

The orbital evolution of satellite galaxies

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In most of previous theoretical and numerical works on the orbital evolution of satellite galaxies, the classical Chandrasekhar's formula has been adopted. However, a recent direct N-body result ("van den Bosch et al. 1999") demonstrated that the orbital evolution is rather different from what is predicted from the Chandrasekhar's formula. The reason for this discrepancy, however, was not clear.

We first verified the result by "van den Bosch et al.". In our simulations the parent galaxy is a King model with $\Psi_0 = 9$ ($N=32768$), and the satellite is a single particle with 1% mass of parent galaxy (Figure1).

We investigated the reason for the difference between the theoretical work and N-body work. We found that the single largest reason is that in theoretical works the Coulomb logarithm was assumed to be constant throughout the orbit. So, we varied the Coulomb logarithm as,

$$\ln\Lambda = \frac{1}{2} \left[\ln \left\{ 1 + \ln \left(\frac{b_{max} V_s^2}{G(M+m)} \right)^2 \right\} - \ln \left\{ 1 + \ln \left(\frac{\epsilon_s V_s^2}{G(M+m)} \right)^2 \right\} \right]$$

where $b_{max} = r/1.4$ and $b_{min} \simeq \epsilon_s$. This improved model with varying the Coulomb logarithm shows very good agreement with the N-body result.

Next, we varied the softening of the satellite galaxy $\epsilon_s = 0.1, 0.0316, 0.01$ in N-body simulations and semi-analytical simulations. In both results, the orbit of the satellite whose softening is smaller become circular in earlier stage (Figure2).

For large objects such as satellite galaxies, the assumption that the Coulomb logarithm is constant is invalid. The assumption overestimate dynamical friction around the galactic center and circularization and underestimate the timescale of orbital evolution.

References

- Binney J., Tremaine S., 1987, *Galactic Dynamics*. Princeton Univ. Press, Princeton
- Frank C., van den Bosch, G. F. L., G. L., J. S. , 1999, *ApJ*, 515, 50
- Ing-Guey Jiang, J. Binney, 2000, *MNRAS*, 314, 468

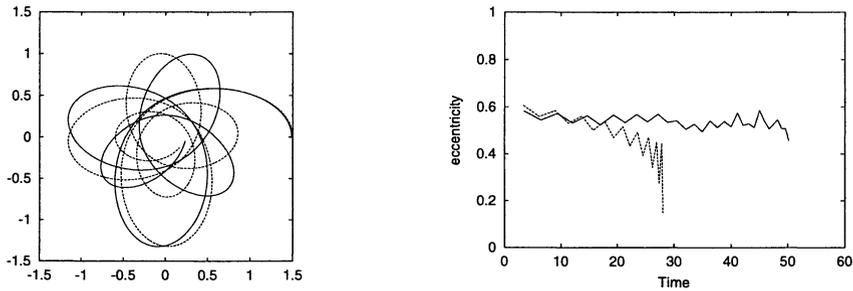


Figure 1. The left panel shows the orbits of satellites, and the right panel shows the orbital eccentricities of the satellites as a function of time. In both panels, full curve shows N-body simulation and dashed curve is that of semi-analytical simulation.

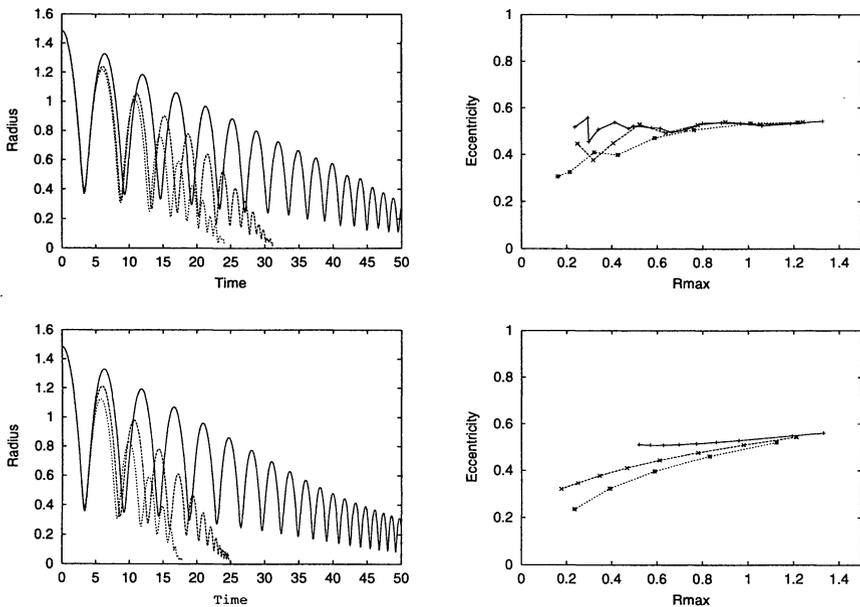


Figure 2. The upper panels are results of N-body simulations, and the lower panels are results of semi-analytical simulations with the modified dynamical friction formula. The left panels show distance of the satellites from the galactic center versus time, and the right panels show orbital eccentricity versus apocentric distance. Full, dashed, and dotted curves represent the results with $\epsilon_s = 0.1, 0.0316, 0.01$, respectively.