

NPC, A NEW COMBINED POSITION CATALOGUE OF STARS IN THE NORTHERN POLE REGION

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Abstract. The main task of this work was to construct a new catalogue of positions and proper motions of stars in the north polar cap, and using the results obtained to study the systematic features of the source catalogues as well as the reference system in the polar region. To achieve the goal, thirteen source catalogues, both meridian and photographic, covering an epoch span near one and a half century, were collected. Most of them were not involved in the process of both the PPM and the ACRS constructions. The new combined catalogue of positions and proper motions, named North Polar Catalogue (NPC), lists 4272 stars on the FK5 system, J2000.0. The catalogue was constructed using the method of infinitely overlapping circles. It represents the second use of this procedure to construct a star catalogue. External rms accuracy of positions for the epoch of 2000 is equal to $0''.15$ and $0''.25$ for the centennial proper motions, respectively. The internal rms errors of positions at the mean catalogue epoch, near 1940, are equal to $0''.18$ and $0''.15$ for right ascension and declination correspondingly, while external rms position errors for that epoch are of the order of $0''.05$. Comparisons were made between the NPC and three modern catalogues, the FK5, PPM, and ACRS. The systematic differences between the new catalogue and PPM one for the epoch of 2000 are valuable, especially in the case of errors depending on right ascension. The authors consider as an essential part of the systematic differences that, which is caused by some shortcomings of the PPM in the polar zone. For the mean epoch of the new catalogue positions the systematic differences are smaller.

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This paper gives a result of the first part of investigation of near-polar regions, both the northern and southern. The principal aims of it are as follows: (1) to collect available catalogues of near-polar stars, including both the rarely utilized catalogues, and the modern ones; (2) to investigate in unified approach each catalogue using the comparison with the modern observations, as well as combined and fundamental catalogues; (3) to estimate their random and systematic accuracy and to choose the suitable catalogues as to our interest; (4) to use the conventional rectangular coordinate system to avoid specific problems which arise due to the equatorial system singularity; to use the appropriate catalogue intercomparison method which should be effectively working in the polar region; (5) to compile a combined catalogue of positions and proper motions of near-polar stars above a declination of $+80^\circ$ using a wide observational base; (6) to make the resulting system on the FK5, J2000.0; and (7) to use a catalogue constructed to study the systematic features of modern catalogues in the zone under consideration.

Thirteen source catalogues, eight meridian and five astrographic, have been collected (see Table 1). Most of them have not been utilized when constructing as the PPM catalogue by Roeser and Bastian (1991), as the ACRS one by Corbin and Urban (1991).

As a standard, the FK5 system has been chosen. We used three FK5 representatives: FK5 Basic by Fricke *et al.* (1988), FK5 Extension by Fricke *et al.* (1991), and the improved positions and proper motions of 995 stars by Schwan *et al.* (1993). Two stars from the FK5 Extension were omitted, i.e. BD83:365 and 85:222, because of their obviously mistaken positions. This subset of stars compiled in such a way we refer to FK5*. To improve the tie between the source catalogues and the fundamental system, the FK5* stars were used down to 75° declination. Beginning from early stages of our work we dealt with the J2000.0 equinox. The source catalogues have been transformed to this equinox according to Smith *et al.* (1989). If needed, the computations were carried out considering each coordinate separately when the observation epoch of α and δ differed. To estimate the quality of the source catalogues, fields were built of the systematic differences by averaging the individual differences between the PPM catalogue and contributed ones. It might be noted, that the more remote a catalogue observation epoch is, the larger the differences. We can explain that by the features of the PPM catalogue proper motions.

While dealing with heterogeneous observational data we really are faced with the problem of how to combine the catalogues, both meridian and astrographic, having so different epochs of observations and produced in several different ways. Lack of clear ways to resolve the problem within the framework of available data leads us to use the magnitude and color system of the PPM catalogue, as representative of the FK5 system with the same

TABLE 1. Source catalogues.

