

## 1995-1998 Large-Scale Campaigns on $\lambda$ Boo Star 29 Cygni

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**Abstract.** We present the main results of a photometric and spectroscopic study of the  $\lambda$  Boo star 29 Cyg carried out from 1995 – 1998.

29 Cyg was the first  $\lambda$  Boo star in which well-defined pulsations were detected (Gies & Percy 1977). It was selected by the Central Asian Network (CAN) group (Mkrtichian et al. 1998) and collaborators as a key object to investigate the pulsational characteristics of  $\lambda$  Boo stars. We intensively studied 29 Cyg during 1995 (Mkrtichian & Kusakin 1996; Paunzen & Handler 1996), 1996 (Mkrtichian et al. 1998), 1997 and 1998 multisite campaigns; the main results can be summarized as follows:

- We established a well-defined multiperiodicity in 29 Cyg and determined frequencies and amplitudes for nine low-degree  $\ell$  modes with amplitudes greater than 0.9 mmag. The stationary amplitude solution for the 1996–1997 V filter light curves obtained for these nine frequencies using the differential correction code (Andronov 1994) is given in Table 1. We found

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long-term and short-term variability of amplitudes for all of these modes in 29 Cyg – the first time this has been done for a  $\lambda$  Boo star.

- For the first time for  $\lambda$  Boo stars we have detected pulsational multiperiodic radial velocity variations.
- Based on spectroscopic radial velocity and line-profile data we found that both low- and high-degree nonradial modes are excited.
- Based on multicolor  $W, B, V, R$  photometry we determined the amplitude-wavelength dependence for low-degree modes in 29 Cyg.
- The photometric behavior of 29 Cyg is similar to the pulsational characteristics of  $\delta$  Scuti stars.

Table 1. Nine-frequency solution for the 29 Cyg 1996–1997  $V$ -data.

| Designation | Frequency<br>(c/d) | $\pm\sigma_f$<br>(c/d) | Semi-ampl.<br>(mmag) | $\pm\sigma_a$<br>(mmag) |
|-------------|--------------------|------------------------|----------------------|-------------------------|
| $f_1$       | 37.425904          | 0.000005               | 10.72                | 0.06                    |
| $f_2$       | 34.720723          | 0.000012               | 4.80                 | 0.06                    |
| $f_3$       | 29.775771          | 0.000020               | 2.82                 | 0.06                    |
| $f_4$       | 25.188602          | 0.000024               | 2.41                 | 0.06                    |
| $f_5$       | 27.503653          | 0.000035               | 1.63                 | 0.06                    |
| $f_6$       | 28.159113          | 0.000050               | 1.14                 | 0.06                    |
| $f_7$       | 25.459053          | 0.000056               | 1.04                 | 0.06                    |
| $f_8$       | 34.911538          | 0.000058               | 0.95                 | 0.06                    |
| $f_9$       | 32.626003          | 0.000059               | 0.93                 | 0.06                    |

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