \dot{b}_0 is the surface balance rate (9 cm ice a^{-1} at UpB (Shabtaie and Bentley, 1987)). The temperature at the bed is given by the pressure-melting temperature, T_m , which can be found from the Clausius Clapeyron relationship (Robin, 1976) $T_m \approx \mu p$, where $\mu = -7.4 \times 10^4 \text{ }^\circ \text{C m}^2 \text{ N}^{-1}$ and $p = \bar{\gamma}_i g H$ is the hydrostatic pressure. For UpB, $T_m = -0.7^\circ \text{C}$.

Equation (5) can be solved numerically for steady state ($\partial T / \partial t = 0$). Our procedure was to assume different values of \dot{b}_H , from which G_b and W could be calculated from Equations (4) and (6), respectively. The modeled temperature-depth curves were then matched against the measured temperatures in the upper 200 m of the ice stream. We found for the best-fitting model $\dot{b}_H = 0.07 \text{ m a}^{-1}$ (the fit worsens markedly outside the range of $0.05 < \dot{b}_H < 0.1$) and $G_b = 0.05^\circ \text{C m}^{-1}$ (Fig. 3b). The modeled melt rate is twice that, 0.032 m a^{-1} , generated by the shear-heating term in Equation (4), even though that value is an upper limit. That could indicate an elevated geothermal flux or other unrecognized heat source, or it could reflect a violation of the

assumption of steady state or of assumptions (b) or (c), all of which are questionable for such a high-energy system as an ice stream. For resistivity modeling, fortunately, why $T(z)$ has a particular shape is irrelevant.

Recent deep drilling at UpB has made it possible to obtain temperatures all the way to the bed (Engelhardt and others, 1990). The latest measurements (personal communication from H. Engelhardt, 1993) are plotted in Figure 3b. There is satisfactory agreement for our purposes between modeled and measured temperatures, although there is a slight offset between them deeper than 600 m and the measured basal temperatures show fluctuations larger than would be expected so deep in the ice sheet. We have not deemed it fruitful to re-do our resistivity modeling on the basis of the measured temperatures, as we are concerned with effects far larger than those caused by the small temperature differences.

RESISTIVITY

In an ice sheet, density, temperature, impurities, crystal size and other chemical and physical properties change with depth and can or might affect the resistivity. The two parameters whose effects upon $\rho(z)$ are known are those whose determination we have already discussed, temperature and density. We will try to evaluate the effects of other factors from the measured resistivities.

The effect of temperature on the resistivity of both firn and solid ice is governed by the Arrhenius law with activation energy E :

$$\frac{\rho(z)}{\rho_0} = \exp\left[\frac{E}{k}\left(\frac{1}{T_K(z)} - \frac{1}{T_0}\right)\right] \tag{7}$$

where $\rho_0 = \rho(0)$ is the resistivity of the surface layer with temperature $T_0 = T_K(0)$, $k = 8.6 \times 10^{-5} \text{ eV K}^{-1}$ (Boltzmann’s constant) and T_K is the absolute temperature. An activation energy of 0.25 eV for firn and solid ice was determined from resistivity and temperature measurements at Dome C, on the East Antarctic plateau (Shabtaie and Bentley, in press a). That value also works well at UpB, as is shown by the modeling discussed below.

Firn can be modeled as a mixture of two components, ice and air, with volume fractions ν_1 and ν_2 , where $\nu_1 + \nu_2 = 1$. The mixture equation derived by Looyenga (1965) describes the electrical behavior of firn well. In general, Looyenga’s equation gives the complex permittivity of a mixture, in terms of two component permittivities ϵ_1 and ϵ_2 :

$$\epsilon^{\frac{1}{3}} = \nu_2(\epsilon_2^{\frac{1}{3}} - \epsilon_1^{\frac{1}{3}}) + \epsilon_1^{\frac{1}{3}}. \tag{8}$$

Equation (8) is symmetrical in ν_2 and ν_1 , i.e. the result is the same regardless of whether the mixture is assumed to be ice in air or air in ice. For the d.c. limit, Equation (8) reduces to

$$\sigma^{\frac{1}{3}} = \nu_2(\sigma_2^{\frac{1}{3}} - \sigma_1^{\frac{1}{3}}) + \sigma_1^{\frac{1}{3}} \tag{9}$$

where σ is the d.c. conductivity. For firn, $\sigma = \sigma_{\text{firn}}$, $\sigma_1 = \sigma_{\text{ice}}$, $\sigma_2 = \sigma_{\text{air}} = 0$ and $\nu_2 = 1 - \nu_1 = 1 - \gamma_{\text{firn}}/\gamma_{\text{ice}}$, where γ_{firn} and γ_{ice} are firn and ice densities, respectively.