No	Catalogue	Number of stars	Mean epoch	Equinox	System	Dsc	Rms
1	Carrington (Carrington, 1857)	3735	1855	1855.0	Car- rington	0.82	0.70 0.60
2	Fabritius (Tel'nyuk-Adamchuk and Molotaj, 1992)	550	1880	B1950.0	FK4	0.42	0.30 0.30
3	Greenwich meridian (Greenwich, 1909)	1501	1901	1900.0	Green- wich	0.39	0.30 0.30
4	Greenwich astrographic (Astrographic catalogue, 1914)	4502	1901	1900.0	Green- wich	0.28	0.20 0.20
5	Berlin (Freundlich, 1916)	1484	1912	1910.0	Pola- rissima	0.56	0.30 0.40
6	Kharkiv-1, Decl. only (Kuz'menko <i>et al.</i> , 1982)	1294	1912	B1950.0	FK4	0.36	— 0.30
7	Pulkovo (Belyavsky, 1947)	4155	1930	B1950.0	?	0.31	0.15 0.15
8	AGK2 (Dieckvoss <i>et al.</i> , 1975)	4317	1930	B1950.0	FK4	0.12	0.25 0.30
9	Yale (Barney and van Woerkom, 1954)	1031	1951	B1950.0	USNO	0.08	0.10 0.10
10	AGK3 (Dieckvoss <i>et al.</i> , 1975)	4317	1960	B1950.0	FK4	0.12	0.15 0.15
11	Kharkiv-3, R.A. only (Derkach, 1994)	452	1964	B1950.0	FK4	0.35	0.20 —
12	Kharkiv-2, R.A. only (Vantzan, 1993)	550	1984	B1950.0	FK4	0.22	0.15 —
13	CAMC (Carlsberg, 1989, 1991, 1992, 1993, 1994)	1462	1990	J2000.0	FK5	0.18	Ori- ginal

Remark to Table 1: Dsc – random rms differences between positions in the PPM catalogue and source catalogue for the mean epoch of observations; Rms – right ascension (above) and declination (below) rms accuracy in arcsec adopted for catalogue positions while determining the positions and proper motions of the combined catalogue.

star density. Removing step by step the systematic components “PPM minus Source Catalogue” depending upon α , δ , magnitude, and spectral class from individual differences we treated the two last components as corrections to star positions in the individual catalogues to bring them onto the PPM magnitude and color system which can be treated as close to the FK5 system. Thus, the system of resulted combined catalogue with respect to

magnitude and color is borrowed from the PPM catalogue. For the sake of the final bringing to the FK5 system and homogenization of individual catalogues we used the filter determined by the method of infinitely overlapping circles (Taff *et al.*, 1990). As a result the Combined Catalogue of Positions and Proper Motions of 4272 Near-Polar Stars North of 80° Declination was obtained (afterwards referred to as North Polar Catalogue, NPC). A quick statistical summary of the NPC catalogue's properties follows: total number of star positions used – 28,728; total number of stars – 4,272; average number of catalogue places per star – 6.6; weighted epoch for α and δ – 1942 and 1937, respectively; the total internal random rms errors: in α at the mean epoch of α – 0''08, in δ at the mean epoch of δ – 0''08, that for the epoch of 2000 – 0''21 and 0''23, respectively, in μ_α and μ_δ – 0.32''/cy and 0.33''/cy, respectively; the total external random rms error as a result of the NPC comparison with the PPM and ACRS catalogues: in α for the epoch of 2000 – 0''18, that for the mean epoch – 0''05, in δ for the epoch of 2000 – 0''15, that for the mean epoch – 0''04, in μ_α – 0.25''/cy, in μ_δ – 0.22''/cy.

To study the external accuracy the NPC was compared with the FK5*, PPM, and ACRS (part 1) catalogues. The position differences are too small especially for the mean epoch (see Table 2). Differences in δ are essentially larger as compared to that in α . The variations of the differences in the proper motions illustrate a strong similarity with those of the positions. Thus, we conclude that position differences are caused by differences in the proper motion system of catalogues. Table 2 illustrates also the significant shift of PPM positions as compared with the ACRS and NPC.

TABLE 2. Systematic shift, in arcsec, in the positions and centennial proper motions between the PPM, ACRS, FK5* , and NPC catalogues in the near-polar zone at the epoch of 2000 and at the mean epoch of NPC (near 1940, in brackets).

