SUPERCLUSTERS AND SUPERVOIDS

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1 Introduction

A new catalog of superclusters from the entire sky cluster catalog of Abell et al. (1989) is compiled. In fact the last version of Reference Catalog of ACO Clusters of Galaxies (Kalinkov and Kuneva 1994) is used. The RC contains information of optical, radio and X-ray observations gathered from catalogs, lists and papers.

There are more than 1000 ACO clusters with measured redshift. We have applied multiple regression analysis to obtain redshift estimate for the rest clusters. Estimators in the last version of the RC are almost the same as in Kalinkov *et al.* (1994).

We use measured as well as estimated redshift for luminosity distance to ACO clusters assuming $q_0 = 1/2$ and $H_0 = 100$ km s⁻¹Mpc⁻¹.

2 Catalog of Superclusters of Galaxies

Our searching procedure for systems of galaxies is described by Kalinkov and Kuneva (1986) and Kuneva and Kalinkov (1990).

We use density contrast f=10, 20, 40, 100, 200, and 400. But we define $f=D_{\nu}/D_{100}$, where D_{ν} is the density in a supercluster with multiplicity ν and D_{100} is the density in a sphere with 100 Mpc radius around clusters. Evidently D [cl Mpc⁻³]. Thus f is a local density enhancement. After a supercluster is found f refers to a mean local density enhancement over all supercluster members.

There are 893 superclusters with $\nu \geq 2$ in our catalog altogether. We list all superclusters for $|b| \geq 30^\circ$ and $R \leq 627$ Mpc (z=0.2) independently whether redshift is measured or estimated. A few superclusters with $|b| < 30^\circ$ and R > 627 Mpc are added – in all cases its cluster members have measured redshifts.

For $R \leq 627$ Mpc there are 302 superclusters with $b \geq 40^{\circ}$ and 394 with $b \leq -40^{\circ}$.

There are 4 superclusters with R < 60 Mpc: Nr 173 (A262, 347, 426), Nr 246 (A3163, 3226), Nr 721 (A3656, 3698, 3706, 3742). Hydra-Centaurus supercluster is really remarkable – Nr 542 (A1060, 3526, 3537, 3656, 3574) – for f = 10 - 100 the distance is 40 Mpc, while for f = 200 - 400 is 33 Mpc (A3526, 3565) and for f = 200 is 49 Mpc (A3537, 3574).

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The main Table contains data for superclusters with $\nu \geq 3$, having at least one cluster with measured redshift:

 ν – multiplicity

f – density contrast level

ACO - clusters

RA(1950)Dec and b – galactic latitude

R - luminosity distance ($q_0 = 1/2, H_0 = 100 \text{ km s}^{-1} \text{ Mpc}^{-1}$), Mpc

 $\triangle RA, \triangle Dec, \triangle R$ - corresponding extent, Mpc

 r_{ν} - distance from the supercluster centre to the most distant member (radius; $r_{\nu} = \Delta r/2$ for $\nu = 2$), Mpc

 $r_{\nu+1}$ - distance from the supercluster centre to the nearest cluster outside the configuration, Mpc

 $D = D_{\nu}/D_{100}$ - ratio of the supercluster density to the mean local density, defined within a sphere of 100 Mpc radius

 \mathcal{M} – supercluster mass in units $10^{14}\mathcal{M}_{\odot}$, according to the relation of Bahcall and Cen (1993)

3 Catalog of Voids

Operational definitions of minivoids (without any ACO-cluster in volume $V < 10^6$ Mpc³) and of supervoids, together with searching procedures are given by Kuneva and Kalinkov (1991).

Catalogs of superclusters, minivoids and supervoids could be received from authors.

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