

A consolidated online database of Galactic planetary nebulae

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Abstract. Since the issue of the unifying Strasbourg-ESO Catalogue of Galactic Planetary Nebulae (SECGPN) a large number of new discoveries have been made thanks to improved surveys and discovery techniques. The increasingly heterogeneous published population of Galactic PNe, that we have determined totals <2850 PNe, is becoming more difficult to study on the whole without a centralised repository. We introduce a consolidated and interactive online database with object classifications that reflect the latest multi-wavelength data and the most recent results. The extensible database, hosted by the Centre de Données astronomique de Strasbourg (CDS), will contain a wealth of observed data for large, well-defined samples of PNe including coordinates, multi-wavelength images, spectroscopy, line intensities, radial velocities and central star information. It is anticipated that the database will be publicly released early 2012.

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1. How many Galactic PNe are there?

The database was meticulously constructed by visually inspecting images of purported PNe (sourced from major catalogues, SIMBAD and other literature sources) in the SuperCOSMOS Halpha Survey (SHS, Parker *et al.* 2005), the INT Photometric Halpha Survey (IPHAS, Drew *et al.* 2005) and the SuperCOSMOS Sky Surveys (SSS, Hambly *et al.* 2001). During this process both PNe and obvious non-PNe (e.g. Galaxies, photographic flaws, HII regions, duplicates, etc.) were added to the database. The non-PNe were flagged as different object types to keep them separate and to keep track of objects already considered for inclusion in the database. This laborious process allowed us to make the most accurate estimate to date of the total number of Galactic PNe, N_{gal} , instead of the usual approach of adding total numbers found in new catalogues. We find $N_{\text{gal}} < 2850$, though this upper limit may be reduced as more non-PNe are identified after we further check and refine our database before its first release. Our value is based on currently published catalogues and will also be revised upwards once future PN discoveries are published. Figure 1 depicts the current Galactic distribution of PNe.

2. Database interface

The database will be accessible through a web-based interface that allows users to search for PNe by object name (IAU PN G and common name) and J2000 coordinates (single or list format). Static releases will also be made intermittently to Vizier for virtual observatory compatibility. Results from the web-interface may be browsed in a gallery format with thumbnail images of each PN, as well as a variety of other formats. Groups of PNe may be selected based on their coordinates, object size, name, catalogue and

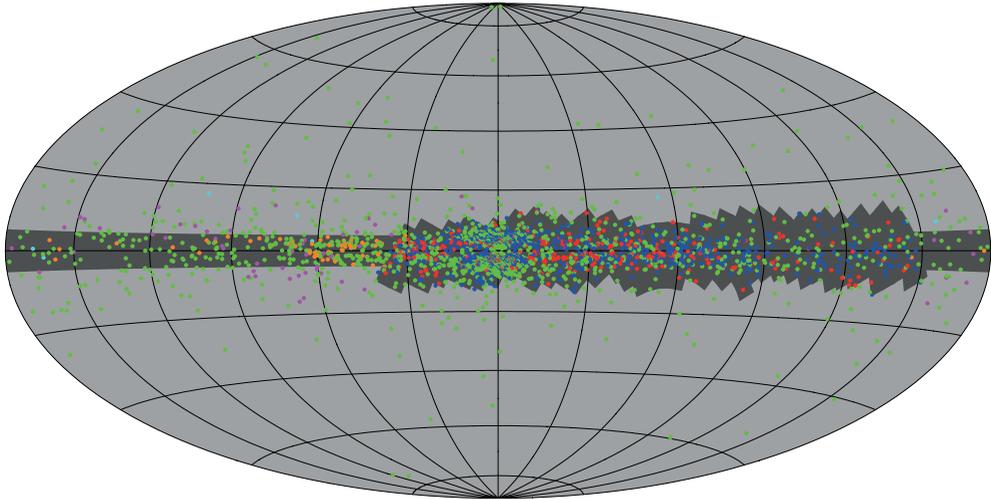


Figure 1. Current database contents in a Galactic Hammer projection. Major published catalogues are coloured separately to most PNe (green) and include MASH (blue, Parker *et al.* 2006), MASH-II (red, Miszalski *et al.* 2008), IPHAS (orange, Viironen *et al.* 2009), Deep Sky Hunters (magenta, Jacoby *et al.* 2010) and the first few discoveries from ETHOS (cyan, Miszalski *et al.* in prep). The darker grey bands represent the IPHAS (Drew *et al.* 2005) and SHS (Parker *et al.* 2005) survey footprints.

so on. Links to online resources are also made available including e.g. the ESO archive, VizieR and Staralt.

3. The First Release

The first release of the database is planned for early 2012 and is intended to produce a working database with the cleanest and largest set of entries for published PNe. This will allow for queries to be made, selecting large samples of PNe to be studied in a unified fashion with accurate coordinates and PN G designations. The initial dataset will include at least image cut-outs from the SHS and SSS (in JPEG and FITS format), SECGPN spectra in FITS format (Acker *et al.* 1992), radial velocities (e.g. Durand *et al.* 1998) and central star magnitudes (e.g. Tylenda *et al.* 1991). Additions to the database will mostly include large observational datasets in refereed publications that are of broad interest to the whole sample.

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