

Twelve years of Glitches in the Vela Pulsar

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1. Introduction

The Vela pulsar (PSR B0833-45) has been observed from Hartebeesthoek RAO since 1984, as part of an ongoing monitoring project. During this time, the pulsar has undergone five sudden and relatively large spin-ups (“glitches”). Good observational coverage was obtained for the four most recent of these events.

2. Post-glitch Timing Behaviour

The post-glitch behaviour of $\dot{\nu}(t)$ is modelled by:

$$\begin{aligned} \dot{\nu}(t - T_G) = & \dot{\nu}_0 + \ddot{\nu}_0(t - T_G) + \sum_i \Delta\dot{\nu}_i \exp(-(t - T_G)/\tau_i) \\ & + \Delta\dot{\nu}_{nl} \left[1 - \frac{1}{1 + e^{-(t - T_G)/\tau_{nl}} (e^{t_0/\tau_{nl}} - 1)} \right], \end{aligned} \quad (1)$$

where T_G is the glitch epoch and subscript *nl* refers to a “non-linear” recovery component of the form discussed by Alpar *et al.* (1989). The results of a preliminary fit of this model to all five glitches are given in table 1. The “baseline model” describes the long-term behaviour; no single simple model provides an adequate fit to all three of the longer inter-glitch eras observed (Flanagan 1996). The immediate post-glitch behaviour (the “transients” in table 1) are modelled with simple exponentials; three such components are required for the larger and better-observed glitches of 1988 and 1991. Following Alpar *et al.* (1993), a model incorporating common timescales for all five glitches was also fitted. The data are consistent with a recovery described by at least three exponential decays of $\Delta\dot{\nu}$, over timescales 50 ± 20 d, 5 ± 2 d and 0.49 ± 0.08 d, following each of the most recent five glitches. In addition, the moment of inertia $I_i = \Delta\dot{\Omega}_i \tau_i / \Delta\Omega_c I$ associated with the 5 d component is $(5 \pm 3) \times 10^{-3}$. The data are inconsistent with the 0.49 d recovery being associated with the same moment of inertia (as defined here) at every glitch.

Table 1. Model Parameters.

	1985	1988	1991	1994	1994A
baseline model:					
$\dot{\nu}(T_G)$ $10^{-13} \text{ Hz s}^{-1}$	-155.54(3)	-156.45(3)	-156.62(4)	-156.54(3)	-156.59(3)
$\ddot{\nu}_0$ $10^{-22} \text{ Hz s}^{-2}$		+9.3(7)	+8.3(2)		
$\Delta\dot{\nu}_{nl}/\dot{\nu}$ %		+0.11(2)			
τ_{nl} days		9(6)			
t_0 days		468(6)			
$\Delta\dot{\nu}/\dot{\nu}$ %	+0.87(2)				
τ days	760(40)				
transients:					
$\Delta\dot{\nu}_1/\dot{\nu}$ %		+0.26(4)	+0.35(6)		
τ_1 days		70(20)	49(6)		
$\Delta\dot{\nu}_2/\dot{\nu}$ %	+1.8(3)	+1.2(5)	+2.3(7)		+0.4(2)
τ_2 days	6.8(4)	5(1)	4.9(6)		6(7)
$\Delta\dot{\nu}_3/\dot{\nu}$ %		+16(6)	+40(10)		
τ_3 days		0.53(9)	0.59(4)		

The components listed in column 1 are those of eqn. 1.

T_G is the glitch epoch.

Errors are $2\text{-}\sigma$ in the least significant digit quoted.

References

- Alpar, M. A., Cheng, K. S. & Pines, D. 1989, *ApJ*, 346, 823
 Alpar, M. A., Chau, H. F., Cheng, K. S. & Pines, D. 1993, *ApJ*, 409, 345
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