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A detailed investigation of the CrB supercluster of galaxies was initiated. More than 50 new radial velocities were determined. The observations were carried out with the 6-m telescope.

Cluster A2001 may be regarded as a projection of two clusters at different distances (Ciardullo et al., 1983). A2004 is the same case. We show that A2067 is also a result of projection of two clusters. The distance is determined with  $H_0 = 100 \text{ km/s/Mpc}$  and

 $q_0 = 1$ . It may be that the only galaxy with measured radial velocity in A2056 is a foreground galaxy. Eastern of this cluster is A2056E, which is in fact not an A-clusters, but Zw 165-49 (mc, pop. 166, MD).

Our results definitely show that the CrB supercluster consists of two superclusters - a near one (CrB-N) and a distant one (CrB-D).

In Table 1 velocity dispersions for 6 clusters are given. For A2065 (namely CrB cluster), the velocity dispersion is taken from Spinrad (1977), and for A2069 - from Struble and Rood (1986).

Assuming  $R_g = 1$  Mpc for each cluster, the virial masses are in the range  $(0.2 \div 2.5) \cdot 10^{15}$  M0 with  $\overline{M} = 1.0 \cdot 10^{15}$  M0. Then the mass of CrB-N or CrB-D will be about  $7 \cdot 10^{15}$  M0.

A	ර <sub>corr</sub> km s <sup>-1</sup>	A	σ <sub>corr</sub> km s <sup>-1</sup>
2061 2065 2067A 2067B	685 1070 751 532	2069 2079 2092	565 549 272

Table 1

The mean radial velocity of CrB-N is 22178 km/s and of CrB-D it is 35214 km/s. Formally estimating  $\sigma$ (CrB-N)=1736 km/s and  $\sigma$ (CrB-D)= 1266 km/s and formally calculating the virial mass, we have M(CrB-N)  $\approx$  1.7 \cdot 10<sup>17</sup> MO for R<sub>g</sub> = 26 Mpc and M(CrB-D) $\approx$ 7.0 · 10<sup>16</sup> MO for R<sub>g</sub> = 20 Mpc.

## REFERENCES

Ciardullo, R. et al. 1983, ApJ, 273, 24. Spinrad, H. 1977, Publ. Astron. Soc. Pacific, 89, 116. Struble, M. F. and Rood, H. J. 1986 (preprint).

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