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UNSTABLE CURRENT SYSTEMS AND PLASMA INSTABILITIES IN ASTROPHYSICS

Edited by MUKUL R. KUNDU and GORDON D. HOLMAN





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INTERNATIONAL ASTRONOMICAL UNION UNION ASTRONOMIQUE INTERNATIONALE

UNSTABLE CURRENT SYSTEMS AND PLASMA INSTABILITIES IN ASTROPHYSICS

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SERGEI I. SYROVATSKII 1925 - 1979

DR. S. I. SYROVATSKII

Sergei Ivanovitch Syrovatskii was born on March 2, 1925 in the city of Bereznegovatoye. In 1941 when the second World War broke out, sixteen years old Syrovatskii went to war. He was awarded two orders and many medals for his valiant conduct during the war.

After graduation from the Physics faculty of the Moscow State University in 1951, Syrovatskii started post graduate studies in the P.N. Lebedev Physical Institute, and studied under the direction of Prof. S. Z. Belenky. Since that time he has worked in the Theoretical Department of the Institute. The first works of Syrovatskii were devoted to magnetohydrodynamics, in particular to different types of surface discontinuities and shock waves. "The Theory of Discontinuities in Magnetohydrodynamics" was the subject of his dissertation in 1954. In 1957 his results on this work were published and were among the first fundamental researches on magnetohydrodynamics carried out in the USSR.

In 1959 Syrovatskii published in the Soviet Astronomical Journal an analysis of relativisitic electron distribution in the Galaxy and of the determination of the Galactic background radio emission. This paper led to the beginning of a long collaboration between Syrovatskii and Ginzburg in radioastronomy.

It was a great achievement in cosmic ray physics when a phenomenological diffusion model of cosmic ray propagation was developed by Ginzburg and Syrovatskii. In the 1960's Syrovatskii along with Ginzburg and Sazonov solved many problems associated with magnetobremsstrahlung radiation, and elaborated many of its characteristic features. Using the first measurements of relativistic galactic cosmic ray electron flux and of the Galactic background radio intensity, Ginzburg and Syrovatskii investigated in 1961 the question of origin of relativistic electrons in the Galaxy. They showed that direct acceleration of electrons in the Galactic sources was important for radio emission rather than their origin as secondary products due to nuclear interactions in cosmic rays. In 1963, Ginzburg and Syrovatskii published the famous monograph "The Origin of Cosmic Rays".

In mid 1960's, Syrovatskii started studies of plasma behavior in strong magnetic fields. He showed that in the neighborhood of the zero point of magnetic fields in a plasma, there can develop a current sheet with a width much larger than the thickness. These studies were used for modelling plasma behaviour in the solar corona, where the magnetic field energy density is much higher than the plasma energy density. The processes of formation, stability and disruption of current sheets has been studied for many years theoretically, numerically and under laboratory conditions by Syrovatskii and his group. A very important property of current sheets is their stability. In the typical solar corona their lifetime is several hours. This property enables a very large energy to be accumulated in the solar atmosphere in the form of magnetic energy of a current sheet. Current sheet disruption causes release of the accumulated energy in different forms, which is a solar flare in the Syrovatskii's model.

Over the years, the problems studied by Syrovatskii's group involved many new branches of solar physics. These problems include hydrodynamics of the convective zone of the Sun, solar dynamo, particle acceleration in solar flares and secondary processes on the Sun. Syrovatskii had more than 200 scientific publications to his credit.

Syrovatskii hecame a Professor at the Lehedev Phyisical Institute in 1979. On September 26, 1979 he died of a heart attack.

In 1982, Sergei Syrovatskii (posthumously) and his collaborators were awarded the State Prize of the USSR for a series of papers on "Dynamics of Current Sheets and Solar Activity".

> - Volodya Dogiel P.N. Lebedev Physical Institute Moscow, U.S.S.R.





1. Tanaka 2. Lui 3. Elphic 4. Sato 5. Goossen 6. Trehan 7. Mahajan 8. Lotko 9. Smith 10. Krishan 11. Waldron 12. Underhill 13. Degaonkar 14. Uchida 15. Priest 16. Kundu 17. Stenzel 18. Hardee 19. Eilek 20. Benford 21. Davila 22. Ionson 23. Smith 24. Wu 25. Gilden 26. Kaastra 27. Wentzel 28. Van Hoven 29. Shevgaonkar 30. Ray 31. Steinolfson 32. Birn 33. Cheng 34. Coroniti 35. Nocera 36. Papadopoulos 37. Eichler 38. Vasliunas 39. Karpen 40. Sonnerup 41. Ray 42. Krautter 43. Goodrich 44. Huba 45. Bratenahl 46. Henriksen 47. Mullan 48. Ghosh 49. Dum 50. Chiuderi 51. Sturrock 52. Nelson 53. Aly 54. Wiita 55. Norman 56. Sakai 57. Nishikawa 58. Goldstein 59. Spicer 60. Lokanadham 61. Tsinganos 62. Rust 63. Spangler 64. Berger 65. Gergely 66. Bridle 67. Heyvaerts 68. Schmahl 69. Holman 70. Ferrari 71. Einaudi 72. Peacock 73. Nordlund 74. Migliuolo 75. Bell 76. Drake 77. Hasegawa 78. Vlahos.