Differences	$\Delta\alpha$	$\Delta\delta$	$\Delta\mu_\alpha$	$\Delta\mu_\delta$	N_1	N_2
PPM – NPC	+0''09 (–0''03)	–0''10 (+0''10)	+0.21	–0.32	4174	59
ACRS – NPC	–0.05 (0.00)	+0.01 (–0.01)	–0.08	+0.04	1868	24
FK5* – NPC	+0.01 (0.00)	–0.00 (0.00)	+0.01	–0.01	114	0
PPM – ACRS	+0.10 (–0.02)	–0.13 (+0.10)	+0.20	–0.36	2300	45

Remark to Table 2: N_1 denotes the number of individual differences used to form averaged values; N_2 is the number of stars whose position differences exceed 1.5 arcsec and are ignored, ($N_1 + N_2$) is a number of common stars in these catalogues.

Figure 1 shows the maps of the PPM–NPC, ACRS–NPC, PPM–ACRS, and FK5*–NPC differences, averaged within the circles of one-degree radius. Because of deficit of FK5* in the case of the FK5*–NPC maps, we

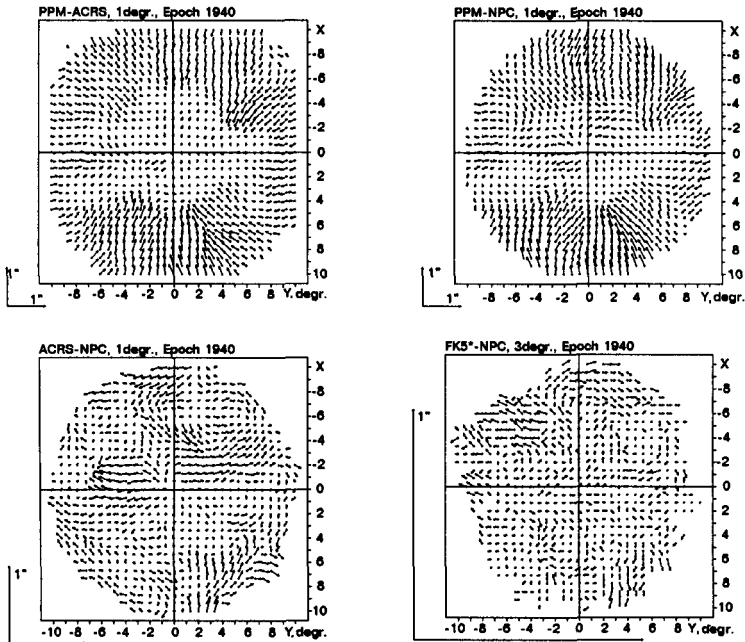


Figure 1. The maps of systematic differences PPM–ACRS, PPM–NPC, ACRS–NPC, and FK5*–NPC in rectangular coordinates (as seen from the North Pole) for the NPC mean epoch. Scales of one arcsec are shown in the bottom left corner of each map.

used three-degree radius. A moving average filter was applied by computing the mean differences in the circles centered in uniformly distributed points. Very similar pictures as well as some central symmetry of maps display the PPM–ACRS and PPM–NPC differences either for the epoch 2000 or for the mean epoch of the NPC. At the epoch of 1940, differences are substantially smaller, symmetry is clearly visible. Moreover, differences are clearly visible between the PPM near-polar stars and stars south of 85° declination. As compared to the PPM–ACRS/NPC the FK5*–NPC differences are quite small, especially at the mean epoch of NPC where they are negligible. Individual differences NPC–PPM, plotted against α , show very interesting peculiarities. The near-two-hour periodicity in the $\Delta\alpha_\alpha$ is clearly seen in $0 - 8^h$ range for 2000 epoch. $\Delta\delta_\alpha$, calculated for the NPC mean epoch, illustrates a clearly visible splitting of the differences into two distinguished subsets above and below 85° declination. Such peculiarities are not seen in the patterns of the ACRS – NPC differences, while they are visible in the PPM–ACRS ones. So, one can link these peculiarities with the PPM catalogue. Moreover, we can deduce that AGK data are responsible of such a periodicity.

The NPC is now available in machine-readable form and can be obtained by sending an e-mail message to vtel@aoku.freenet.kiev.ua.

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