A NEW SAMPLE OF FAINT HALO B STARS

KENNETH J. MITCHELL
General Sciences Corporation
6100 Chevey Chase Dr., Laurel, MD 20707
REX A. SAFFER
Space Telescope Science Institute
3700 San Martin Dr., Baltimore, MD 21218
AND
STEVE B. HOWELL

STEVE B. HOWELL

Planetary Science Institute
2421 E. 6th St., Tucson, AZ 85719

1. Summary

This poster reports the isolation of a new, complete sample of 24 faint halo B stars. These B stars were selected as blue- and ultraviolet-excess objects in the US survey at high Galactic latitudes (Usher and Mitchell 1990), and were given preliminary classifications using low-resolution spectrophotometry. The new sample is complete over 206 square degrees of sky to faint magnitude completeness limits in the range B=16.5 and B=18.3.

Stromgren color indices for the US B stars have been derived from the spectrophotometry through numerically convolved filters (Howell 1986). The colors have been used to help define the red completeness limit of the sample at (b-y)=-0.01, at the B9.5/A0 classification boundary. In addition, surface gravity and temperature estimates useful for separating the hotter B stars from sdB stars have been provided by atmospheric model fitting to the existing spectra (cf. Saffer $et\ al.\ 1994$).

References

Howell, S.B. (1986) Astronomical Journal, 91, 171. Saffer, R.A. et al. (1994) Astrophysical Journal, 432, 351. Usher, P.D. and Mitchell, K.J. (1990) Astrophysical Journal Supplement, 74, 885.