## II. CURRENT RADIAL-VELOCITY PROGRAMS FOR S'TARS BRIGHTER THAN NINTH MAGNITUDE R. M. Petrie

This report is prepared from information communicated by a number of astronomers and from data published in the Draft Reports (see volume XIA). Limitations of space and time demand such condensed accounts of programs that they can only be mentioned with a minimum of description.
The report is summarized very briefly in the accompanying tabulation. The numbers of stars, spectral types, magnitudes, instrumentation, etc., were included where available. An effort was made to include programs which might be of interest, even indirectly, to the subject of the Joint Discussion; possibly this aspect of the tabulation has been overdone. Excluded from the summary are: programs recently completed which have been published; programs of binary stars and variable stars as such; programs of stars of special interest unless such programs are extensive. This report is restricted to stars brighter than ninth magnitude but some fainter stars are included if they are part of a program containing bright stars. Objective prism programs are not included.

Twelve institutions, at least, are carrying out radial-velocity programs on the brighter stars. There are only a few long-term programs that include large numbers of stars distributed generally over the sky, the emphasis being placed on shorter programs having specific objectives. It is noted that several observatories are making spectroscopic observations of galactic clusters and associations; several programs are directed at improving our knowledge of stellar velocities perpendicular to the galactic plane. The former poverty of radial-velocity data of stars in the southern sky has vanished under the energetic attack of the Radcliffe, Cape, and Mount Stromlo observatories.

It is estimated that about 5000 stars, mostly brighter than ninth magnitude, are included in current radial-velocity programs. Approximately three-quarters of these are in general programs which will add substantially to the material now available for the study of stellar dynamics.

The tabulation "Summary of Radial-Velocity Programs" is given on the following two pages, 408 and 409.

## Summary of Radial-Velocity Programs

Stars Brighter than 9th Magnitude

\left.| Ref. | Institution | Program | Details |
| :--- | :--- | :--- | :--- |
| No. | (I) General Programs |  |  |$\right]$

## Summary of Radial-Velocity Programs

Stars Brighter than 9th Magnitude

| Ref. <br> No. | Telescope Dispersion | Status | Application |
| :---: | :---: | :---: | :---: |
|  |  | (1) General Programs |  |
| 1 | 36-in. refl. $125 \AA / \mathrm{mm}$. | - | Accel. perp. to Gal. Plane |
| 2 | 74 -in. refl. 33 and 66 $\AA / \mathrm{mm}$. | Two-thirds complete | To supply r.v. of faint stars near zones. |
| 3 | $74-\mathrm{in}$. refl. $66 \AA / \mathrm{mm}$. | Observing and measuring complete | To supply r.v. of faint stars near zones. |
| 4 | $73-\mathrm{in}$. refl. $518 / \mathrm{mm}$. | Observing and measuring complete | Gal. rotation, peculiar motions, interstellar calcium |
| 5 | " | 90\% complete | Accel. perp. to Gal. Plane |
| 6 | $36-\mathrm{in}$. refl. $125 \AA / \mathrm{mm}$. |  | Accel. perp. to Gal. Plane |
| 7 | 200- and ioo-in. refl. 13 to $30 \AA / \mathrm{mm}$. | Complete, 1962 | To give reliable space motions |
| 8 | $74-\mathrm{in} .49$ and $84 \AA / \mathrm{mm}$. | Complete, 1962 | Gal. rotation; dynamics of Galaxy |
| 9 |  | Complete, 1961 | Motions of L.P.V. |
| 10 | $36-\mathrm{in}$. refl. $120 \AA / \mathrm{mm}$. | Observing $80 \%$ complete; measuring $20 \%$ complete | To improve r.v. of northern A stars |
| II | $\begin{aligned} & 74 \text {-in. refl. } 21 \text { to } 84 \\ & \AA / \mathrm{mm} \text {. } \end{aligned}$ | 55\% complete | Motions of near stars and highvelocity stars |
| 12 | , | Completed | Accel. perp. to Gal. Plane |
| 13 |  |  |  |
| 14 |  |  |  |
| 15 |  |  |  |

(2) Special Programs
$36-\mathrm{in}$. ref. $125 \AA / \mathrm{mm}$. Partially complete

Motions of faint stars
Improvement of wave-length tables
r.v., spectral class
r.v. and luminosities

Accurate determination of r.v.
(3) Clusters and Associations

I 73 -in. refl. $51 \AA / \mathrm{mm} ; 30 \%$ complete $90 \AA / \mathrm{mm}$.
r.v.; spectral type; absolute mag.; membership; internal motions
r.v.; motions; membership; galactic rotation
r.v.; spectral class; absolute mag.; motions
r.v.; membership; rot. vel. r.v.
r.v.; membership

