

# Gravitational potential and X-ray luminosities of early-type galaxies observed with XMM-Newton and Chandra

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We study the dark matter content in early-type galaxies and investigate whether X-ray luminosities of early-type galaxies are determined by the surrounding gravitational potential. We derived gravitational mass profiles of 22 early-type galaxies observed with XMM-Newton and Chandra. Sixteen galaxies show constant or decreasing radial temperature profiles, and their X-ray luminosities are consistent with kinematical energy input from stellar mass loss. The temperature profiles of the other 6 galaxies increase with radius, and their X-ray luminosities are significantly higher. The integrated mass-to-light ratio of each galaxy is constant at that of stars within  $0.5-1r_e$ , and increases with radius. The scatter of the central mass-to-light ratio of galaxies was less in K-band light. At  $3r_e$ , the integrated mass-to-light ratios of galaxies with flat or decreasing temperature profiles are twice the value at  $0.5r_e$ , where the stellar mass dominates, and at  $6r_e$ , these increase to three times the value at  $0.5r_e$ . This feature should reflect common dark and stellar mass distributions in early-type galaxies: Within  $3r_e$ , the mass of dark matter is similar to the stellar mass, while within  $6r_e$ , the former is larger than the latter by two-fold. In contrast, X-ray luminous galaxies have higher gravitational mass in the outer regions than X-ray faint galaxies. We describe these X-ray luminous galaxies as the central objects of large potential structures; the presence or absence of this potential is the main source of the large scatter in the X-ray luminosity.

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## JD1 - Poster Session

**Abstract.** During the Joint Discussion #1 a significant number of exciting new results were presented in the form of a poster. Below we list the title and authors of the posters.

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- (a) M. D. Suran, *ΛCDM hydrodynamical cosmological simulations*
  - (b) I. A. Lacerna, *Spatial Correlations of Halo Assembly*
  - (c) P. M de Novais, *Merging pairs of galaxies in the SDSS*
  - (d) R. Salinas, *Kinematics of the field elliptical NGC 7507 – A galaxy with little dark matter?*
  - (e) E. Iodice, *Dark Matter Content in the Polar Disk Galaxy NGC4650A*
  - (f) T. Verdugo, *The whole picture of a galaxy group: Combining Strong Lensing, Weak lensing, Dynamics and N-body simulations in SL2SJ02140-0532*
  - (g) R. Gonzalez, *Galaxy properties within DM halos and Large scale structure*
  - (h) J. A. Magana, *Structure Formation with  $\phi^2$  Dark Matter*
  - (i) P. da Cunha Ferreira, *Predicting the length-to-width ratio on gravitational arcs*
  - (j) V. E. Timofeev, *Observation of ionization jerk in the ionization chamber ASK-1*
  - (k) A. D. Ernest, *Gravitational Eigenstates in the Cosmos: The Answer to Dark Matter?*
  - (l) M. A. Dantas, *Current lookback time-redshift bounds on dark energy*
  - (m) G. B. Caminha, *Fraction of arcs in galaxy clusters: redshift evolution and the importance of magnification*
  - (n) A. S. Iribarrem, *Radial statistics of galaxy number counts at high redshifts*
  - (o) R. R. Rosa, *s Characterizing Extreme Event Dynamics in Galaxy-Sized Dark Matter Haloes*
  - (p) V. A. P. Martin, *Log Slit Spectroscopy and broad-band photometry of the peculiar galaxy ESO 287-G40*
  - (q) D. Bettoni, *The Core Fundamental Plane of low redshift radio galaxies*
  - (r) S. Bryan, *Orbits in Dark Matter Haloes*
  - (s) F. E. M Costa, *Current constraints on dark matter-dark energy interaction*
  - (t) L. Marassi, *Mass Functions in an Homogeneous Dark Energy Model*
  - (u) M. D’Onofrio, *Comparing the FP of early-type galaxies in the V and K bands*
  - (v) E. R. Carrasco *Disentangling the monster. Gemini/GMOS observations of a massive galaxy in the core of Abell 3827*

(w) E. S. Rykoff, *The Origin and Evolution of Fossil Groups*

(x) G. Caminha, *Cross section for arc formation in the perturbative approach*