The resulting mixture equation for the resistivity of firn is

$$\rho_{\text{firn}} = \rho_{\text{ice}} \left(\frac{\gamma_{\text{firn}}}{\gamma_{\text{ice}}}\right)^{-3}. \tag{10}$$

Bentley (1977) and Shabtaie and Bentley (1979, in press b) found that Equation (10) closely described resistivities measured at several stations on the Ross Ice Shelf and at Dome C.

BED RESISTIVITY

Another parameter needed for calculating ρ_a is the resistivity in the bed. The resistivity of wet, porous sediments, ρ_r , is generally approximated by Archie’s law (Archie, 1942):

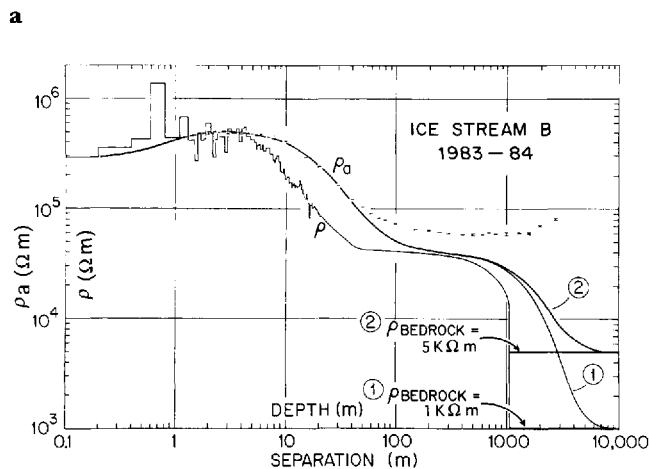
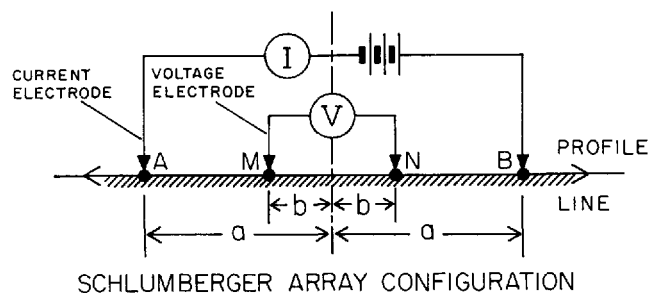
$$\rho_r = \rho_w a \phi^{-m} \tag{11}$$

where ρ_w is the pore-water resistivity, $a \sim 1$ and $m \sim 1.5$ for most sedimentary rocks, and ϕ is the porosity. The resistivity we need is the mean resistivity of the upper few hundred meters of the subglacial sediments. Seismic work (Rooney and others, 1991) indicates that 600 m of low-velocity ($\sim 2 \text{ km s}^{-1}$) sediment underlies the ice at UpB, so no deeper layers need be considered. The wave velocity implies a porosity of about 30% (Blankenship and others, 1987). Using a typical ground-water resistivity of $10 \Omega \text{ m}$ (Shabtaie and others, 1987) leads to a resistivity of only $60 \Omega \text{ m}$. Measurement between two probes 20 m apart in the bed at UpB, however, suggested a resistivity greater than $100 \Omega \text{ m}$ (personal communication from H. Engelhardt, 1990). To be conservative, we take a high resistivity value for modeling purposes ($1000 \Omega \text{ m}$, except where otherwise noted)—adoption of a lower value would only strengthen our conclusion that the basal ice has high resistivity.

APPARENT RESISTIVITY MODELS

From the data presented in Figures 2 and 3, and Equations (7) and (10), we can calculate what $\rho(z)$ would be if it depended on γ and T alone (Fig. 4). In our numerical model, the layer thicknesses range from 0.1 m at the top of the ice sheet to 10 m at the bottom. For model 1, $\rho_{\text{bedrock}} = 1 \text{ k}\Omega \text{ m}$, and $E = 0.25 \text{ eV}$. The corresponding model for ρ_a , the apparent resistivity calculated by the numerical techniques discussed by Shabtaie and Bentley (in press a), is also shown in Figure 4. The ρ_a model fits the data for $a < 40 \text{ m}$ but falls much lower than the data at greater separations. This is still true even when resistivity as high as $5 \text{ k}\Omega \text{ m}$ (pure subglacial meltwater (Gow, 1968)) in the bed is assumed (Fig. 4, model 2). It is clear that the cause of the high resistivity anomaly is not a high value of the bed resistivity, so we will use a value of $1 \text{ k}\Omega \text{ m}$ for the remainder of this analysis.

