

Monitoring UT1 from astro-geodetic techniques at the EOP Center of the IERS

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Abstract. Monitoring the Earth rotation is essential in various domains linked to reference frames firstly with applications in orbit determination, space geodesy or Astronomy. Secondly for geophysical studies where are involved mass motions within the different external fluid layers, atmosphere, hydrosphere, core and mantle of the earth, this on time scales ranging from a few hours to decades. The Earth Orientation Centre of the IERS is continuously monitoring the earth orientation variations from results derived from the various astro-geodetic techniques. It has in particular the task of deriving an optimal combined series of UT1 which is now based mainly on Very Long Baseline Interferometry (VLBI) with some contribution of LOD derived from GPS. We give here a brief summary concerning the contribution of the various techniques to UT1 and in particular how the use of LOD derived from GPS can improve the combination. More details are available in Gambis (2004) and Bizouard and Gambis (2009) and the website <http://hpiers.obspm.fr/eop-pc/>

Keywords. Universal Time UT1, Earth Rotation

1. UT1 determination: monitoring and accuracy

Earth Orientation Parameters describe the orientation of the earth with respect to a non rotating reference frame. One of the parameter, Universal Time UT1 represents the rotation of the Earth around its axis. Until the 1970's, UT1 was exclusively monitored by astrometric techniques based on optical instruments like photozenithal tubes, PZT, meridian refractors and astrolabes. In the 1970's the emergence of Lunar Laser Ranging (LLR) allowed to determine UT0. In 1985 its accuracy was in the range of 0.400 ms. Meanwhile VLBI technique was emerging and determined UT1 with an accuracy at least ten times better than LLR (Table 1).

Table 1. Contribution of astro-geodetic technique to the determination of UT1 and LOD

Technique	Since	EOP	Time resolution	Accuracy
ASTROMETRY	1899	UT1	5 days	1 ms
LLR	1969	UT0	1 day	0.4 ms
SLR	1976	LOD	3 days	200 μ s
VLBI	1981	UT1 Standard	3-4 days	5 μ s
	1981	UT1 Intensive	1 day	15 μ s
	1981	LOD	3-4 days	15 μ s
GPS	1993	LOD	1 day	10 μ s

2. Contribution of LOD derived from GPS estimates in UT1 series, method of combined smoothing

VLBI, as the only technique referring to a non rotating celestial reference frame is the main contributor of UT1. Alternatively, satellites techniques, like GPS which are realizing their celestial frame through the orbit determination do not allow determining an accurate UT1; because of mis-modelling of various perturbations, the orbit is affected by long-term systematic variations. GPS techniques can nevertheless determine daily LOD estimates of which values are slightly biased. LOD(GPS) can be used for UT1 computation when calibrated by UT1 derived from VLBI using the so-called method of “Combined Smoothing” Vondrak *et al.*, 1999; Vondrak, 2000. Figure 1 shows the comparison of various combined UT1 series with the external Final Bulletin A series. It appears that the contribution of GPS LOD either by the direct integration of LOD(GPS) or when applying the Combined Smoothing leads to a small improvement of a few μs in the WRMS compared to the solution which does not incorporate any LOD(GPS) data. It is also striking that the contribution of intensive session is only a few μs .

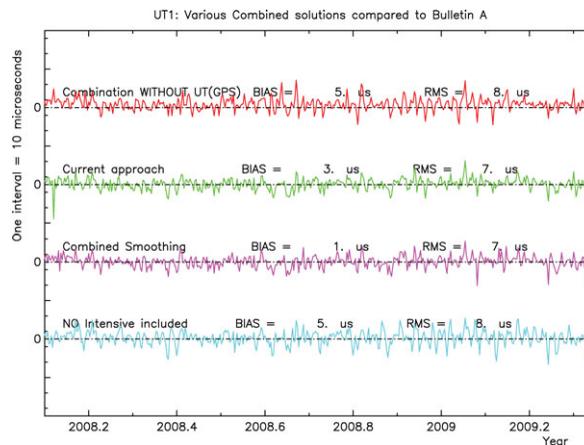


Figure 1. Comparisons of different UT1 combined series using or not UT(GPS). It appears that the contribution of GPS LOD either by the direct integration of LOD(GPS) referred as Current approach, or when applying the Combined Smoothing leads to a small improvement in the WRMS compared to the usual solution. We can remark that the inclusion of Intensive sessions does not significantly improve the final combined UT1 solution.

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