Several models that incorporate high resistivity anomalies (in addition to the temperature and density effects) were tested; a few examples are shown in Figures 5 and 6. The fact that modeled and measured ρ_a differ for $a > 40 \text{ m}$ indicates that the high resistivity



b Fig. 4. a. Schlumberger four-electrode array configuration; a and b are the current and potential electrode separations (i.e. distance from the center), respectively. The apparent resistivity ρ_a for this configuration is given by

$$\rho_a = \frac{\pi a^2}{2b} \left(1 - \frac{b^2}{a^2} \right) \frac{\Delta V}{I}$$

where ΔV and I are potential difference and current, respectively. b. Plots of apparent resistivity, ρ_a , versus electrode separation and modeled resistivity, ρ , versus depth. The heights of the data-point symbols represent the standard deviations of the measurements. The two models show the effects of temperature and density only, with different assumed resistivities in the bed. The vertical and horizontal scales are the same for both ρ_a and ρ .

must start at a shallow depth. It is clear from model 3 (Fig. 5) that a gradual increase starting at a shallow depth will not work. What is required is an increase in $\rho(z)$ by a factor of about $1\frac{1}{2}$, starting at about 20 m (Fig. 5, models 4 and 5).

An increase by several orders of magnitude in $\rho(z)$ deep in the ice is needed to produce the high ρ_a values at large separations. Models that include such an increase below 800 m fit the data fairly well (Fig. 5, model 5; Fig. 6, model 6) but a still better fit is provided by a model in which $\rho(z)$ first decreases to a lesser value between 650 and 800 m (Fig. 6, models 7 and 8). Model 8 fits the data particularly well. Comparison of models 7 and 8 shows the sensitivity of the models changes in the position of the peak and trough in $\rho(z)$ between 650 and 800 m depth.

The shape of the best-fitting model (Fig. 6, model 8) is similar to the one calculated for other stations in Antarctica. Resistivity depth curves from several sites (RI, near the mouth of Ice Stream E on the Ross Ice

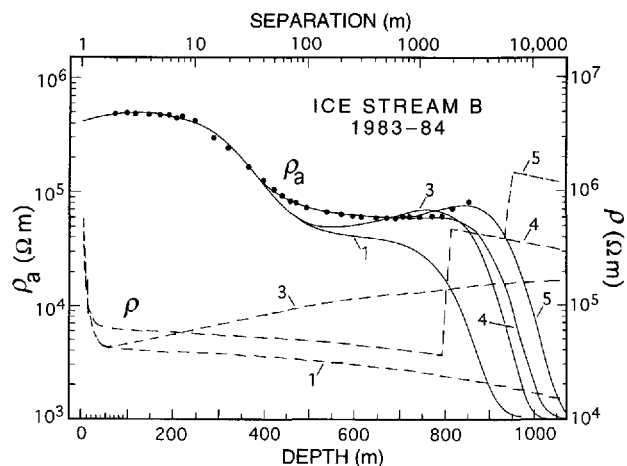


Fig. 5. Plots of modeled resistivity versus depth (dashed lines) and corresponding apparent resistivity versus electrode separation (solid lines) for models 3, 4 and 5. Model 1 is repeated for reference. Dots are the apparent-resistivity data points; the diameter of the dots is equal to the largest standard deviation (the point at $a = 1000$ m in Figure 4). Note the different scales for ρ_a and ρ .

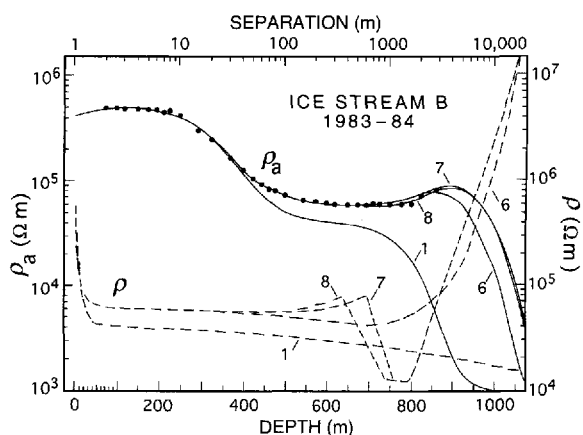


Fig. 6. Plots of modeled resistivity versus depth (dashed lines) and corresponding apparent resistivity versus electrode separation (solid lines) for models 6, 7 and 8. Model 1 is repeated for reference. Dots are the apparent-resistivity data points; the diameter of the dots is equal to the largest standard deviation (the point at $a = 1000$ m in Figure 4). Note the different scales for ρ_a and ρ .

Shelf; J9, also on the ice shelf (Shabtaie and Bentley, 1984, in press a)) are shown along with our model for Ice Stream B in Figure 7. In general, the profiles show similar features; exceptions are the sharp increase in $\rho(z)$ at shallow depth on the UpB profile and the absence at J9 of the order-of-magnitude increase in the basal ice.

The analyses for station J9 (Shabtaie and Bentley, 1984) and Dome C (Shabtaie and Bentley, in press a) show that variation of $\rho(z)$ with depth does not correlate with salts, sulfates, acids or any other impurities found in the chemical analysis of ice cores from those sites. Instead, changes in $\rho(z)$ are highly correlated with crystal-size variation. For example, at Dome C both resistivity and crystal-size increase between 90 and 400 m, and below

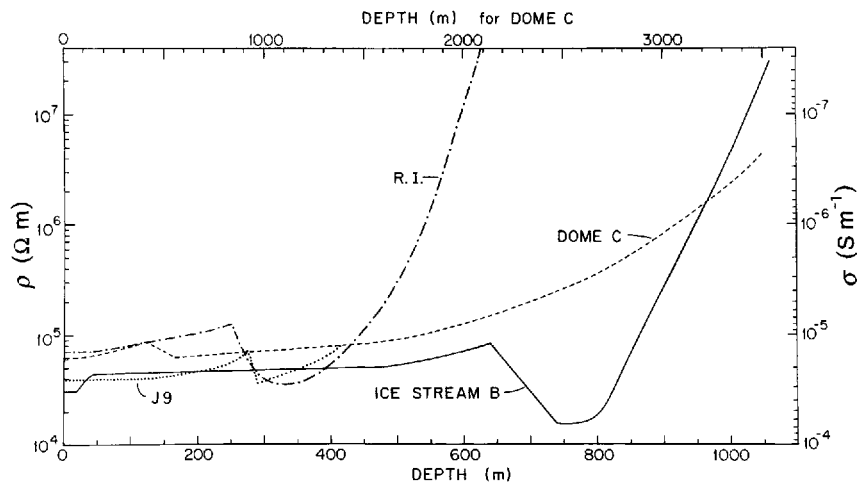


Fig. 7. Best-fitting models of resistivity (or conductivity) versus depth, corrected to solid-ice density and a temperature of -20°C , at four Antarctic locations.

540 m; in between, in the Holocene Wisconsin transition zone, both decrease markedly. This correlation is also found elsewhere (Shabtaie and Bentley, 1984). Furthermore, the shape of $\rho(z)$ at UpB correlates well with crystal-size variation at Byrd Station (Gow and Williamson, 1976). By analogy, therefore, we conclude that the Wisconsin-Holocene transition zone at UpB is at a depth between 650 and 800 m.

ELECTRICAL CONDUCTIVITY BEHAVIOR OF POLAR ICE

It has long been known that resistivities in polar ice sheets are two or three orders of magnitude lower than in temperate and laboratory ice (see Glen and Paren, 1975; Bentley, 1977; Shabtaie and Bentley, 1979, 1984). The resistivities we have obtained from measurements at 13 locations on the East and West Antarctic inland ice and on the Ross Ice Shelf (Shabtaie and Bentley, 1979; work in progress by S. Shabtaie) and measurements elsewhere in the Arctic and Antarctic (Meyer and R othlisberger, 1962;  ostrem, 1967; Glen and Paren, 1975) all show this for the upper layers of the ice sheet. In deep polar ice, however, in both East and West Antarctica, the range of the resistivity values is close to that of temperate ice, although the temperatures are well below freezing.

Fitzgerald and Paren (1975) found that the electrical behavior of polar ice samples changed after melting and refreezing to that of temperate or pure ice. Glen and others (1977) proposed that temperate glaciers are purified by the flushing out of interstitial impurities by percolating meltwater. But phenomena that require melting cannot explain high resistivities in deep polar ice that has always been frozen. Furthermore, field measurements both in the Arctic (V ogtli, 1967) and in the Antarctic (Reynolds, 1982), in zones where there is surface melting and refreezing, yielded resistivities that are typical of polar ice rather than temperate ice.

It is possible that soluble impurities are incorporated substitutionally into the ice lattice and that the concentration of these impurities affects the defect

population, thus changing the resistivity of the ice. Laboratory measurements on monocrystalline ice doped with HF, HCl and NaCl (Camplin and others, 1978; Gross and others, 1978) have shown resistivities comparable with those in polar ice. However, the required impurity concentrations are much higher than the levels found in polar-ice samples: 8–50 μM of HF (Camplin and others, 1978) and 3–40 μM of HCl or NaCl (Gross and others, 1978) compared with impurity levels in polar ice sheets that range from less than 1–3 μM (Delmas and others, 1980; Petit and others, 1981; Herron, 1982; Legrand and Delmas, 1988). Furthermore, there is no correlation between measured near-surface resistivities (corrected to solid-ice density and 20°C) and salt concentration in the form of Na^+ (Boutron and Lorius, 1977; Herron and Langway, 1979) for our Antarctic sites (Fig. 8). Resistivity modeling based on measured variations of salt impurities with depth at J9 (Shabtaie and Bentley, 1984) and Dome C (Shabtaie and Bentley, in press a) also failed to match the measured values. Although measurements in a few deep holes in both the

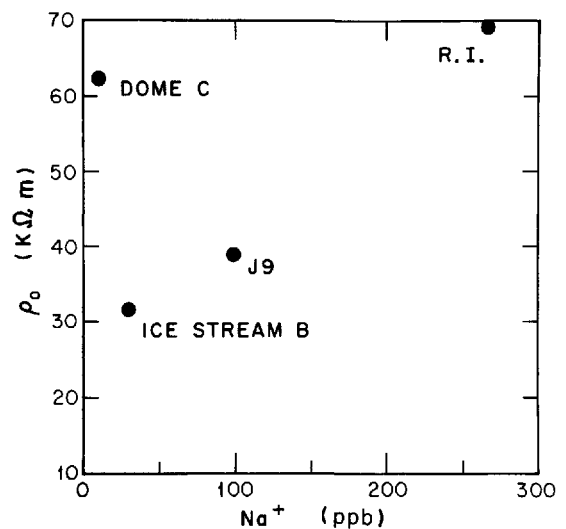


Fig. 8. Near-surface resistivity, corrected to solid-ice density and a temperature of -20°C , versus salt concentration at four Antarctic stations.

Arctic and the Antarctic confirm the existence of impurities throughout the ice column, the impurity concentrations in temperate glaciers, which are found relatively close to industrial centers and/or deserts, are up to an order of magnitude larger than in polar ice (Wagenbach and others, 1988). Clearly, the cause of the very low conductivities in temperate and some polar ice cannot be related to impurities alone; some other factor or factors must come into play.

NEW MODEL

At several locations, we have observed a strong correlation between conductivities and crystal size (Shabtaie and Bentley, 1984, in press a). An explanation for this is found in a theoretical model (unpublished work of S. Shabtaie) in which the impurities surround the ice grains like shells. The impurity coatings must already exist at very shallow depth to explain electrical conduction in near-surface low-density snow (Shabtaie and Bentley, 1990, in press b).

Near the surface, the grains are about 1 mm^2 in size and are ellipsoidal in shape. However, both melting and deformation at high stresses cause extensive recrystallization and crystal growth in the ice. In temperate glaciers and deep in polar ice sheets, crystals can reach over 100 cm^2 in size and are no longer convex in shape; they twist, fold and surround other crystals and become "interlocking" (Rigsby, 1968; Gow, 1970). In our model, this transformation and recrystallization effectively breaks the continuity of the impurity shells surrounding the ice crystals—it is the connectivity of the system that controls the conduction. Breaking the conducting paths does not mean removing the impurities from the system, only disconnecting them. Moderate concentrations of impurities in temperate ice and deep polar ice are accommodated by locating them in isolated domains at grain boundaries.

We believe this new model can explain conduction phenomena throughout an ice sheet. Quantitative details have been worked out for snow and firn by Shabtaie and Bentley (in press a) and for solid ice in a manuscript soon to be submitted.

CONCLUSIONS

A detailed resistivity profile was completed on Ice Stream B at separations from 2 m to 3 km. Resistivities were modeled as a function of depth using measured densities and temperatures, and functional relations that are now well known. Fits were good only in the uppermost part of the ice sheet; below that, model fitting required a 50% increase in resistivity at a depth of only about 20 m and an increase by at least two orders of magnitude in the bottom 200 m or so. A five-fold decrease in resistivity between 650 and 850 m was also suggested; by analogy, we associate this with the Holocene–Wisconsin transition zone. We propose that the principal features of the resistivity–depth profile can be explained by the degree of connectedness between impurities lying on two-grain boundaries. In particular, we believe that the high resistivity at large

depth arises because the extensive recrystallization and grain growth at those depths destroy the connectivity of the two-grain-boundary conduction paths.

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